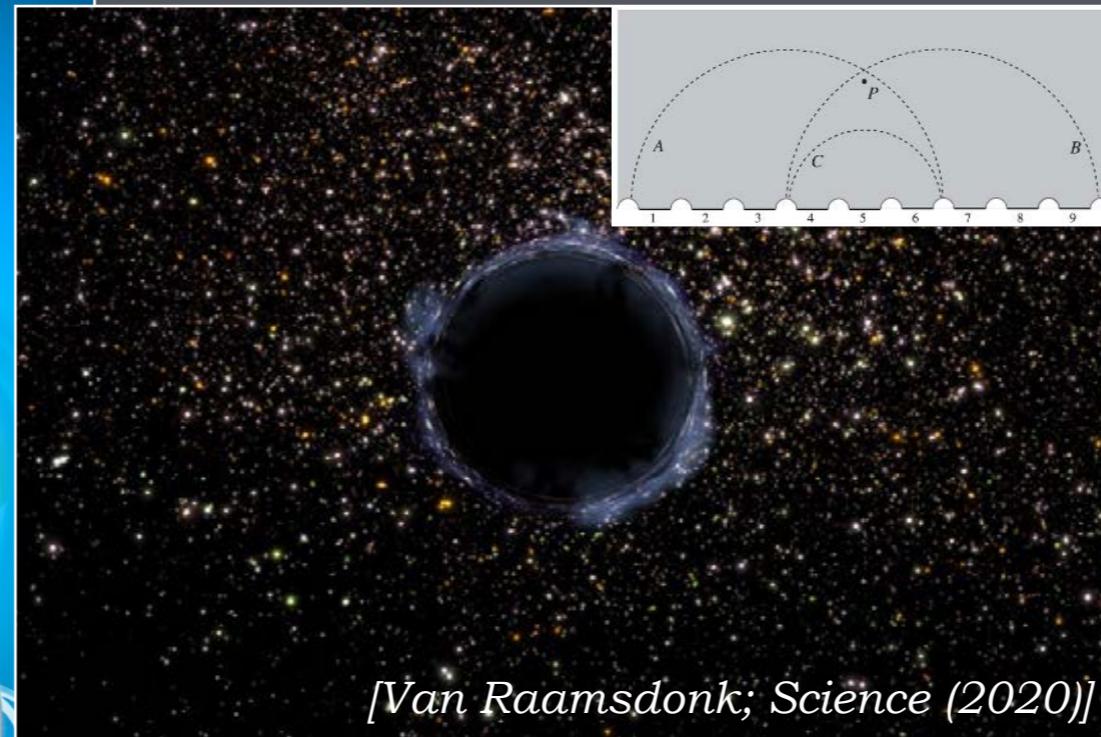
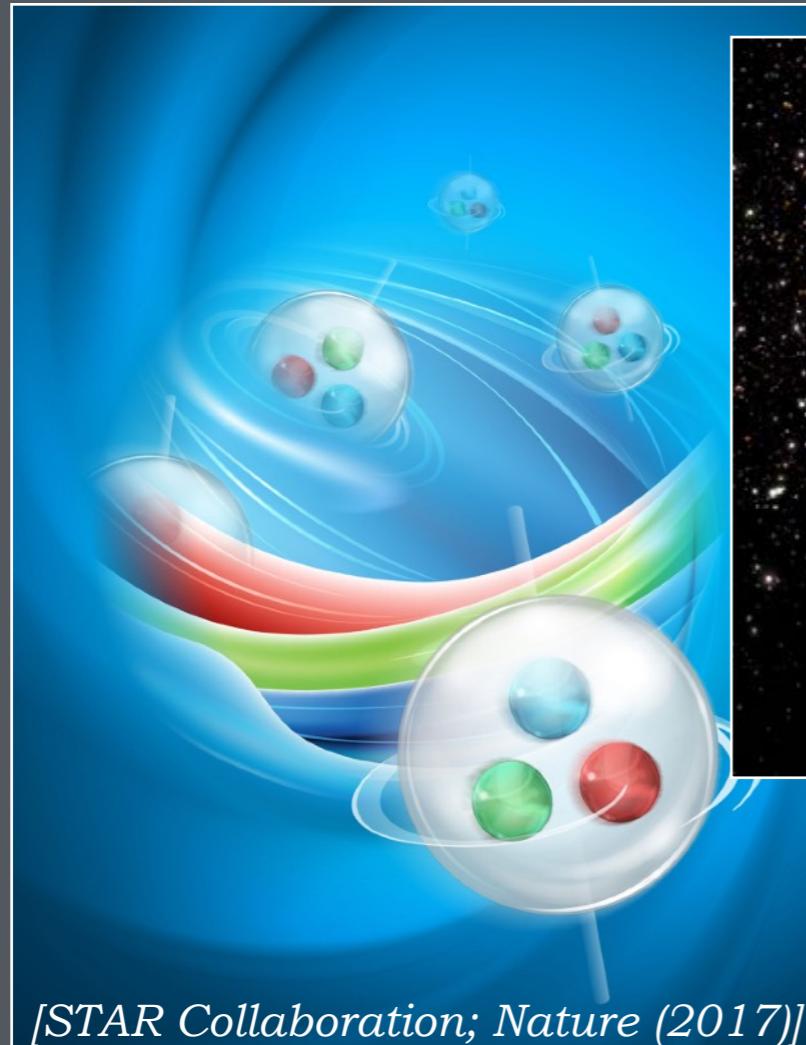


Entanglement, thermalization (and transport) in quark-gluon plasma, in ultra-cold gases, & black holes

Quantum Few- and Many-Body Systems in Universal Regimes, INT, Seattle

October 23rd, 2024



Matthias Kaminski
University of Alabama



Circular rainbow viewed from plane above Seattle



Circular rainbow viewed from plane above Seattle



[<https://deepai.org/>]

Gauge/Gravity Correspondence
=
Holography
=
AdS/CFT

Circular rainbow viewed from plane above Seattle



[<https://deeppai.org/>]

**Universality at large N (number of colors):
Shear viscosity to entropy density bound**

[Kovtun, Son, Starinets; PRL (2005)]

$$\frac{\eta}{s} = \frac{1}{4\pi}$$

(Finite N corrects, finite
coupling respects bound
[Buchel, Myers, Sindha; JHEP (2008)]
[Cremonini, Mod.Phys.Lett.B (2011)])



Gauge/Gravity Correspondence

=

Holography

=

AdS/CFT

Outline

- 1. Statistical versus quantum mechanic dynamics**
- 2. Holography (far from equilibrium)**
- 3. Holographic entanglement entropy (calculation)**
- 4. Quantum gravity experiments**

Thermodynamics: entropy increases

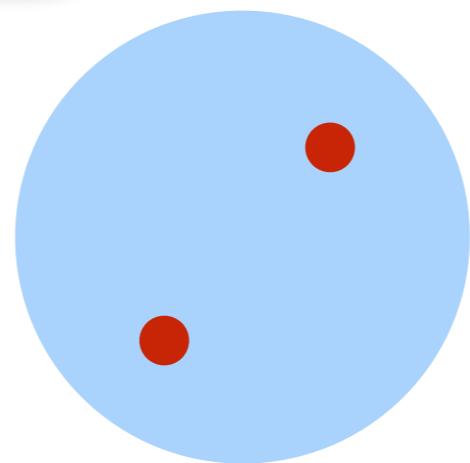
Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.



Entropy = Non-Information

$S \propto \# \text{ of configurations}$



time ↓



Thermodynamics: entropy increases

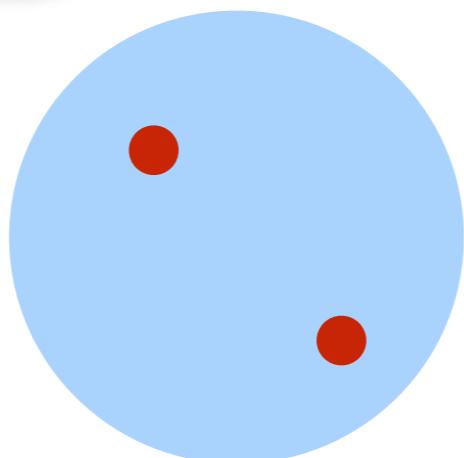
Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.



Entropy = Non-Information

$S \propto \# \text{ of configurations}$



time ↓

Definition via density matrix

$$\rho = \sum_i p_i |n_i\rangle\langle n_i|$$

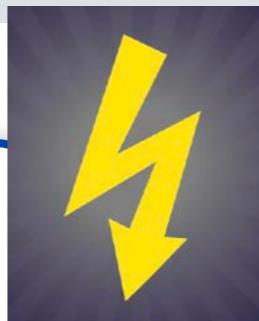
$$S = -\text{tr}(\rho \log \rho)$$



Clash of two fundamental concepts

Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.



Quantum mechanics:

A **pure state** (with zero entropy) remains pure with **zero entropy**.

[Kaufman et al.; Science (2016)]

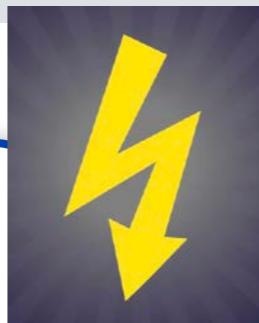
“Quantum thermalization through entanglement in an isolated many-body system”



Clash of two fundamental concepts

Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.



Quantum mechanics:

A **pure state** (with zero entropy) remains pure with **zero entropy**.

[Kaufman et al.; Science (2016)]

“Quantum thermalization through entanglement in an isolated many-body system”



Pure state density matrix

$$\rho = 1 \cdot |\text{eigenstate}\rangle\langle\text{eigenstate}| \text{ with } \text{tr}(\rho^2) = 1.$$

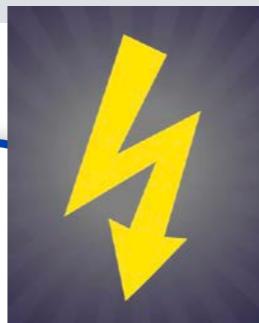
$$\text{Vanishing entropy } S = - \text{tr}(\rho \log \rho) = 0$$

Unitary time evolution $\rho \rightarrow U^{-1}\rho U$ keeps $S=0$.

Clash of two fundamental concepts

Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.

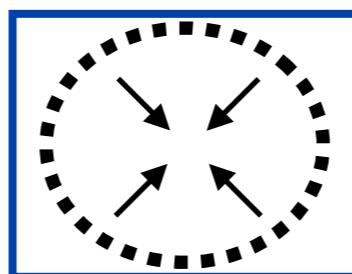


Quantum mechanics:

A **pure state** (with zero entropy) remains pure with **zero entropy**.

Example 1: **Black hole evaporation**

[Almheiri et al.; Rev.Mod.Phys. (2021)]



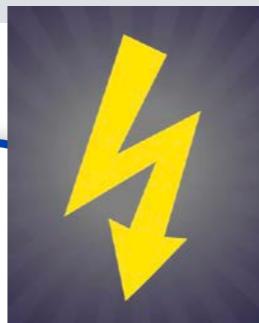
time

Pure state. **Zero entropy.**
(inward falling mass shell)

Clash of two fundamental concepts

Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.

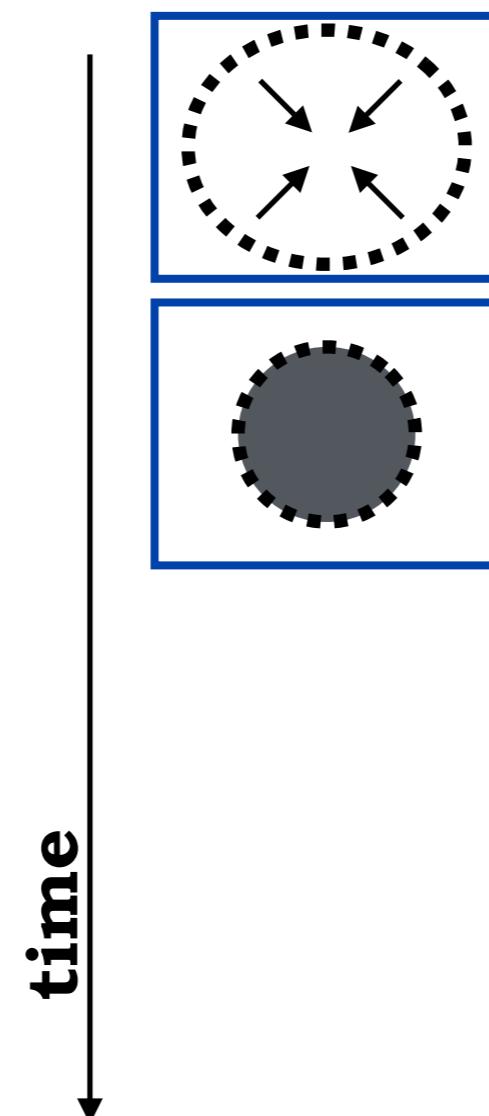


Quantum mechanics:

A **pure state** (with zero entropy) remains pure with **zero entropy**.

Example 1: Black hole evaporation

[Almheiri et al.; Rev.Mod.Phys. (2021)]



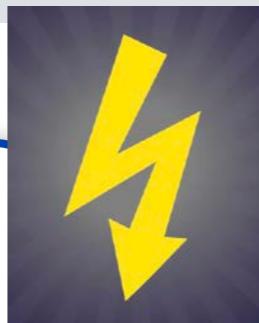
Pure state. **Zero entropy.**
(inward falling mass shell)

Black hole is formed.

Clash of two fundamental concepts

Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.

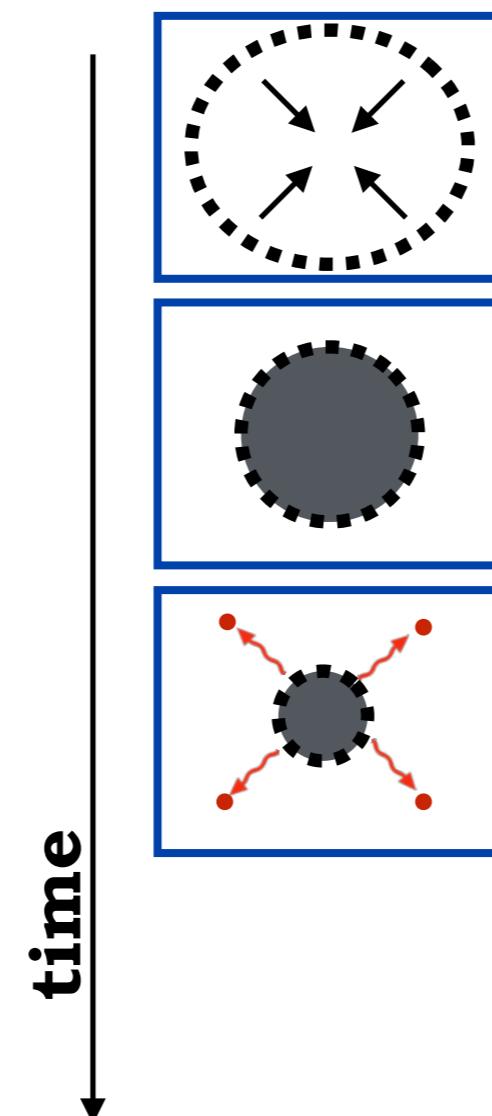


Quantum mechanics:

A **pure state** (with zero entropy) remains pure with **zero entropy**.

Example 1: Black hole evaporation

[Almheiri et al.; Rev.Mod.Phys. (2021)]



Pure state. **Zero entropy.**
(inward falling mass shell)

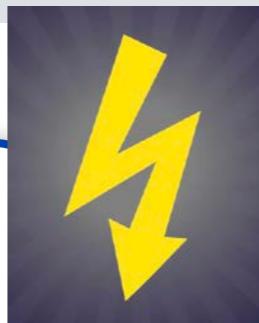
Black hole is formed.

Hawking radiation,
black hole evaporates.

Clash of two fundamental concepts

Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.

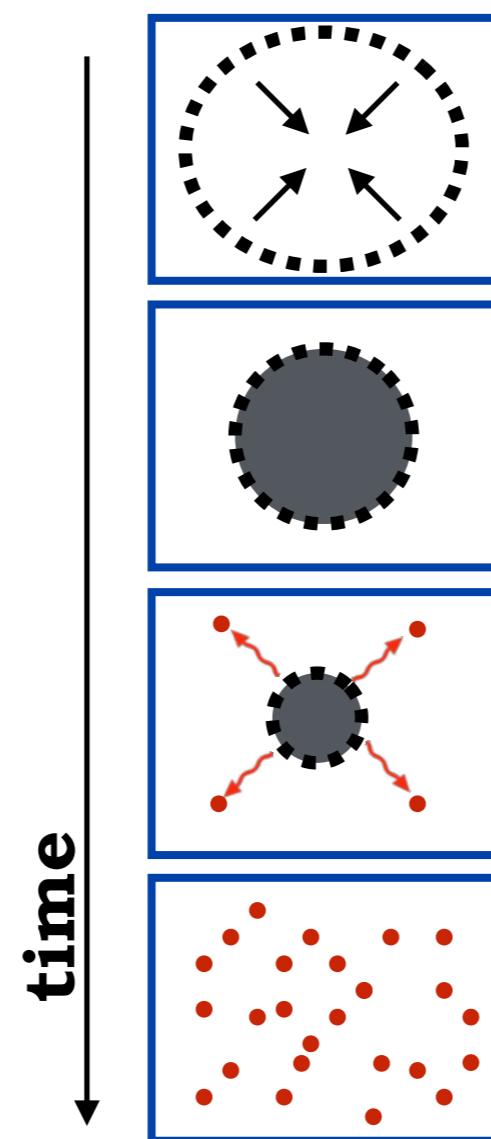


Quantum mechanics:

A **pure state** (with zero entropy) remains pure with **zero entropy**.

Example 1: Black hole evaporation

[Almheiri et al.; Rev.Mod.Phys. (2021)]



Pure state. **Zero entropy.**
(inward falling mass shell)

Black hole is formed.

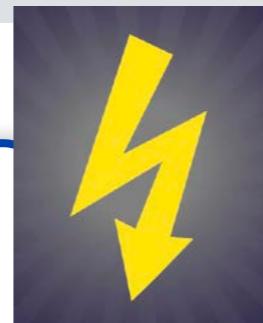
Hawking radiation,
black hole evaporates.

Thermal Hawking radiation.
Nonzero entropy!?

Clash of two fundamental concepts

Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.



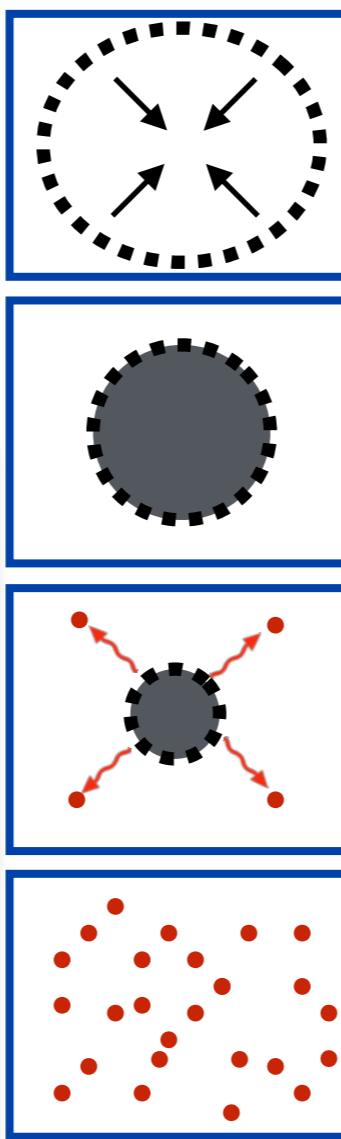
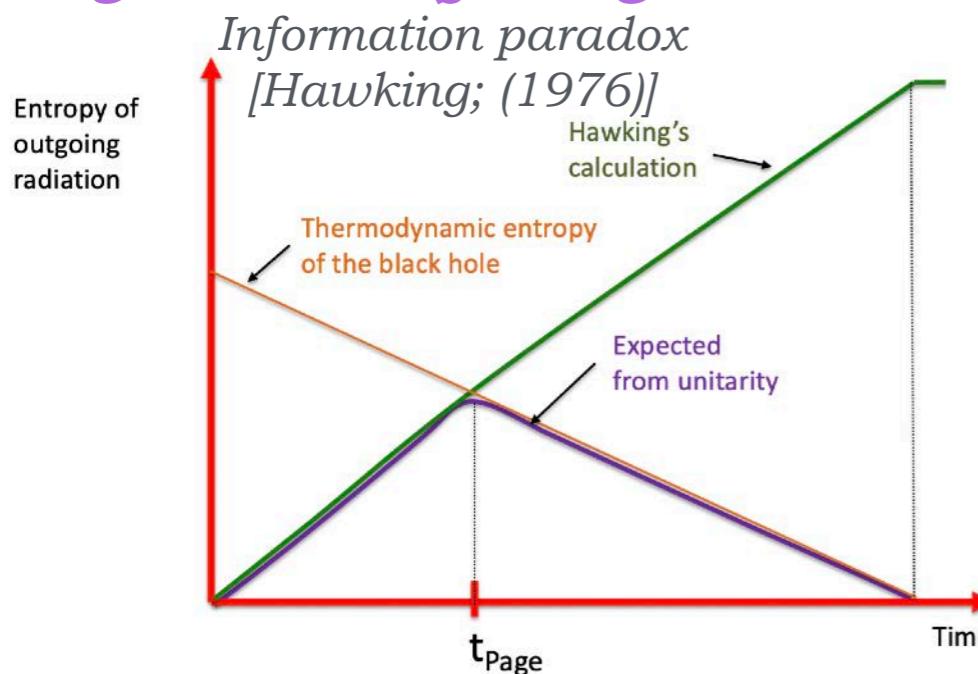
Quantum mechanics:

A **pure state** (with zero entropy) remains pure with **zero entropy**.

Example 1: Black hole evaporation

[Almheiri et al.; Rev.Mod.Phys. (2021)]

Page curve (fine-grained S)



Pure state. **Zero entropy**.
(inward falling mass shell)

Black hole is formed.

Hawking radiation,
black hole evaporates.

~~Thermal Hawking radiation.
Nonzero entropy!?~~

Clash of two fundamental concepts

Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.



Quantum mechanics:

A **pure state** (with zero entropy) remains pure with **zero entropy**.

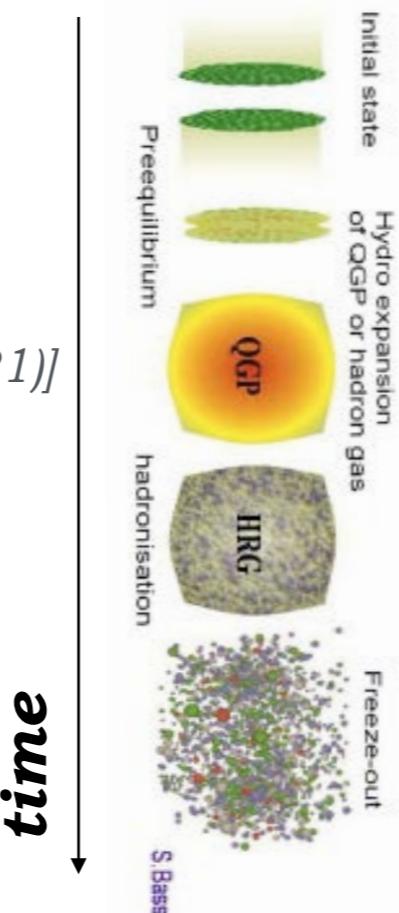
Example 1:
Black hole evaporation

Example 2:
Heavy ion collisions (HIC)

Quantum information approach to HIC
[Kharzeev; Phil.Trans.A.Math.Phys.Eng.Sci. (2021)]

[Zhang et al.; PRD (2021)]

[Florio,Kharzeev; PRD (2021)]



Pure state. **Zero entropy.**
(two colliding ions)

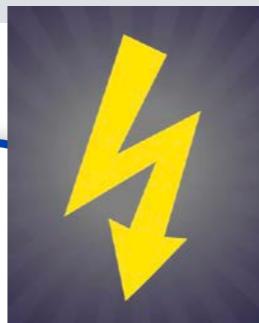
[Müller,Schäfer; (2017)]

Particle distributions in detectors look thermal. **Nonzero entropy!?**

Clash of two fundamental concepts

Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.



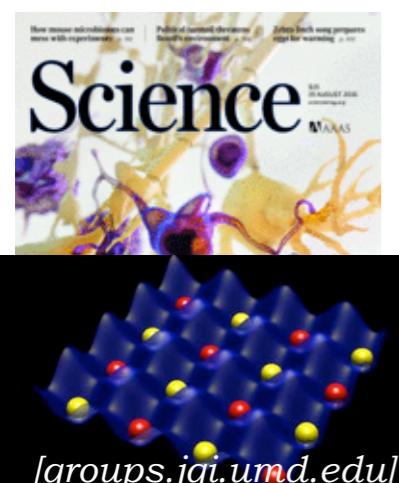
Quantum mechanics:

A **pure state** (with zero entropy) remains pure with **zero entropy**.

Example 1:
Black hole evaporation

Example 2:
Heavy ion collisions

Example 3:
Ultracold atoms



[Kaufman et al.; Science (2016)]
“Quantum thermalization through entanglement in an isolated many-body system”

- Rb atoms in optical lattice
- six-site Bose-Hubbard system
- quench and microscopy

time ↓

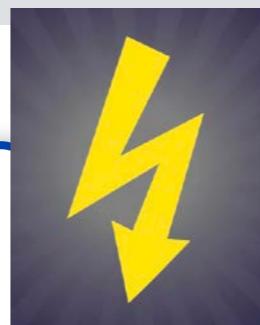
Pure state. **Zero entropy.**
Remains pure!

**BUT subsystems thermalize!
(nonzero entropy)**

Clash of two fundamental concepts

Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.

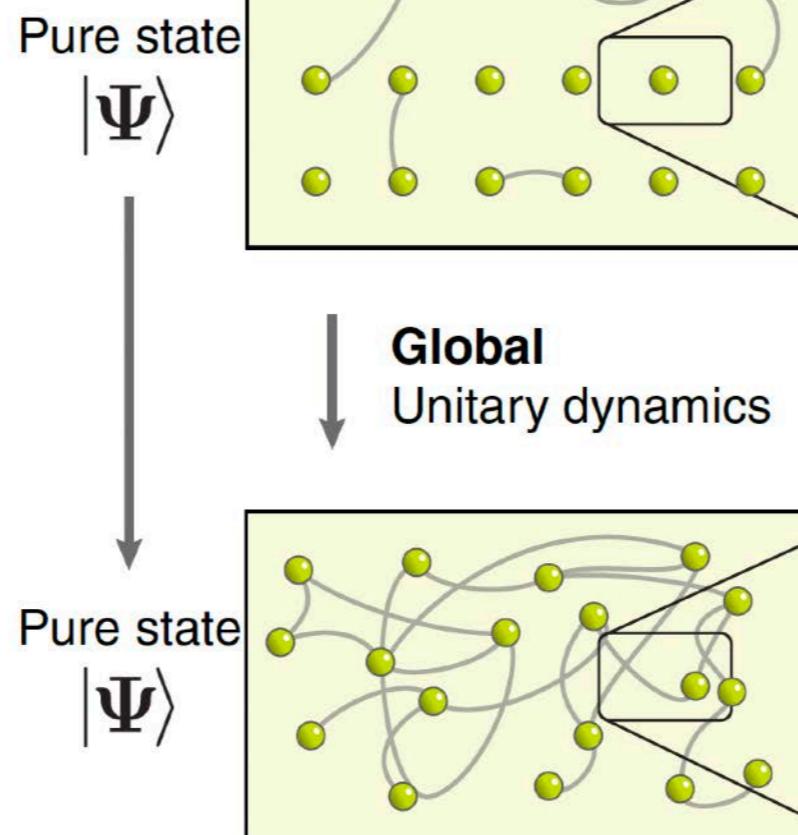


Quantum mechanics:

A **pure state** (with zero entropy) remains pure with **zero entropy**.

General proposed resolution

- Total system remains pure.
 - Subsystems thermalize (nonzero entropy).
 - Eigenstate thermalization hypothesis (ETH)?
[Srednicki; (1993)]
- ... many open questions!



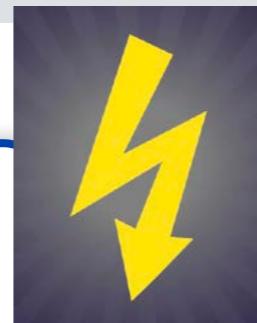
[Kaufman et al.; Science (2016)]

grey links:
entanglement

Clash of two fundamental concepts

Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.



Quantum mechanics:

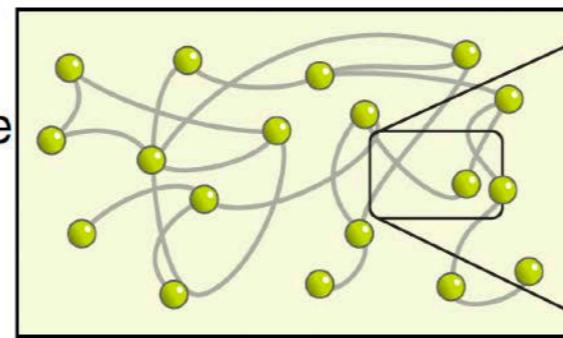
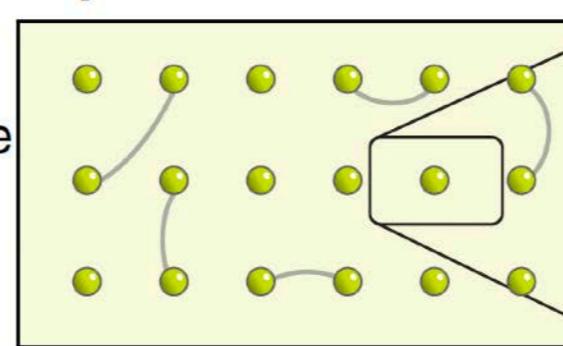
A **pure state** (with zero entropy) remains pure with **zero entropy**.

General proposed resolution

- Total system remains pure.
 - Subsystems thermalize (nonzero entropy).
 - Eigenstate thermalization hypothesis (ETH)?
[Srednicki; (1993)]
- ... many open questions!

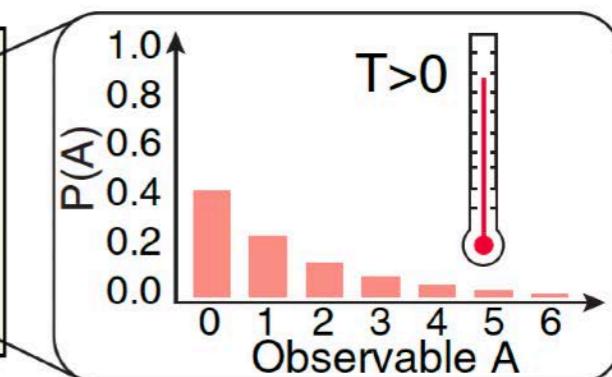
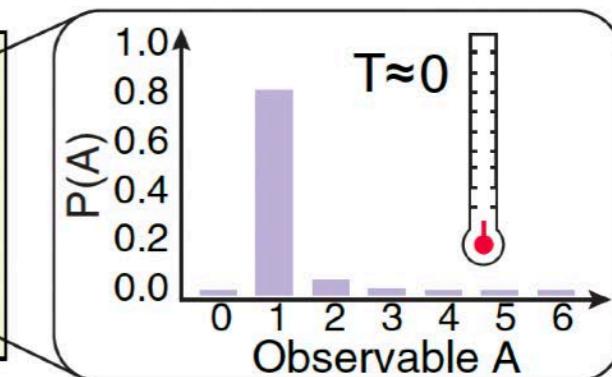
Pure state
 $|\Psi\rangle$

Pure state
 $|\Psi\rangle$



Global
Unitary dynamics

Local
Thermalization



grey links:
entanglement

[Kaufman et al.; Science (2016)]

Clash of two fundamental concepts

Statistical mechanics:

An **isolated system** evolves such that it **maximizes its entropy**.

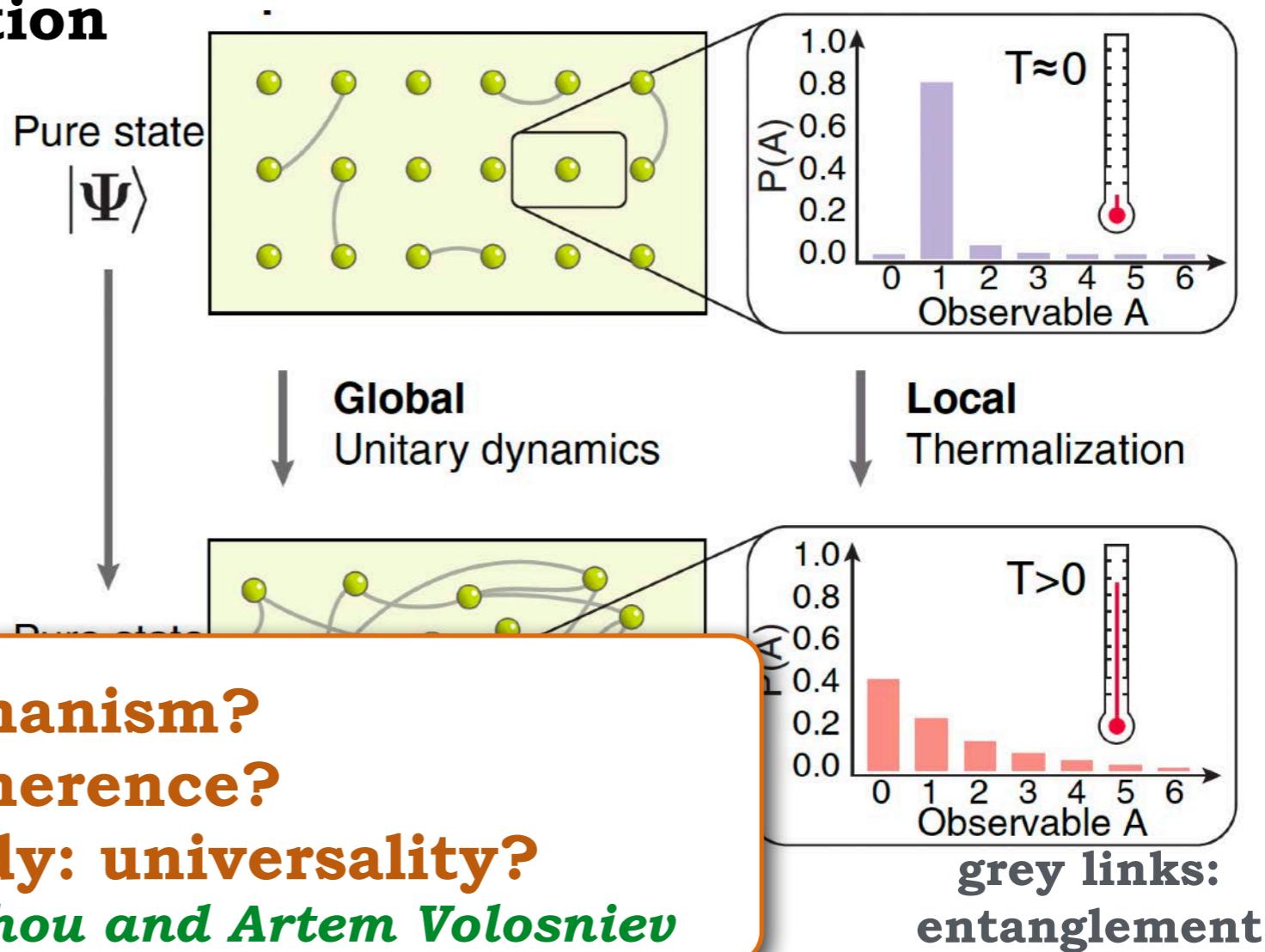


Quantum mechanics:

A **pure state** (with zero entropy) remains pure with **zero entropy**.

General proposed resolution

- Total system remains pure.
- Subsystems thermalize (nonzero entropy).
- Eigenstate thermalization



- **thermalization mechanism?**
- **entanglement/decoherence?**
- **few versus many body: universality?**
- **talks by Qi Zhou and Artem Volosniev**

Physical question

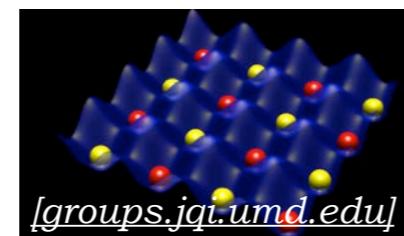
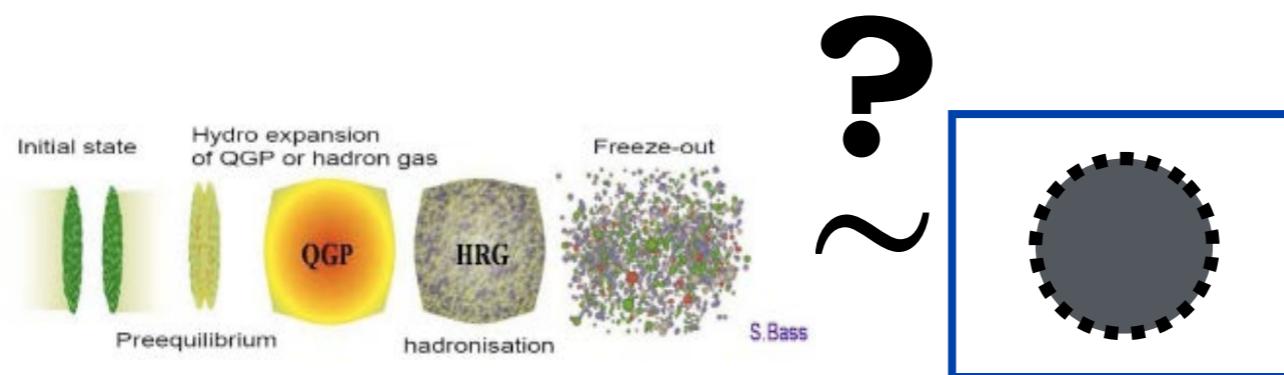
**How do black holes
and other isolated quantum systems
thermalize?**

→ *talk by Aurel Bulgac (slower than ETH-thermalization)*

Outline

1. Statistical versus quantum mechanic dynamics
2. **Holography (far from equilibrium)**
3. Holographic entanglement entropy (calculation)
4. Quantum gravity experiments

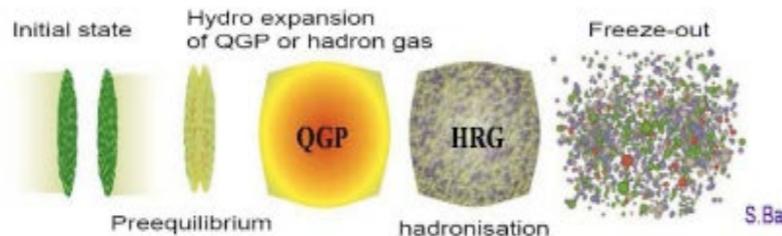
Ion collision/Ultracold atoms ~ Black hole formation?



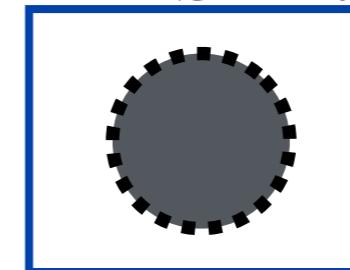
Exact correspondence: Holography (gauge/gravity)



**$N=4$ Super-Yang-Mills
in 3+1 dimensions
with $SU(N)$ and λ
(gauge)**



**Typ II B Superstring Theory
in (4+1)-dimensional
Anti de Sitter space
(gravity)**

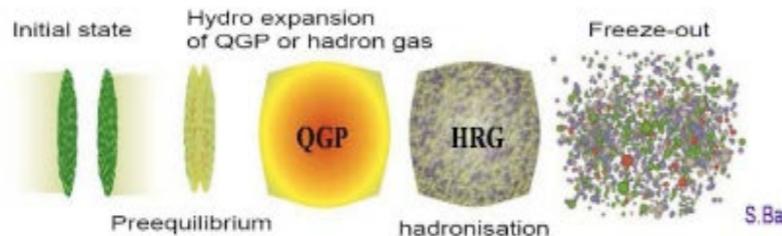


[*t Hooft (1993)*]
[*Susskind; J.Math.Phys. (1995)*]
[*Maldacena; Adv.Theor.Math.Phys. (1997)*]

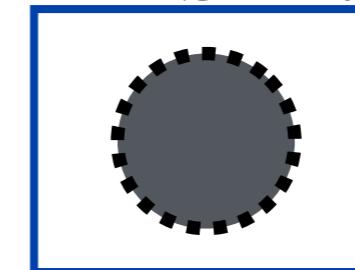
Exact correspondence: Holography (gauge/gravity)



**$N=4$ Super-Yang-Mills
in 3+1 dimensions
with $SU(N)$ and λ
(gauge)**



**Typ II B Superstring Theory
in (4+1)-dimensional
Anti de Sitter space
(gravity)**



[t Hooft (1993)]

[Susskind; J.Math.Phys. (1995)]

[Maldacena; Adv.Theor.Math.Phys. (1997)]

Black hole entropy grows as its **surface area**
(not as its volume).

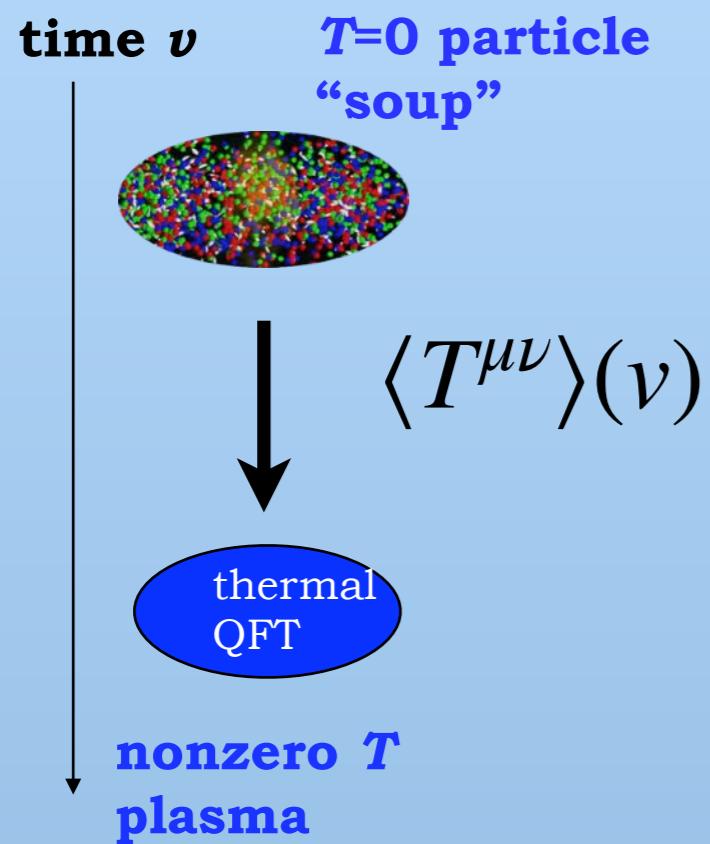
[Bekenstein]

[Hawking]

$$S \propto A_h$$

Far from equilibrium holography

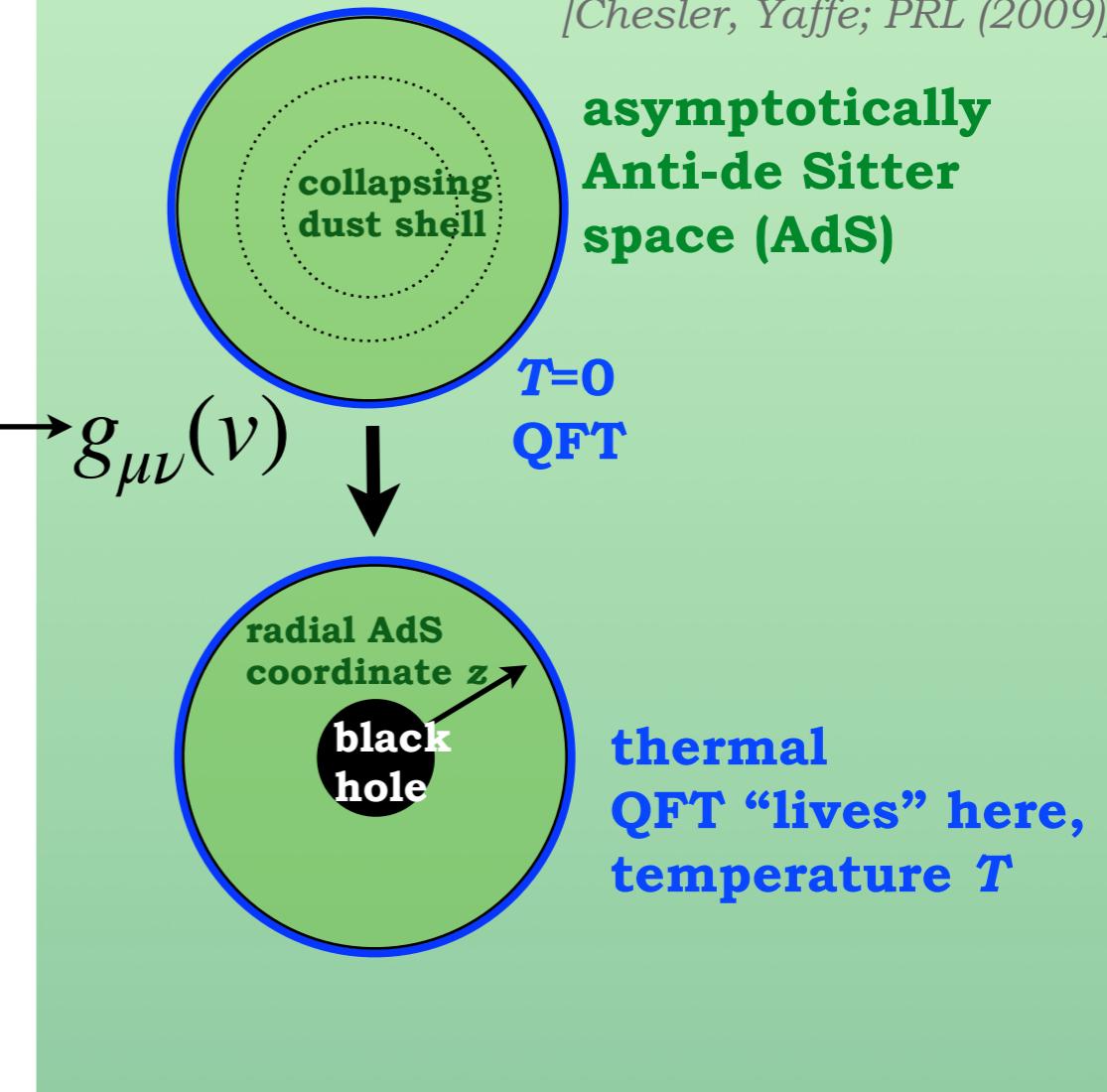
Thermalization in field theory



gauge/gravity correspondence

Horizon formation in gravity

[Janik, Peschanski; PRD (2006)]
[Chesler, Yaffe; PRL (2009)]



→ exact/numerical thermalization results from holography

[Cartwright, Kaminski; JHEP (2019)]

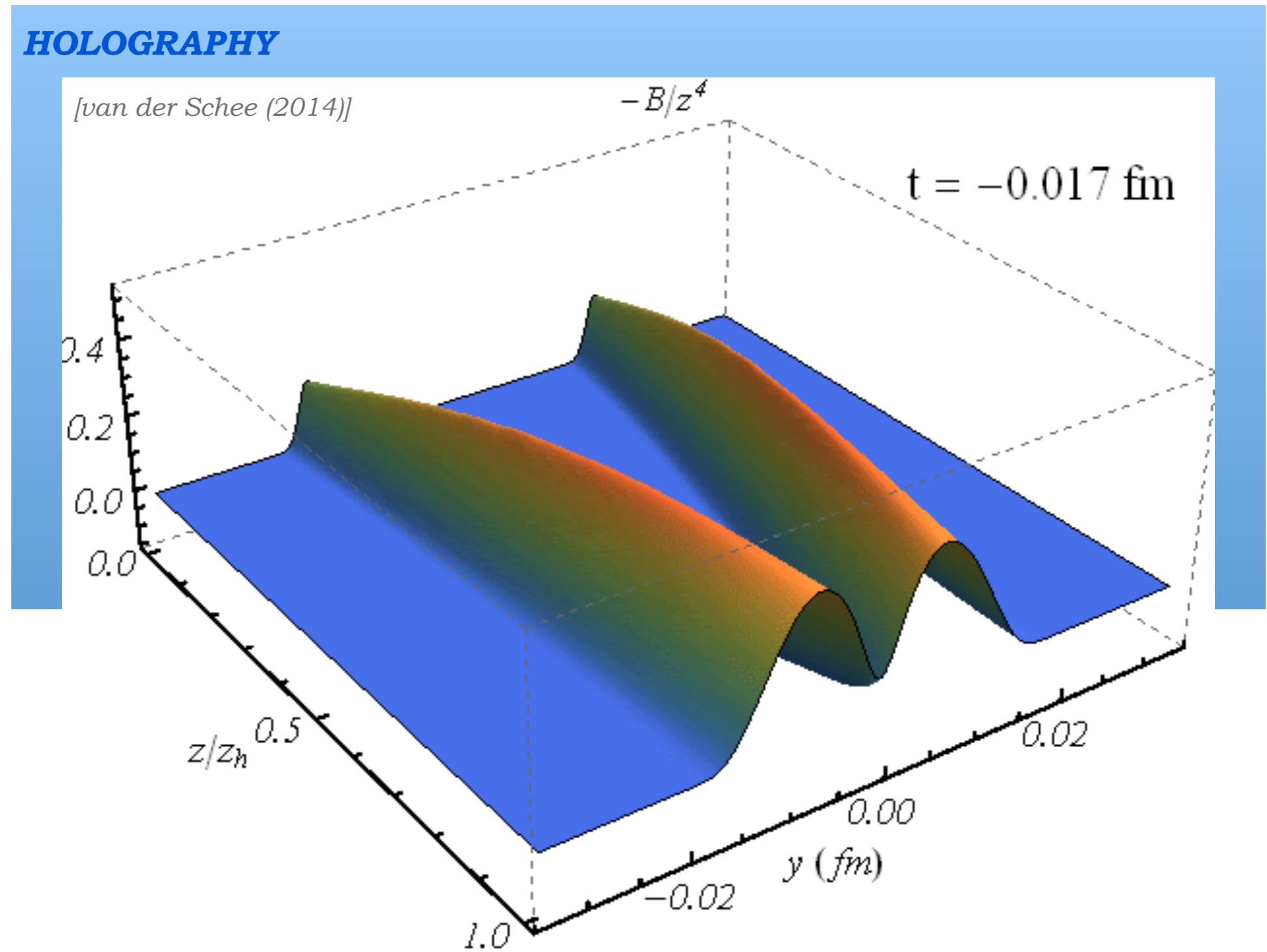
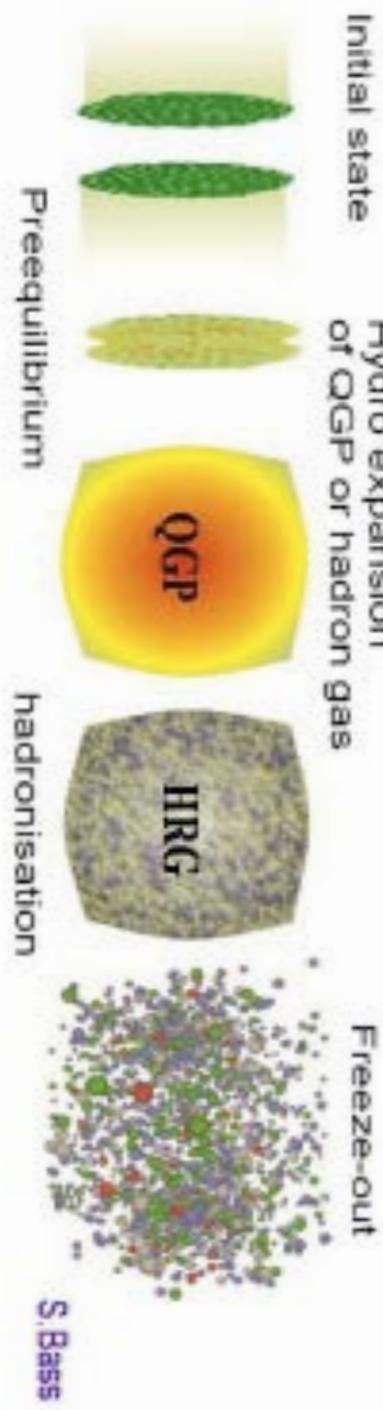
[Cartwright, Kaminski; JHEP (2021)]

[Cartwright, Kaminski, Knipfer (2022)]

[Cartwright, Kaminski, Schenke; PRC (2022)]

Holographic heavy ion collision (numerical, large N)

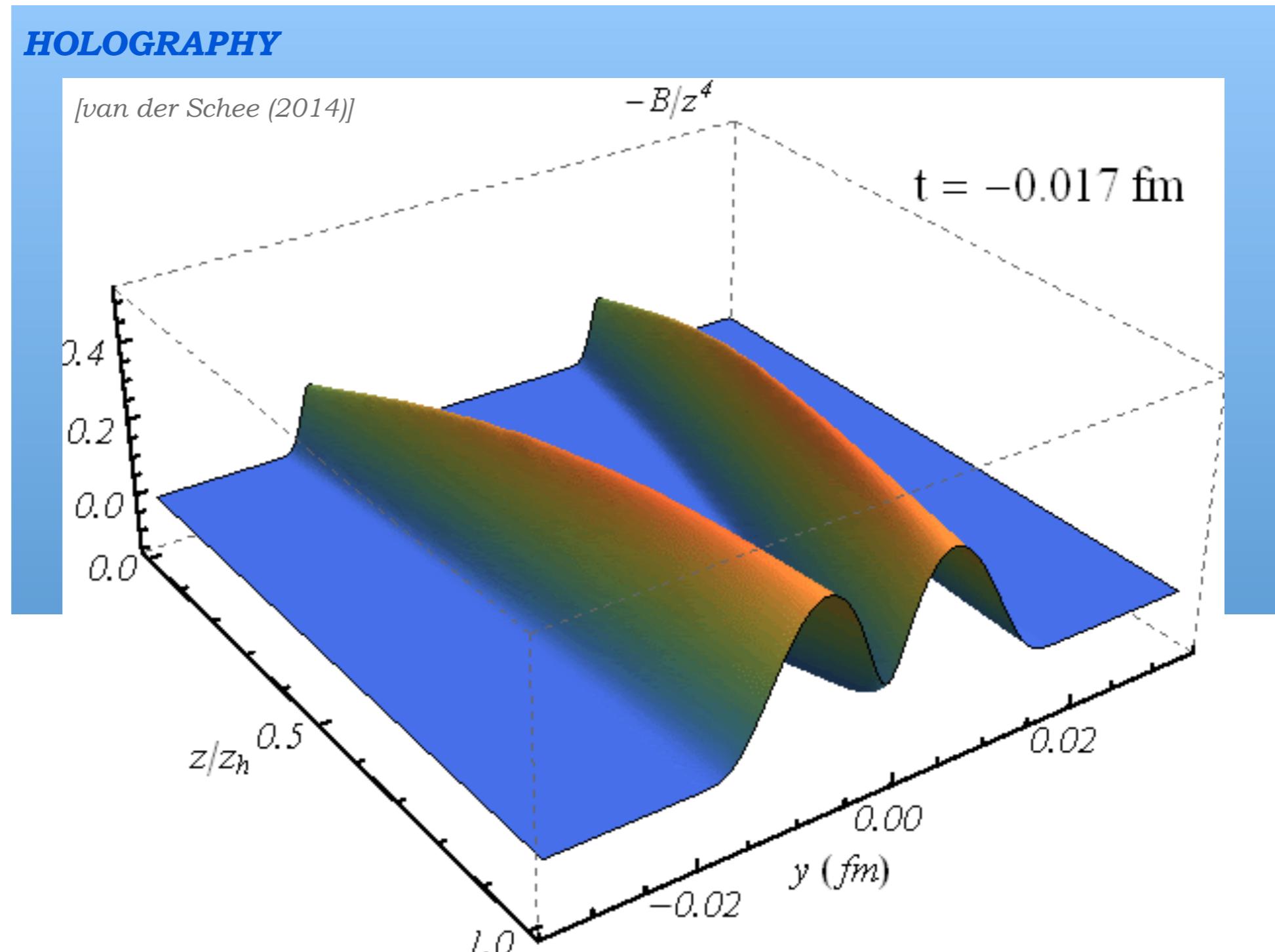
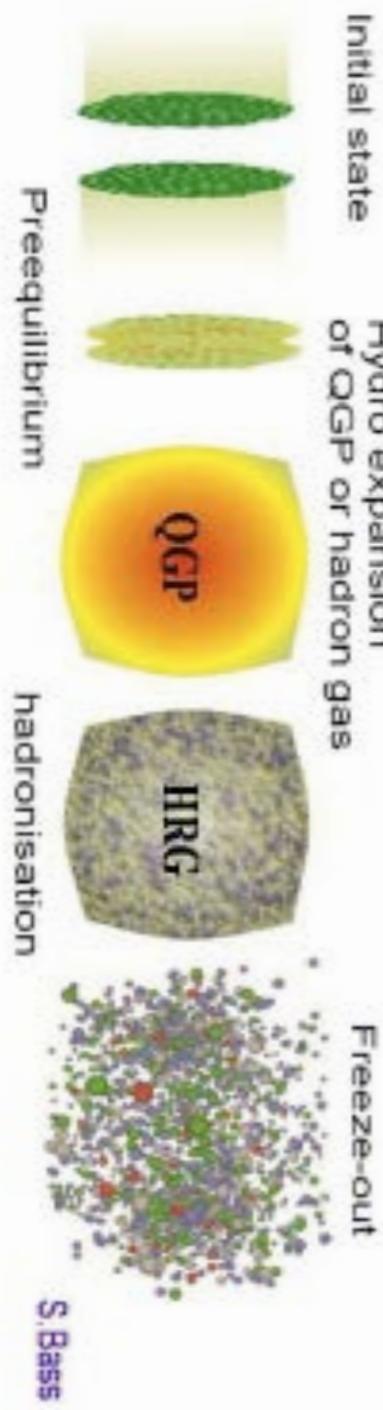
Example:
heavy ion
Collisions



holographic idea: [Janik, Peschanski; PRD (2006)]

Holographic heavy ion collision (numerical, large N)

Example:
heavy ion
Collisions

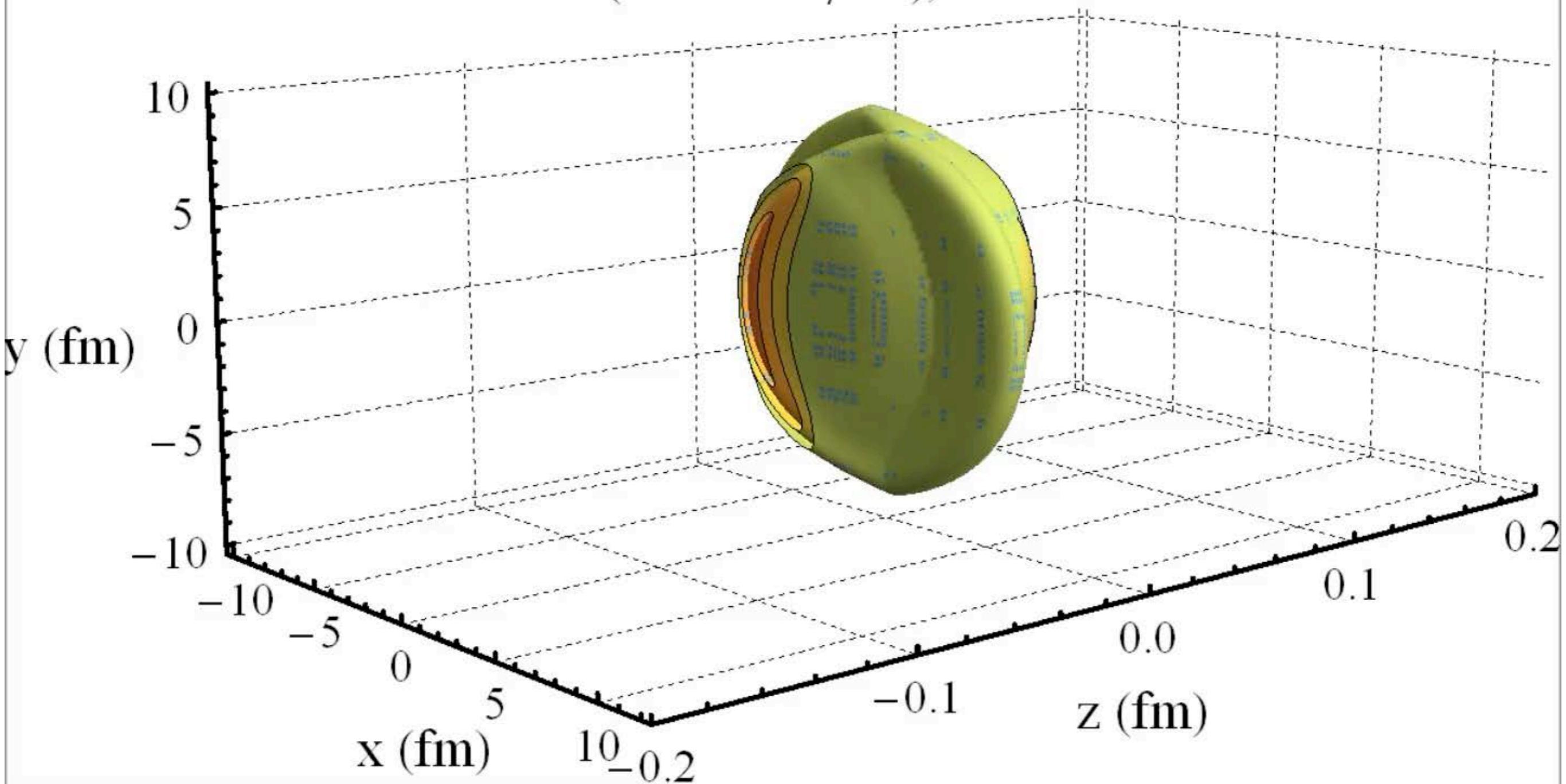


holographic idea: [Janik, Peschanski; PRD (2006)]

Off-center holographic heavy ion collision

[van der Schee (2014)]

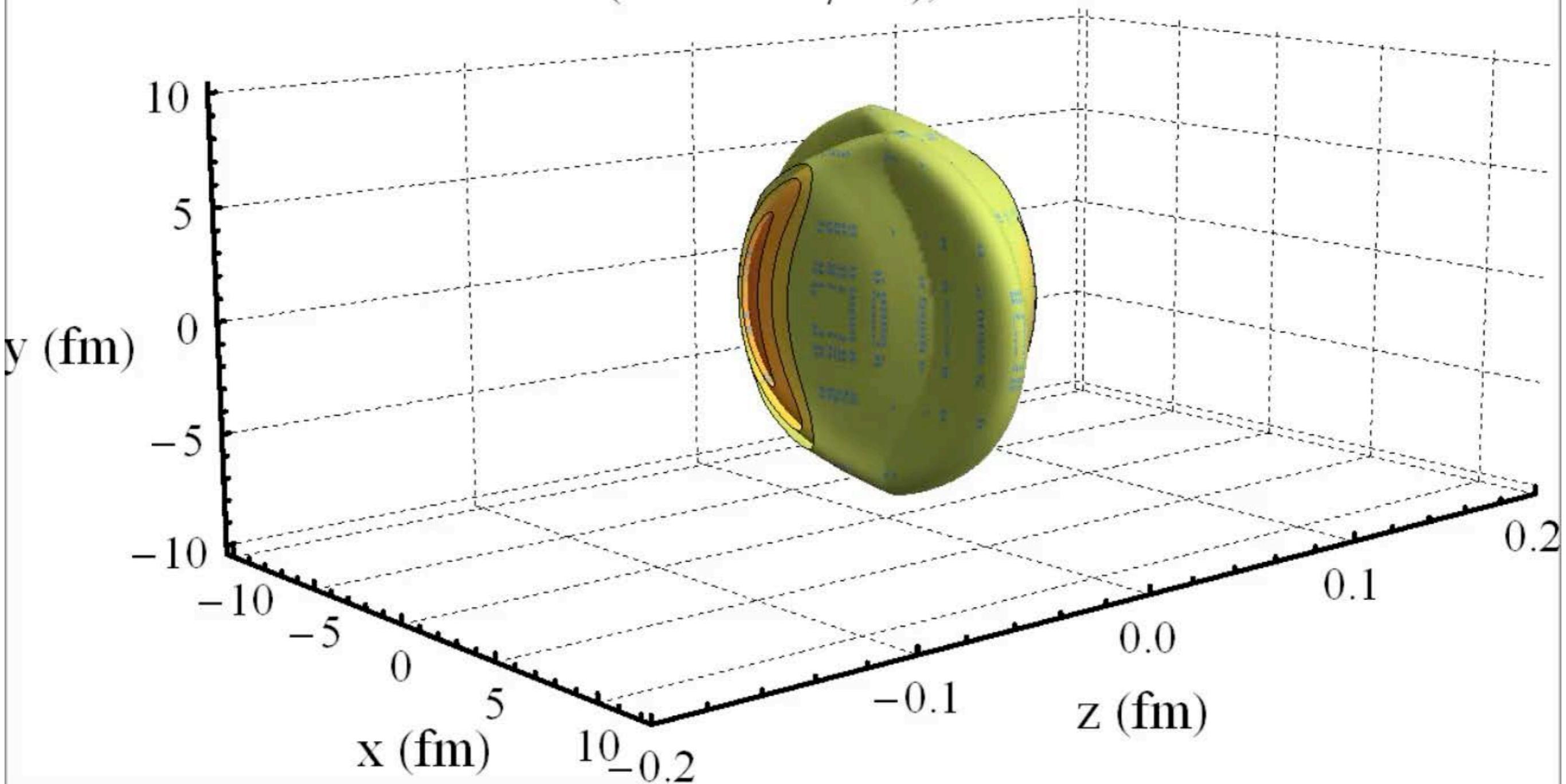
$\mathcal{E} (2 - 50 \text{TeV/fm}^3)$, $t = -0.02 \text{ fm}$



Off-center holographic heavy ion collision

[van der Schee (2014)]

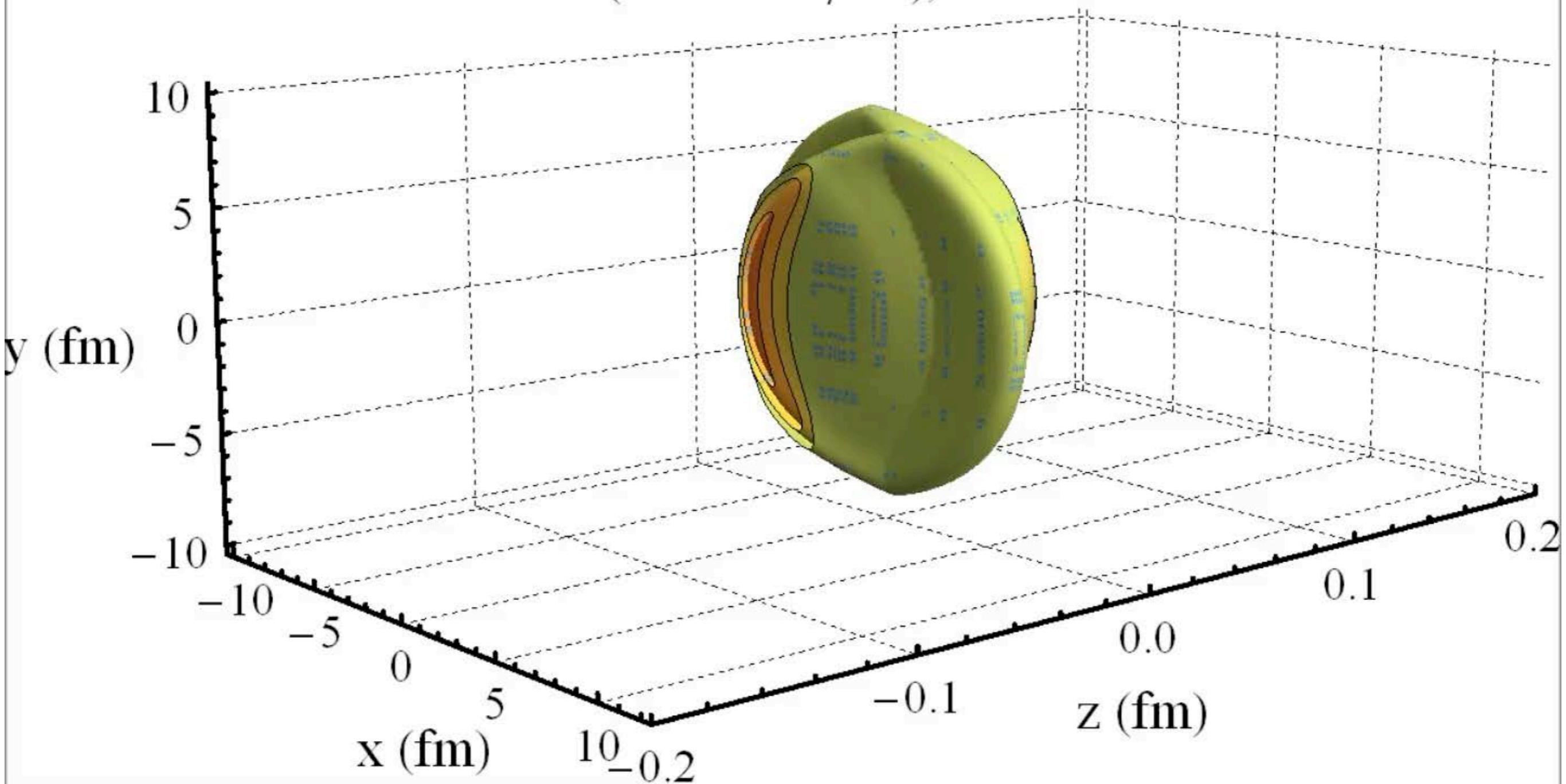
$\mathcal{E} (2 - 50 \text{TeV/fm}^3)$, $t = -0.02 \text{ fm}$



Off-center holographic heavy ion collision

[van der Schee (2014)]

$\mathcal{E} (2 - 50 \text{TeV/fm}^3)$, $t = -0.02 \text{ fm}$

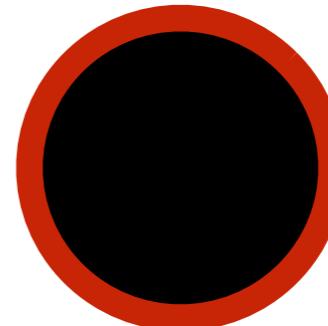


Outline

1. Statistical versus quantum mechanic dynamics
2. Holography (far from equilibrium)
- 3. Holographic entanglement entropy (calculation)**
4. Quantum gravity experiments

Entanglement entropy corresponds to minimal surface

RECALL: Black hole entropy grows as its **surface area** (not as its volume).

$$S \propto A_h$$


Entanglement entropy in quantum field theory corresponds to a particular minimal surface in gravity theory

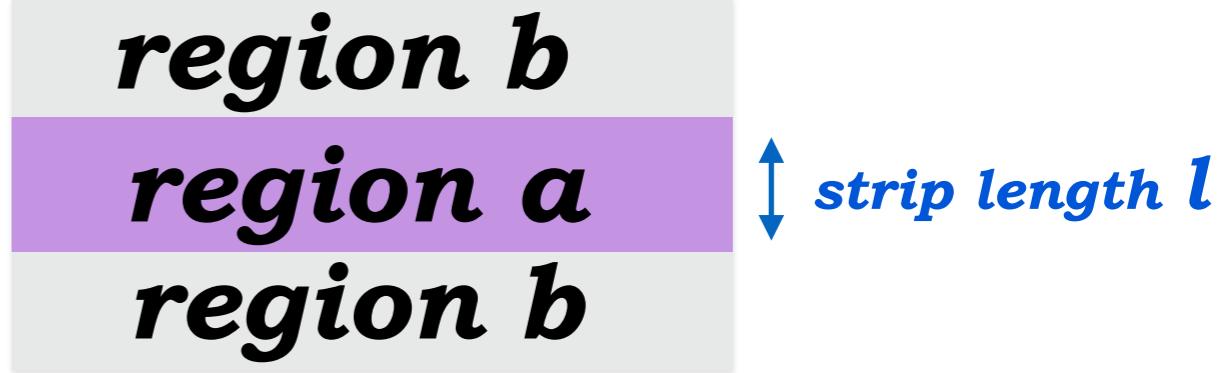
[Ryu, Takayanagi; JHEP (2006)]



Calculation of entanglement entropy in 2D CFT from minimal surface in AdS3

Definition: entanglement entropy

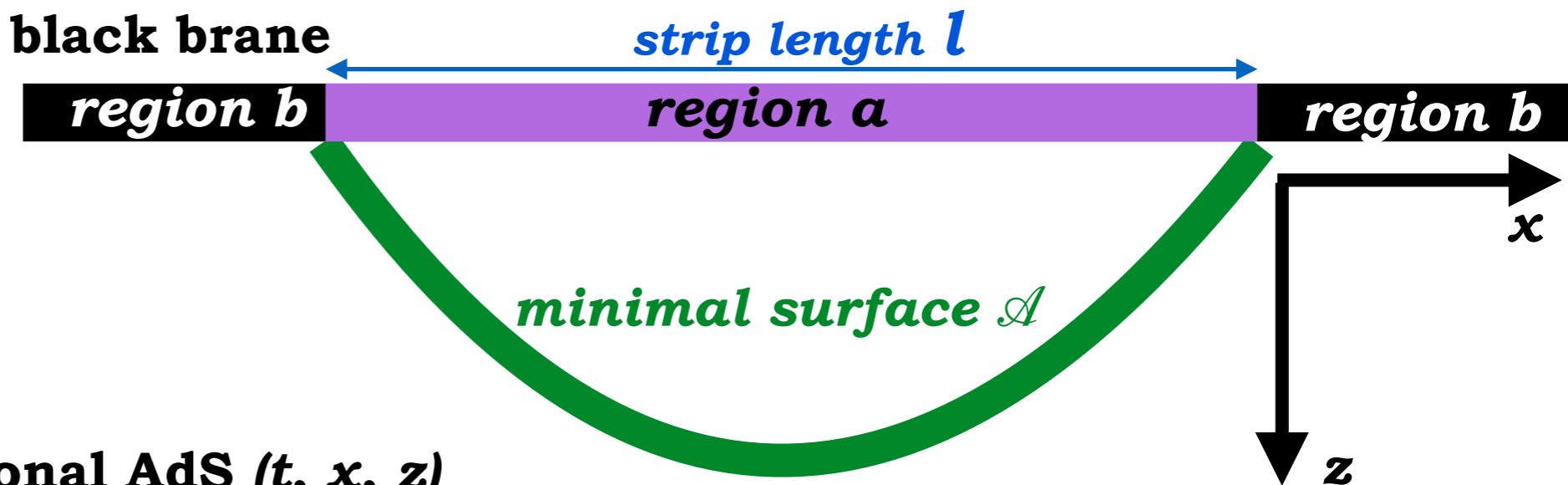
$$S_a = -\text{Tr} \rho_a \log \rho_a, \quad \rho_a = \text{Tr}_b |\psi\rangle \langle \psi|$$



Holographically dual definition

[Ryu, Takayanagi; JHEP (2006)]

$$S_a = \frac{1}{4G} \mathcal{A}$$



**Example: 3-dimensional AdS (t, x, z)
at a given time t ,
minimal surface is shortest path
which is also called a geodesic**

Calculation of entanglement entropy in 2D CFT from minimal surface in AdS3

[Ecker; thesis (2018)]

Example: 3-dimensional AdS in Eddington-Finkelstein coordinates
 $(t, x, z) \rightarrow (v, x, z)$

Calculate minimal surface = shortest path = geodesic

Metric: $ds^2 = \frac{1}{z^2} (-dv^2 - 2dzdv + d\vec{x}^2)$

Clever parametrization of surface:
 $X^\alpha(\sigma) = (Z(\sigma), V(\sigma), X(\sigma))$

Geodesic equation:

$$\ddot{X}^\alpha(\sigma) + \Gamma_{\beta\gamma}^\alpha(X^\delta(\sigma))\dot{X}^\beta(\sigma)\dot{X}^\gamma(\sigma) = J(\sigma)\dot{X}^\alpha(\sigma)$$

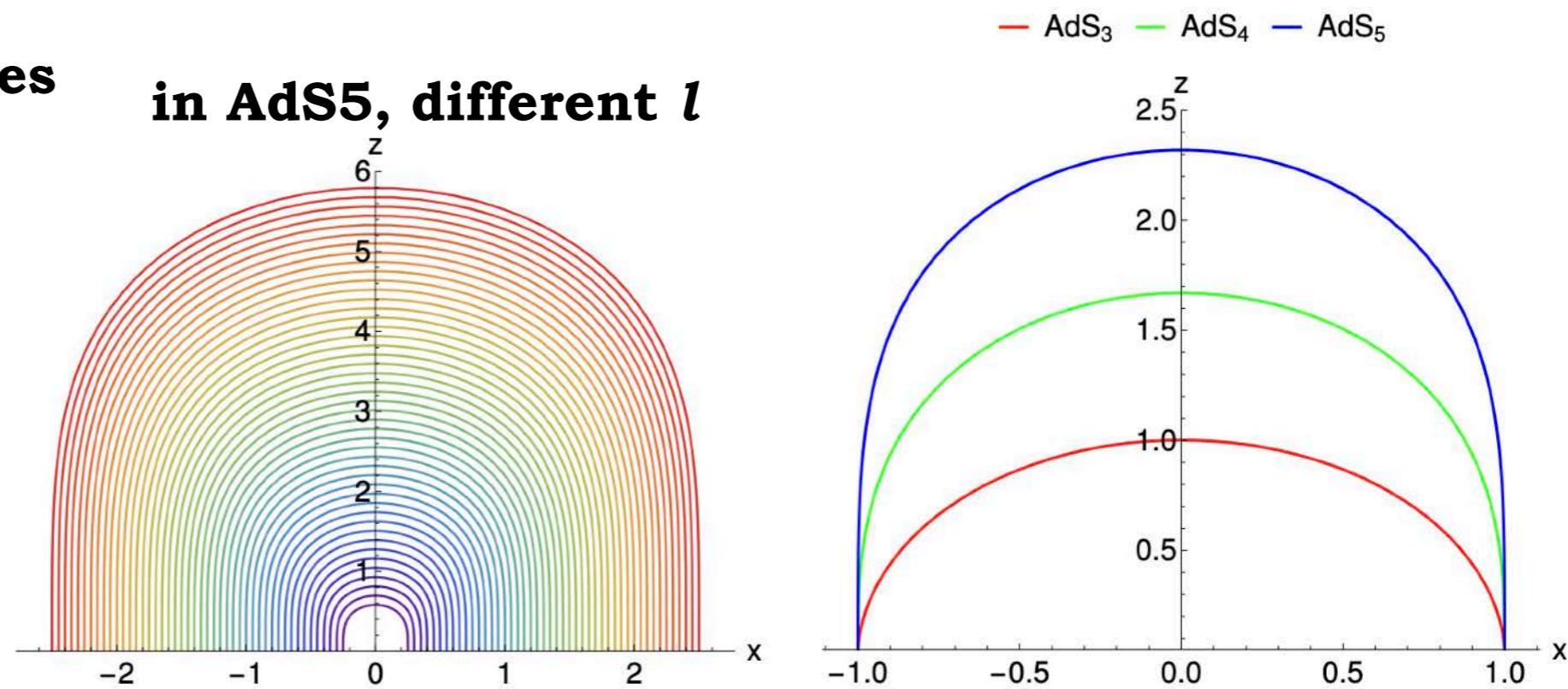
Solution: minimal surfaces

in AdS5, different l

$$Z(\sigma) = \frac{l}{2}(1 - \sigma^2),$$

$$V(\sigma) = v_0 - Z(\sigma),$$

$$X(\sigma) = \frac{l}{2}\sigma\sqrt{2 - \sigma^2}$$



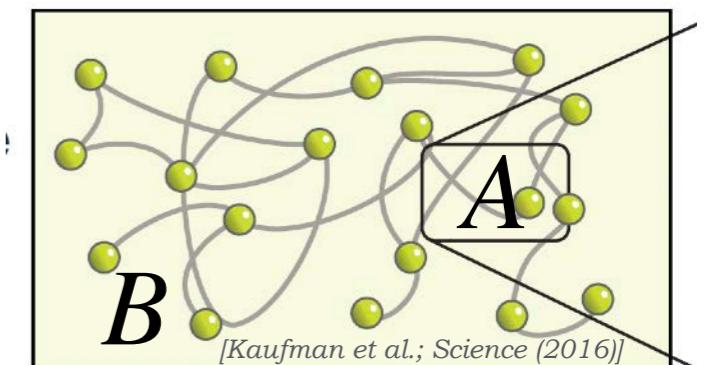
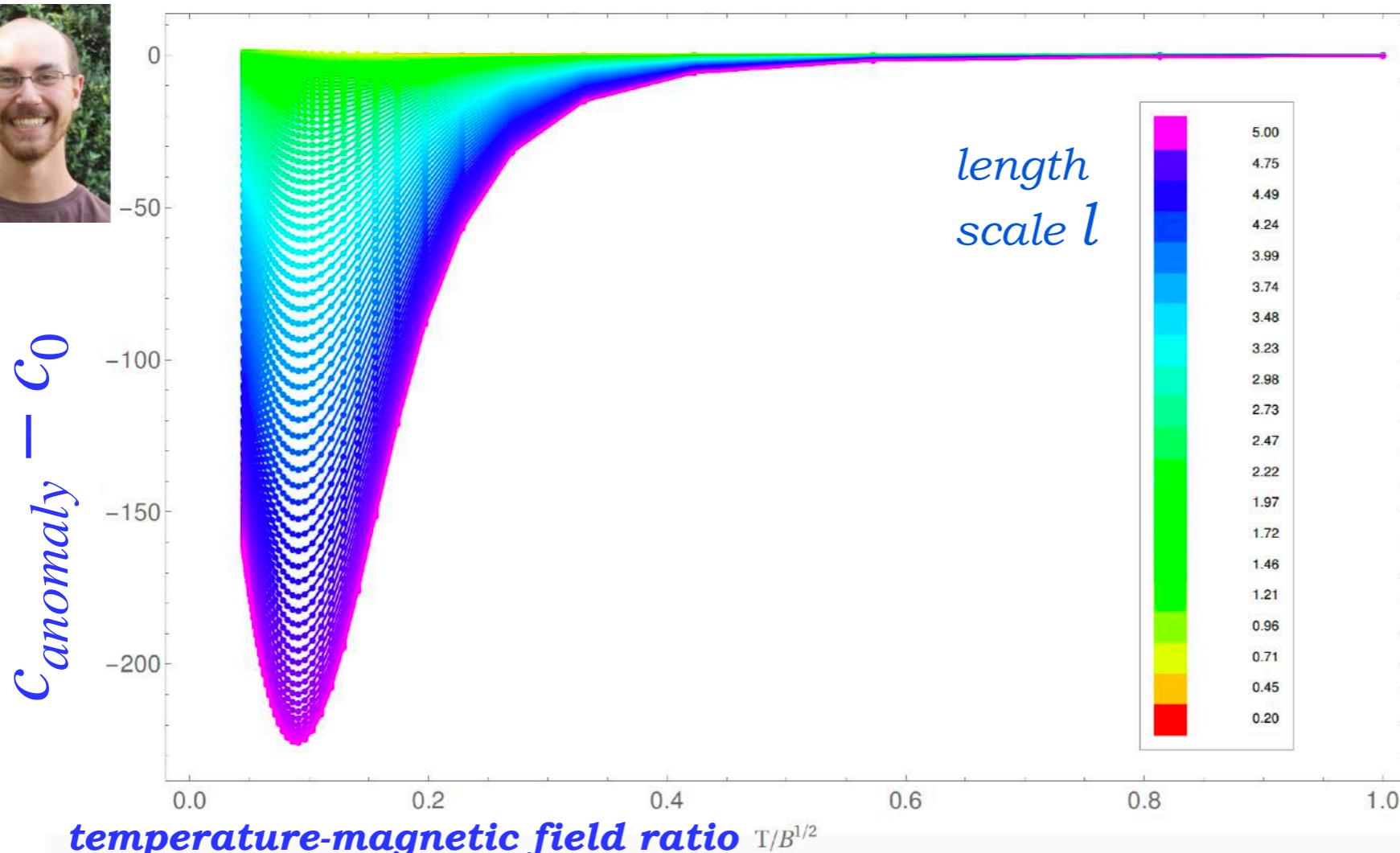
— AdS₃ — AdS₄ — AdS₅

Effect of chiral anomaly and magnetic field on entanglement

Calculation: strongly coupled $N=4$ Super-Yang-Mills theory in strong B ;
compute minimal surfaces in AdS5

Entropic c-function (with anomaly minus without anomaly)

[Cartwright,Kaminski; JHEP (2021)]



Reduced density matrix:
 $\rho_A = \text{tr}_B \rho$

Entanglement entropy:
 $S_A = - \text{tr} (\rho_A \log \rho_A)$

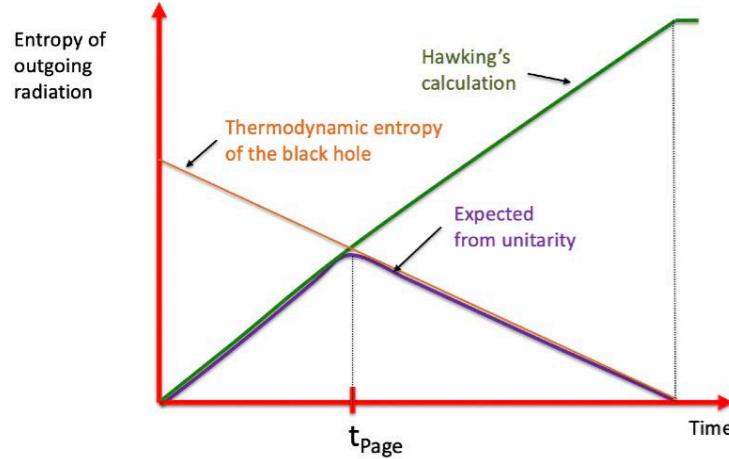
Entropic c-function:
 $c \propto \frac{\partial S_A}{\partial l}$

→anomaly effect peaked at $T/B^{1/2} = 0.1$

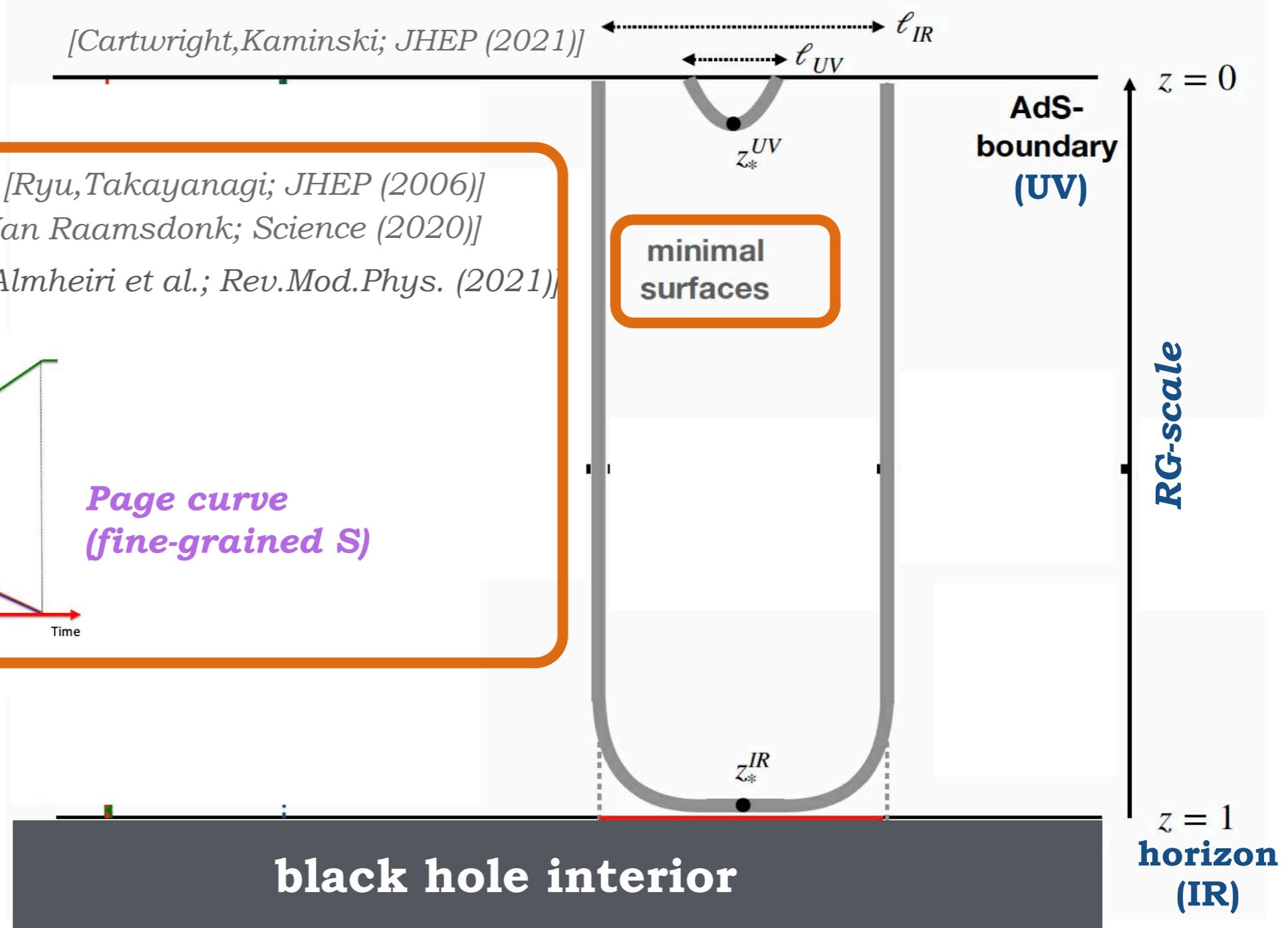
Geometric picture: three faces of *minimal surfaces*

[Cartwright,Kaminski; JHEP (2021)]

- 1) **Entanglement entropy:** [Ryu,Takayanagi; JHEP (2006)]
- 2) **Emergent spacetime:** [Van Raamsdonk; Science (2020)]
- 3) **Fine-grained entropy:** [Almheiri et al.; Rev.Mod.Phys. (2021)]



**Page curve
(fine-grained S)**



wormhole
through
extra
dimension

entangled pair

our universe

→ Entanglement = Spacetime

[Van Raamsdonk; Gen.Rel.Grav. (2010)]

Outline

1. Statistical versus quantum mechanic dynamics

2. Holography (far from equilibrium)

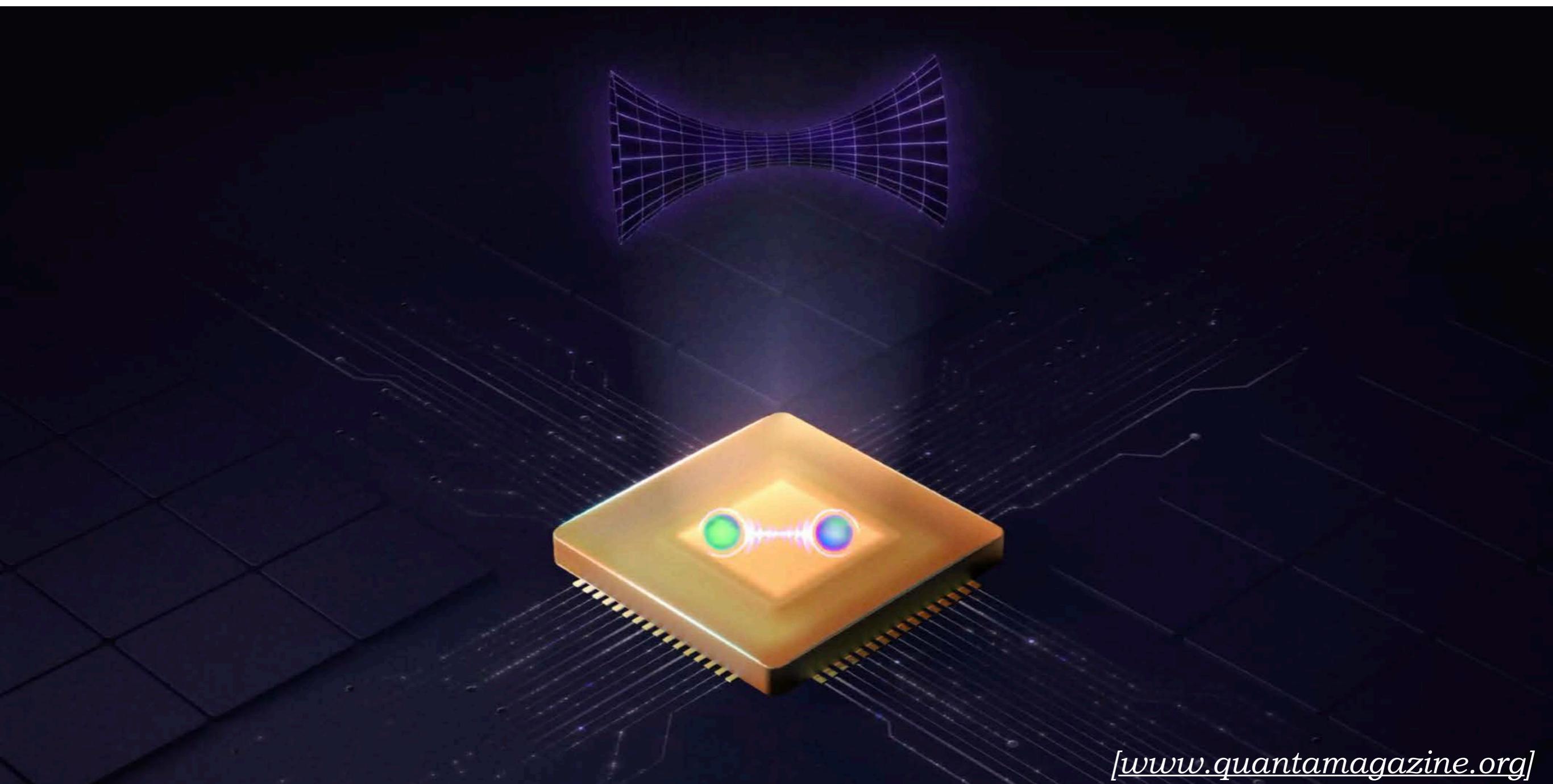
3. Holographic entanglement entropy (calculation)

4. Quantum gravity experiments

- **indirect: traversable wormhole on quantum computer**
- **direct: gauge/gravity correspondence on electric circuit board**

Indirect Experiment: Simulation on Quantum Computer

[Jafferis et al.; *Nature* (2022)]



[\[www.quantamagazine.org\]](http://www.quantamagazine.org)

Examine spacetime with “quantum computing glasses”.

Indirect Experiment: Simulation on Quantum Computer



[www.quantamagazine.org]

[Jafferis et al.; Nature (2022)]

IBM/Quantinuum Competitors:
[Shapoval et al.; Quantum (2023)]

Principle

Quantum gravity problem

send a signal through a wormhole
in Jackiw-Teitelboim (JT) Gravity

**gauge/gravity
correspondence**



Quantum mechanics problem

quantum teleportation protocol in
Sachdev-Ye-Kitaev (SYK) model

**Google's
Sycamore
solves
machine-
simplified
problem**

Signal!



**machine
learning:
learn SYK**

Simplified SYK problem

machine represents problem???
(Hamiltonian of seven Majorana fermions
with five fully-commuting terms)

SYK Model of Majorana Fermions

from [Jafferis et al.; Nature (2022)]

Here, we study the dynamics of traversable wormholes through a many-body simulation of an SYK system of N fermions^{2,3}. The traversable wormhole protocol is equivalent to a quantum teleportation protocol in the large- N semiclassical limit (Fig. 1c). Explicitly, given left and right Hamiltonians H_L and H_R with N Majorana fermions ψ on each side, the SYK model with q couplings is given by

$$H_{L,R} = \sum_{1 \leq j_1 < \dots < j_q \leq N} J_{j_1 \dots j_q} \psi_{L,R}^{j_1} \dots \psi_{L,R}^{j_q}, \quad (1)$$

where the couplings are chosen from a Gaussian distribution with mean zero and variance $J^2(q-1)!/N^{q-1}$. We choose $q=4$ and demonstrate gravitational physics at sufficiently small N , sparsifying $J_{j_1 \dots j_q}$ to enable experimental implementation.

Applying the learning process, we produce a large population of sparse Hamiltonians showing the appropriate interaction sign dependence (Fig. 2a). We select the learned Hamiltonian

$$H_{L,R} = -0.36\psi^1\psi^2\psi^4\psi^5 + 0.19\psi^1\psi^3\psi^4\psi^7 - 0.71\psi^1\psi^3\psi^5\psi^6 \\ + 0.22\psi^2\psi^3\psi^4\psi^6 + 0.49\psi^2\psi^3\psi^5\psi^7, \quad (3)$$

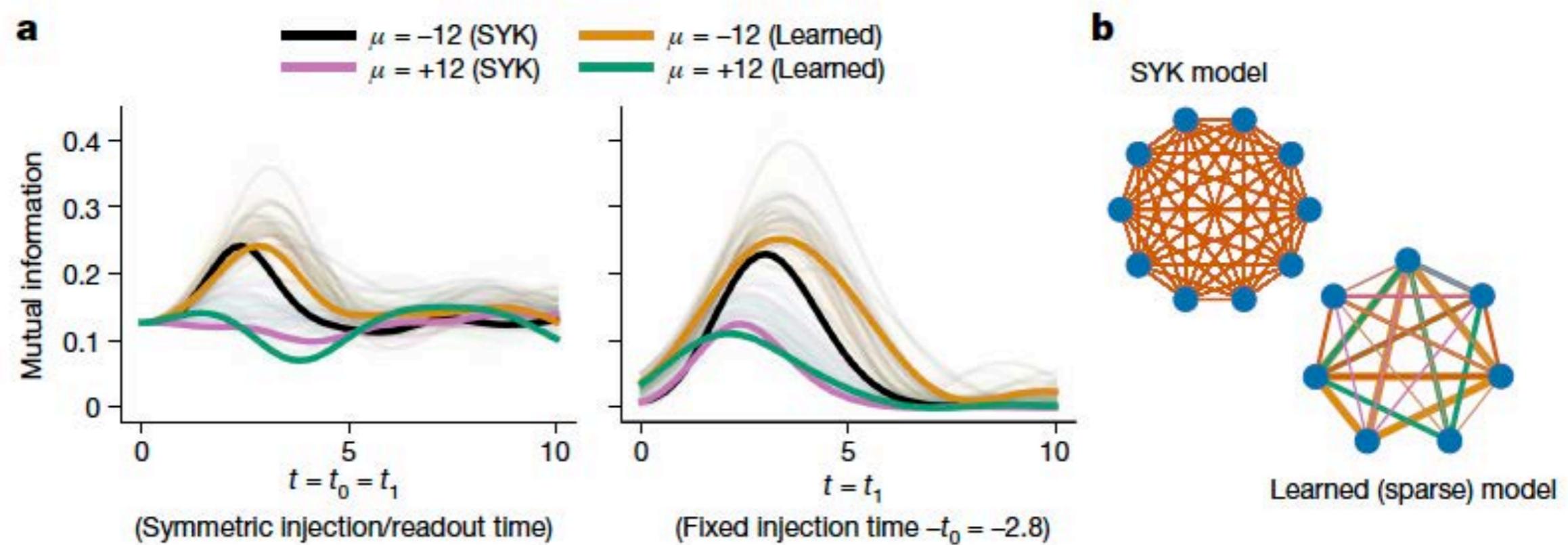
which requires seven of the original $N=10$ SYK model fermions, where ψ^i denotes the Majorana fermions of either the left or the right systems.

Machine Learning (ML): Simplify the Problem

[Jafferis et al.; Nature (2022)]

Use of ML

- use machine learning techniques to construct a sparsified SYK model,
experimentally realized with 164 two-qubit gates on a nine-qubit circuit



Potential shortcomings of Jafferis et al.'s experiment



[Jafferis et al.; *Nature* (2022)]

[\[www.quantamagazine.org\]](http://www.quantamagazine.org)

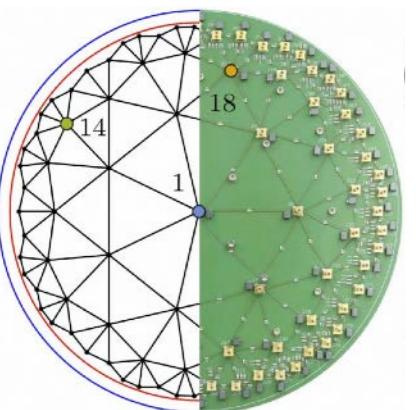
Comment criticizing this experiment

[Kobrin/Schuster/Yao; preprint (2023)]

- **Problem 1:** learned Hamiltonian does not exhibit thermalization
- **Problem 2:** resembles SYK only for operators used in ML training
- **Problem3:** perfect size winding is generic feature of small-size models

see also response to criticism: [Jafferis et al.; (2023)]

Direct Experiment: Black Hole on Electrical Circuit



Principle

Quantum gravity problem

send a signal through a wormhole
in Jackiw-Teitelboim (JT) Gravity

gauge/
gravity
corre-
spondence



Quantum mechanics problem

quantum teleportation protocol in
Sachdev-Ye-Kitaev (SYK) model

Quantum gravity
quantum
information on
quantum circuit

Signal!

Google's
Sycamore
solves
machine-
simplified
problem

machine
learning:
learn SYK

Simplified SYK problem

machine represents problem???

Direct Experiment: Black Hole on Electrical Circuit

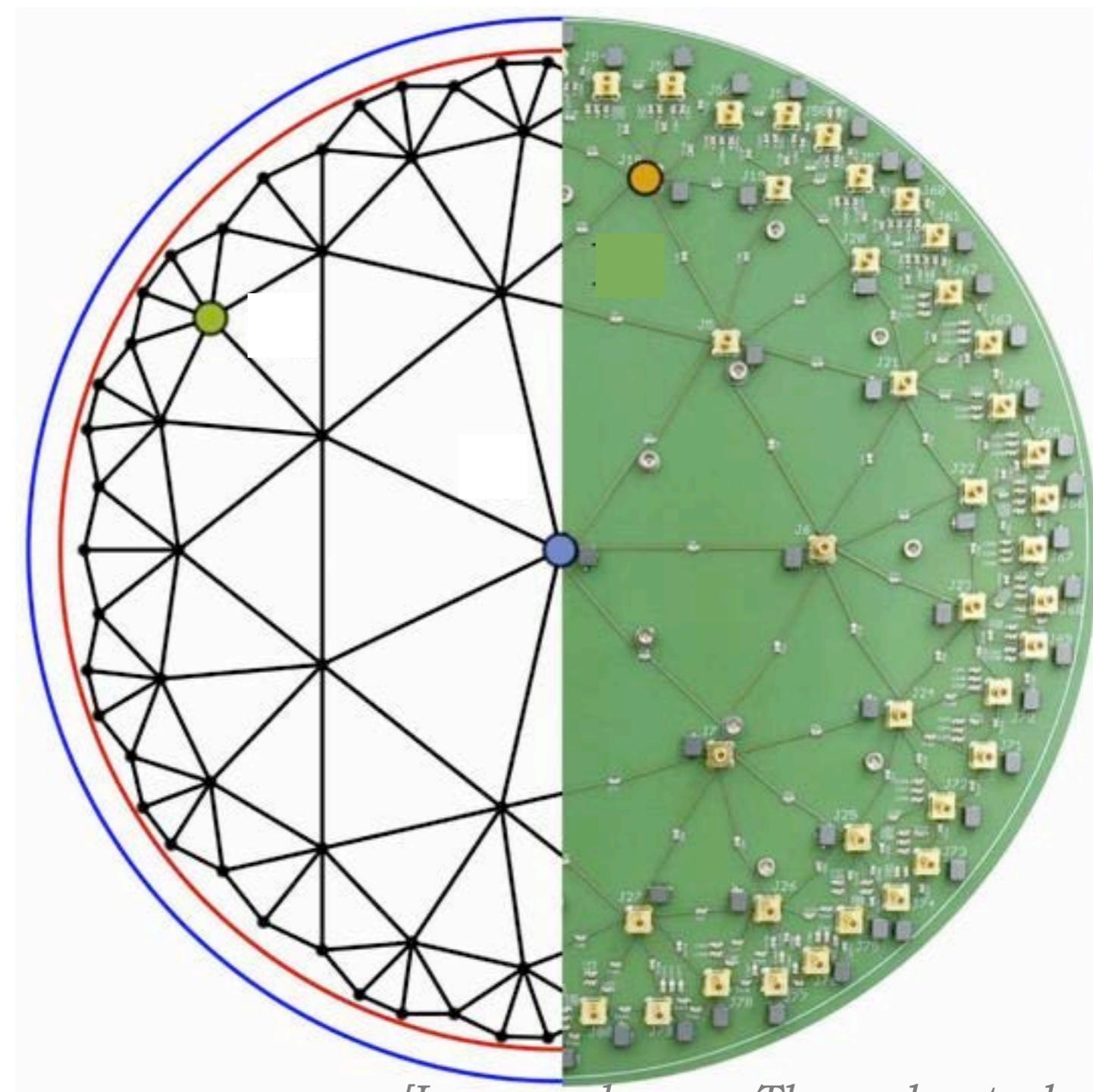
[Dey, Chen, Kaminski, et al.; PRL (2024)]

Model black hole on electrical hyperbolic circuit board

Voltage on circuit satisfies Klein-Gordon equation for massive scalar in AdS

[Basteiro et al.; PRL (2023)]

Testing holography in the lab

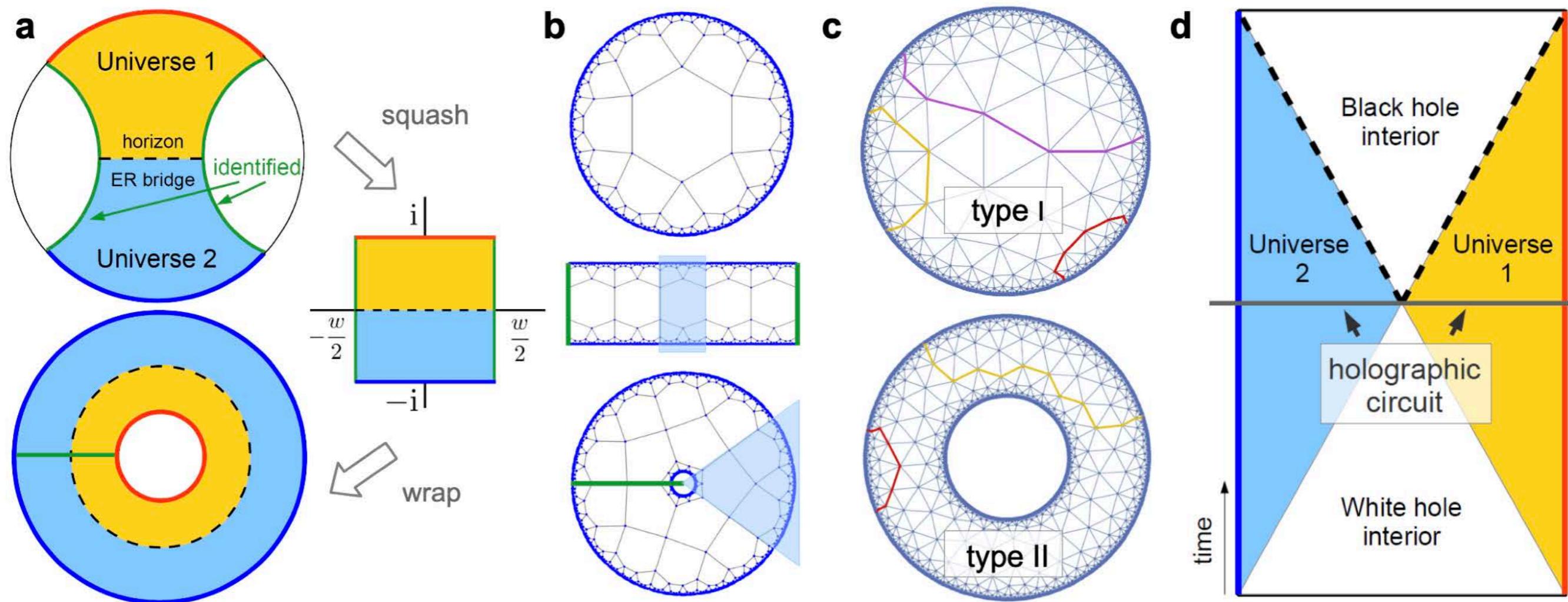


[Lenggenhager, Thomale et al.,
Nat. Commun. (2022)]

Direct Experiment: Black Hole on Electrical Circuit

[Dey, Chen, Kaminski, et al.; PRL (2024)]

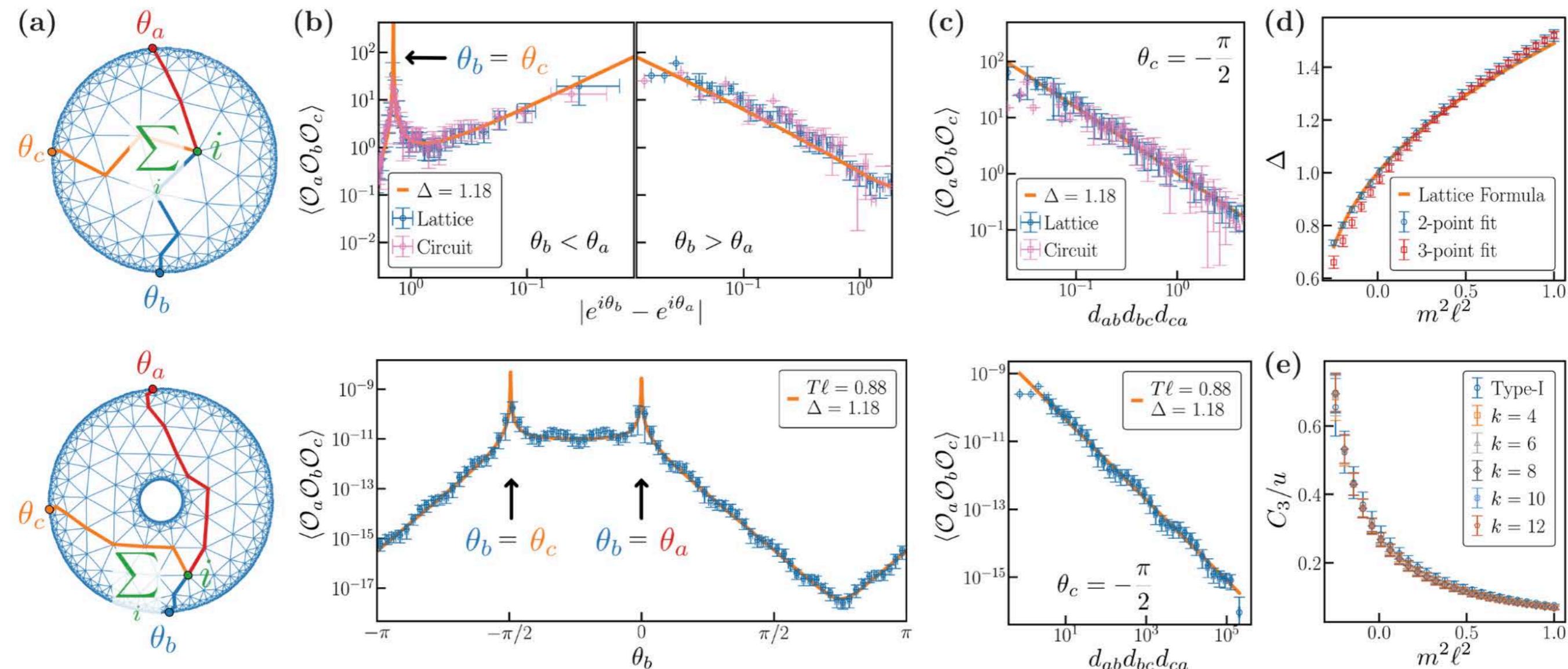
Realize wormhole on classical electric circuit:



Direct Experiment: Black Hole on Electrical Circuit

[Dey, Chen, Kaminski, et al.; PRL (2024)]

Three point functions measured on circuit (gravity side):



CFT expectation (gauge side):

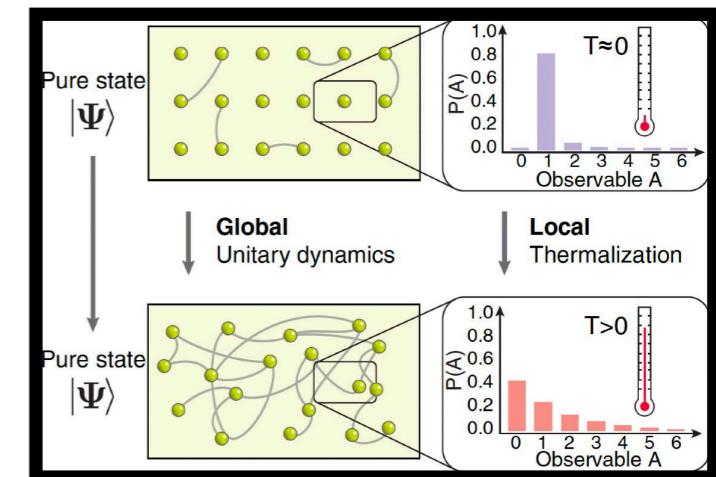
$$\langle \mathcal{O}_a \mathcal{O}_b \mathcal{O}_c \rangle \simeq \frac{C_3}{(d_{ab} d_{ac} d_{bc})^\Delta}$$

$$d_{ab} = \begin{cases} |e^{i\theta_a} - e^{i\theta_b}| & \text{(type-I)} \\ \frac{\sinh(\pi T\ell |\theta_a - \theta_b|)}{\pi T\ell} & \text{(type-II)} \end{cases}$$

Discussion

SUMMARY

- isolated quantum systems thermalize in similar ways
- entanglement = spacetime



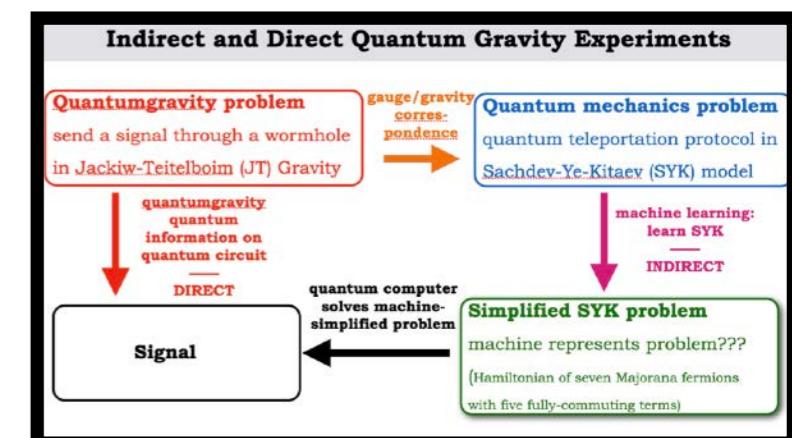
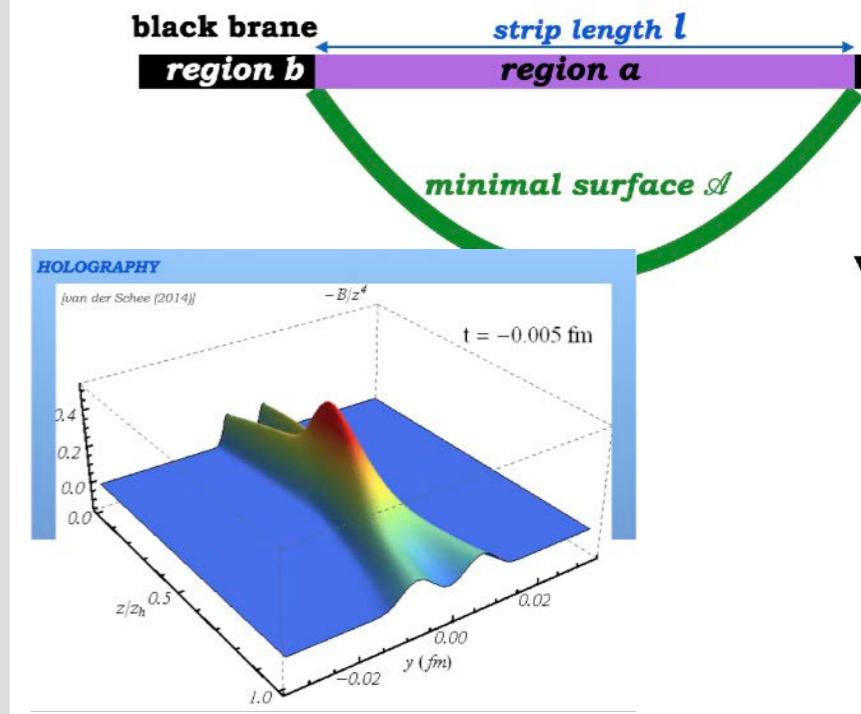
OUTLOOK

- entanglement entropy in holography far-from-equilibrium
- finite N corrections: get closer to few-body dynamics
- demonstrate “entanglement=spacetime” in experiments

(spacetime emerging from entangled quantum bits?)

- improve indirect traversable wormhole on quantum computer
- directly simulate quantum gravity on quantum computer
- use machine learning methods (experiments, ML spacetime)

→ talk by Jane Kim



Indirect and Direct Quantum Gravity Experiments

Quantumgravity problem

send a signal through a wormhole
in Jackiw-Teitelboim (JT) Gravity

gauge/gravity
corre-
spondence



Quantum mechanics problem

quantum teleportation protocol in
Sachdev-Ye-Kitaev (SYK) model

quantum gravity
quantum
information on
quantum circuit

—
DIRECT

Signal

quantum computer
solves machine-
simplified problem

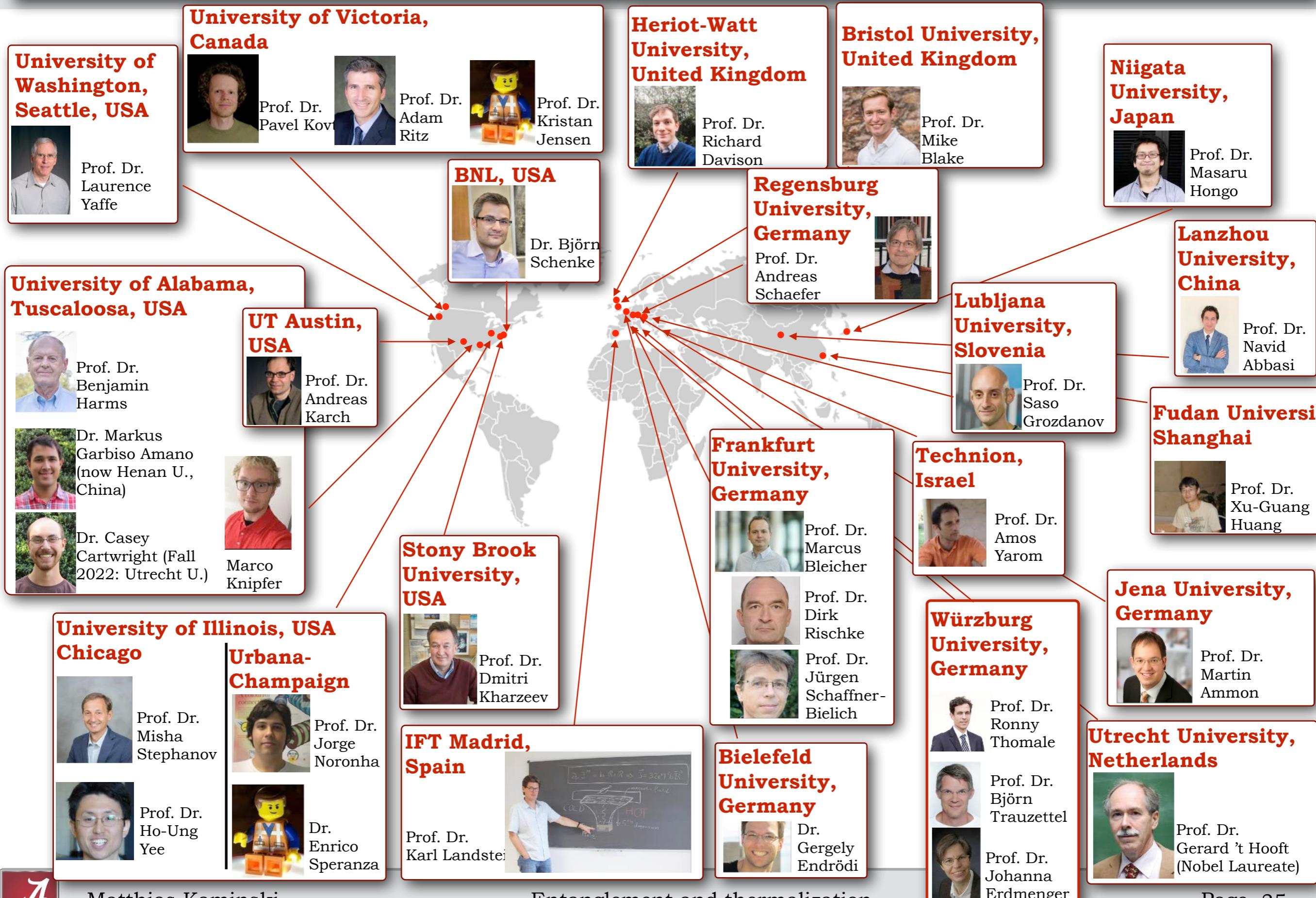


machine learning:
learn SYK
—
INDIRECT

Simplified SYK problem

machine represents problem???
(Hamiltonian of seven Majorana fermions
with five fully-commuting terms)

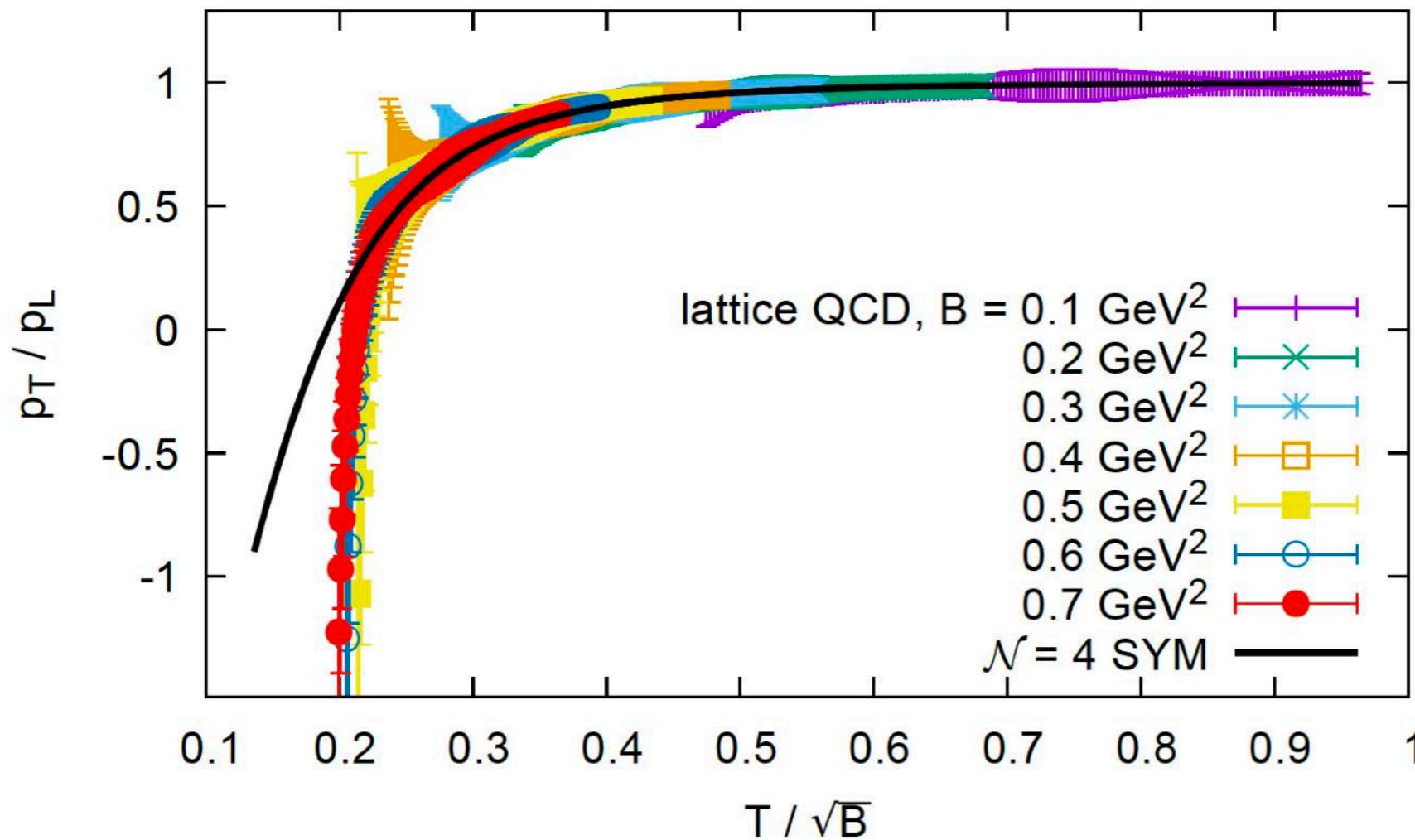
Thanks to my collaborators (since 2012)



APPENDIX

APPENDIX: Universal magneto response in LQCD and N=4 SYM with magnetic field

[Endrödi, Kaminski, Schäfer, Wu, Yaffe; JHEP (2018)]



Lattice QCD with 2+1 flavors, dynamical quarks,
physical masses

*transverse
pressure:*

*longitudinal
pressure:*

$$p_T = -\frac{L_T}{V} \frac{\partial F_{\text{QCD}}}{\partial L_T}$$

$$p_L = -\frac{L_L}{V} \frac{\partial F_{\text{QCD}}}{\partial L_L}$$

F_{QCD} ... free energy
 L_T ... transverse system size
 L_L ... longitudinal system size
 V ... system volume

Direct Experiment: Black Hole on Electrical Circuit

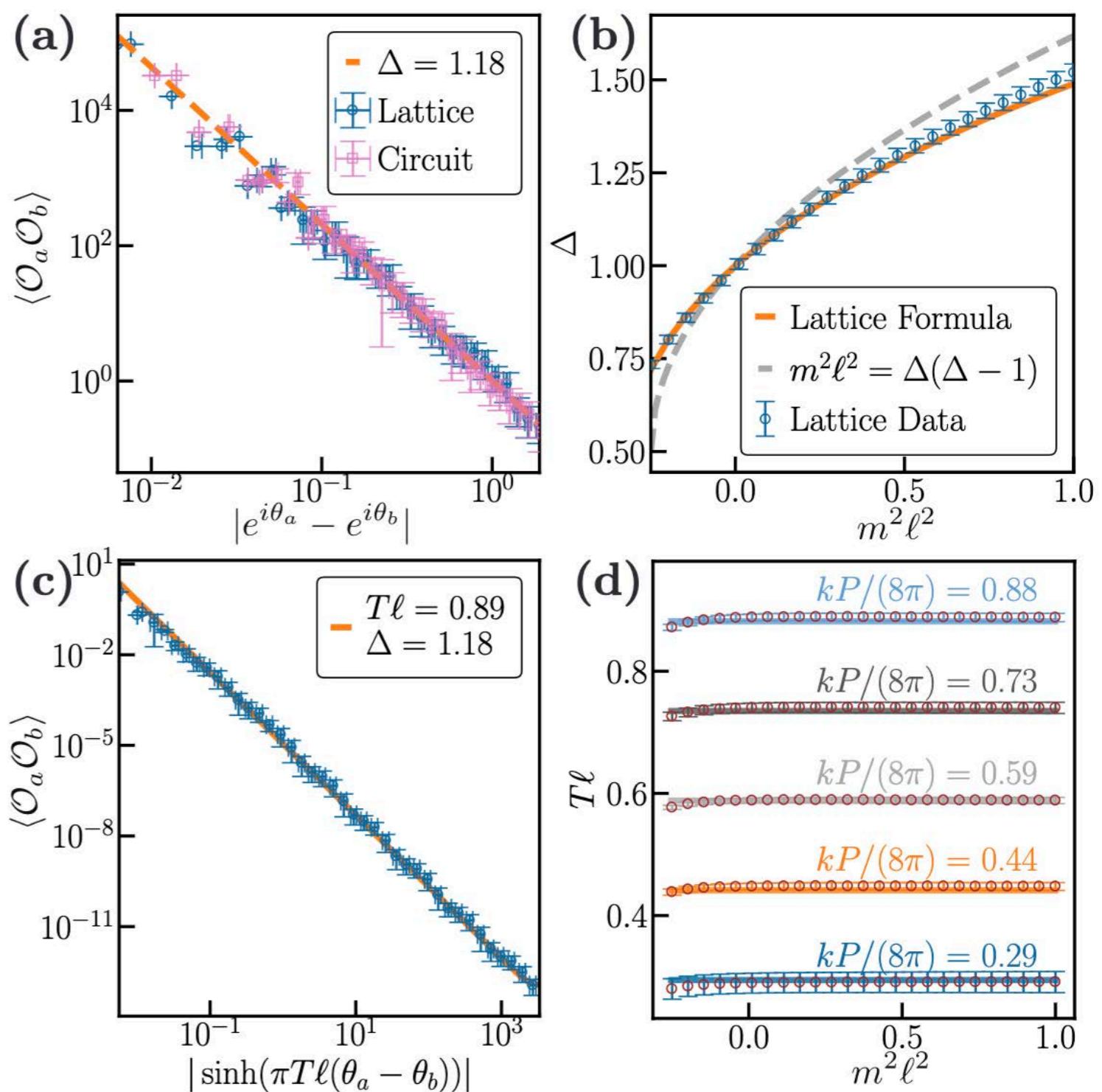
[Dey, Chen, Kaminski, et al.; PRL (2024)]

CFT expectation
(gauge side):

$$\langle \mathcal{O}_a \mathcal{O}_b \rangle \simeq \frac{1}{(d_{ab})^{2\Delta}}$$

$$d_{ab} = \begin{cases} |e^{i\theta_a} - e^{i\theta_b}| & \text{(type-I)} \\ \frac{\sinh(\pi T \ell |\theta_a - \theta_b|)}{\pi T \ell} & \text{(type-II)} \end{cases}$$

Two point functions measured
on circuit (gravity side):



c-Function

**Zamolodchikov's
c-theorem in 2D**

$$(c)_{\text{UV}} \geq (c)_{\text{IR}}$$

energy-momentum tensor: $\langle T^a_a \rangle = -\frac{c}{12}R$
trace anomaly

[Zamolodchikov; JETP Lett.(1986)]

**c-theorem in 4D
(the a-theorem)**

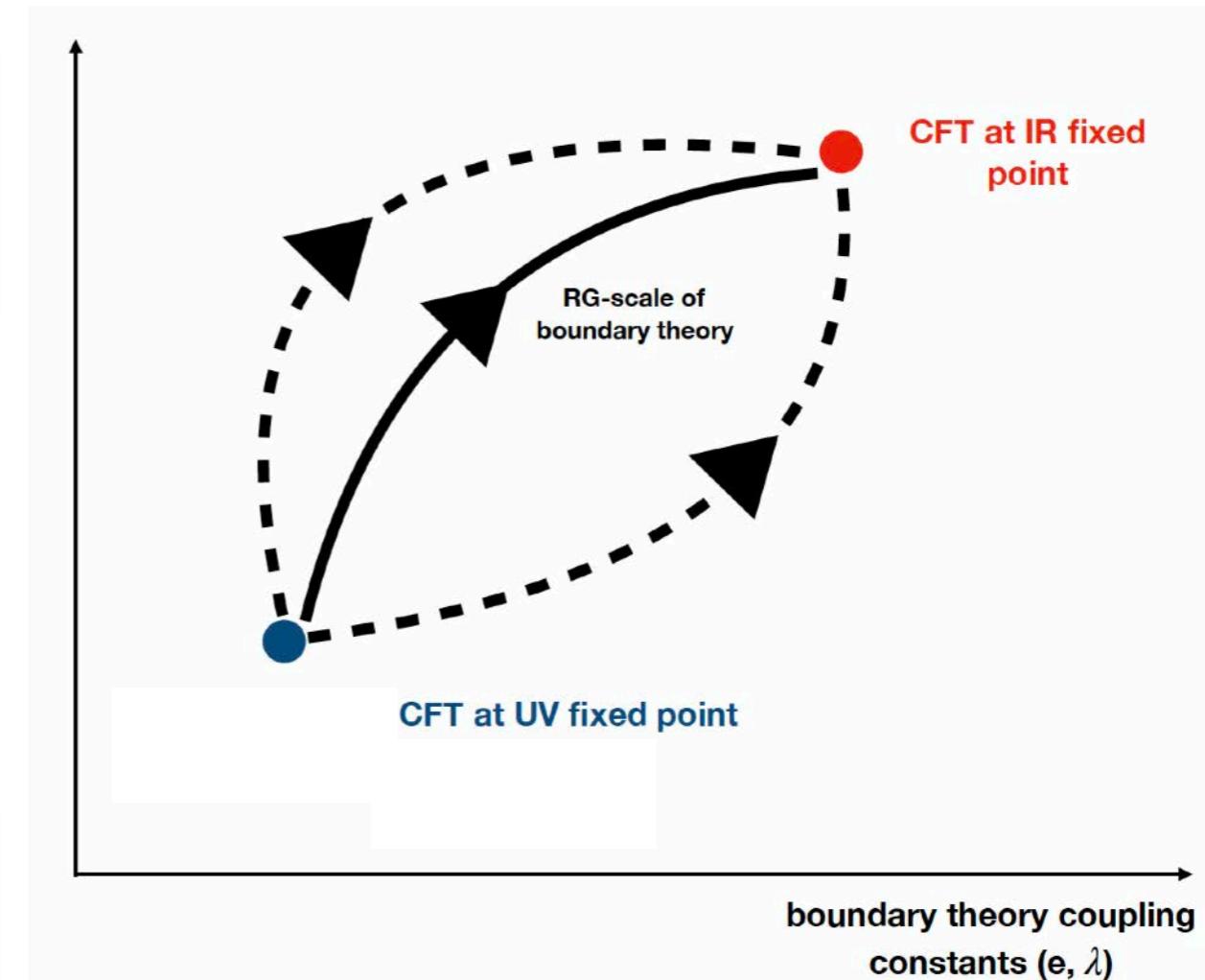
$$(a_4)_{\text{UV}} - (a_4)_{\text{IR}}$$

$$\langle T^b_b \rangle = \frac{c_{TT}}{16\pi^2}C^2 - \frac{a_4}{16\pi^2}\mathcal{E} - \frac{1}{4}F^2$$

[Komargodski, Schwimmer; (2011)]

[Cardy; Phys.Lett.B(1988)]

[Osborn; Phys.Lett.B(1988)]



- IR/UV fixed points: c-function equals central charge of IR/UV CFT
- c-function measures degrees of freedom
- take a CFT: c-function constant

Entropic c-function

2D

$$c_2 = 3\ell \frac{\delta S_a}{\delta \ell}$$

*entanglement
entropy*

[Casini, Huerta; Phys.Lett.B (2004)]

ℓ : length scale (inverse energy scale)

4D

$$a_4 = \beta_4 \frac{\ell^3}{H^2} \frac{\partial S_a}{\partial \ell}$$

[Nishioka, Takayanagi; JHEP (2007)]

[Myers, Sinha; JHEP (2011)]

H : IR-regulator

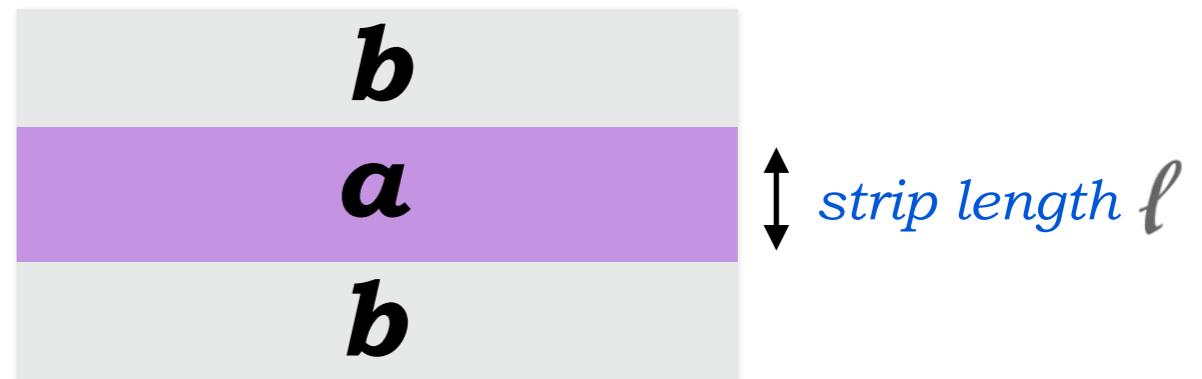
β_4 : known constant

→ c-function defined by
entanglement entropy

Holographic entanglement entropy

Definition

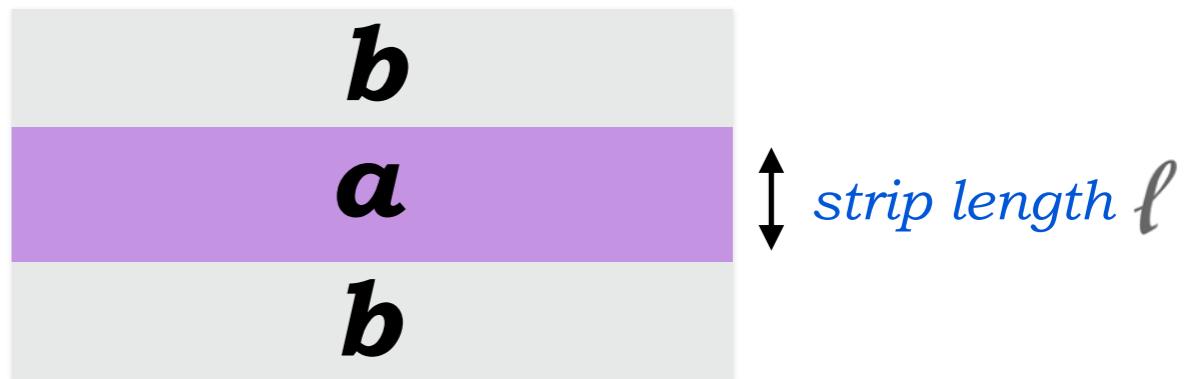
$$S_a = -\text{Tr} \rho_a \log \rho_a, \quad \rho_a = \text{Tr}_b |\psi\rangle \langle \psi|$$



Holographic entanglement entropy

Definition

$$S_a = -\text{Tr} \rho_a \log \rho_a, \quad \rho_a = \text{Tr}_b |\psi\rangle \langle \psi|$$

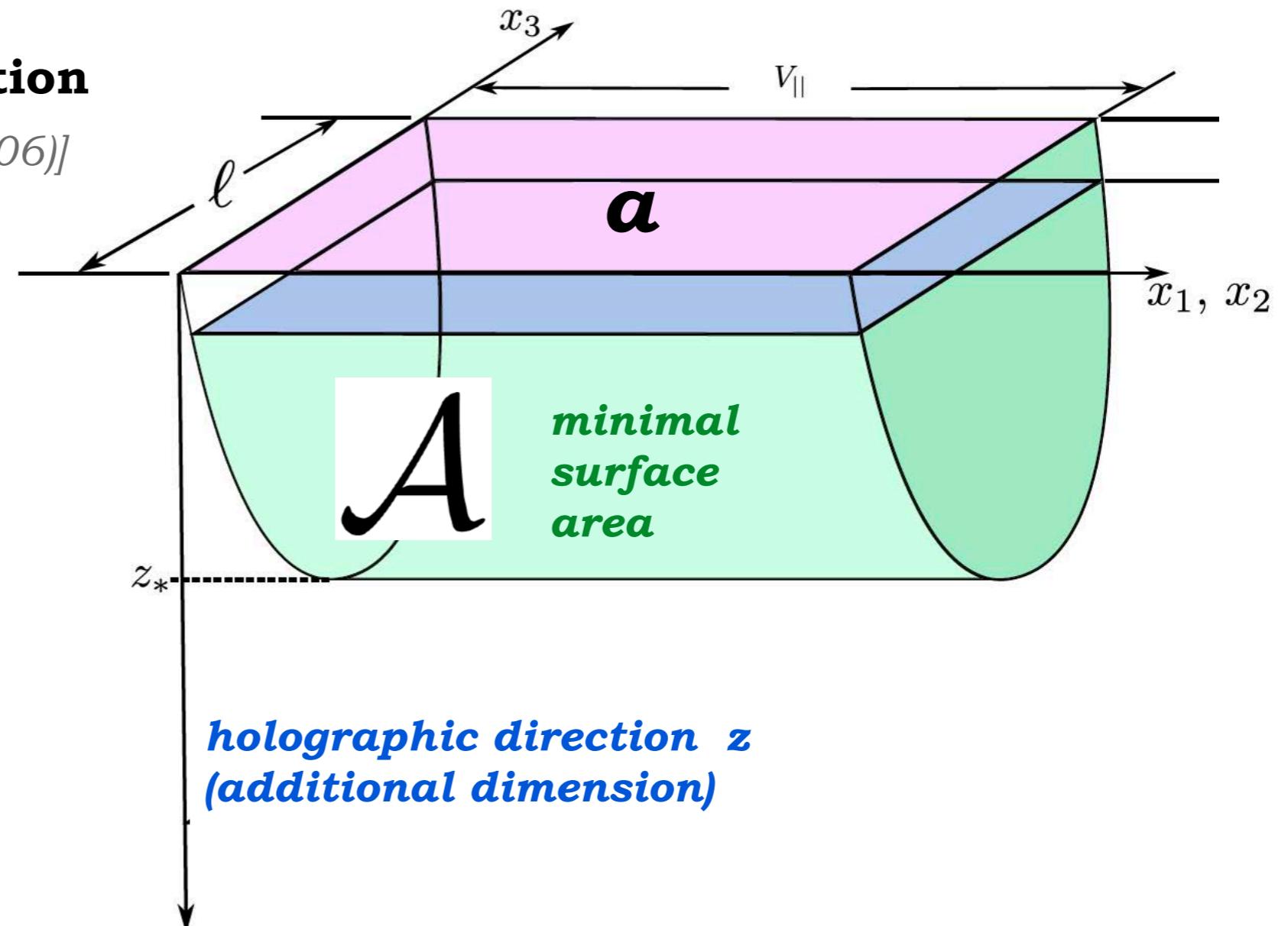


Holographically dual definition

[Ryu, Takayanagi; JHEP (2006)]

$$S_a = \frac{1}{4G_5} \mathcal{A}$$

G_5 is the 5-dimensional gravitational constant of Anti de Sitter spacetime



Gravity dual to $N=4$ SYM theory with magnetic field



Einstein-Maxwell-Chern-Simons action

$$S_{grav} = \frac{1}{2\kappa^2} \left[\int_{\mathcal{M}} d^5x \sqrt{-g} \left(R + \frac{12}{L^2} - \frac{1}{4} F_{mn} F^{mn} \right) - \frac{\gamma}{6} \int_{\mathcal{M}} A \wedge F \wedge F \right]$$

5-dimensional Einstein-Maxwell action encodes $N=4$ Super-Yang-Mills theory with axial **$U(1)$ gauge symmetry**

5-dimensional Chern-Simons term **encodes chiral anomaly**

Einstein-Maxwell equations

$$R_{\mu\nu} + 4g_{\mu\nu} = \frac{1}{2} \left(F_{\mu\alpha} F_{\nu}^{\alpha} - \frac{1}{6} g_{\mu\nu} F_{\alpha\beta} F^{\alpha\beta} \right)$$

$$\nabla_{\mu} F^{\mu\nu} = -\frac{\gamma}{8\sqrt{-g}} \epsilon^{\nu\alpha\beta\lambda\sigma} F_{\alpha\beta} F_{\lambda\sigma} .$$

Solution: charged magnetic black brane metric

[D'Hoker, Kraus; JHEP (2010)]

- magnetic extension of a (charged) Reissner-Nordstrom black brane

$$ds^2 = \frac{1}{z^2} \left(\frac{dz^2}{U(z)} - U(z) dt^2 + v(z)^2 (dx_1^2 + dx_2^2) + w(z)^2 (dx_3^2 + c(z) dt)^2 \right)$$

with numerically known solutions for U, v, w, c

Gravitational *calculation*

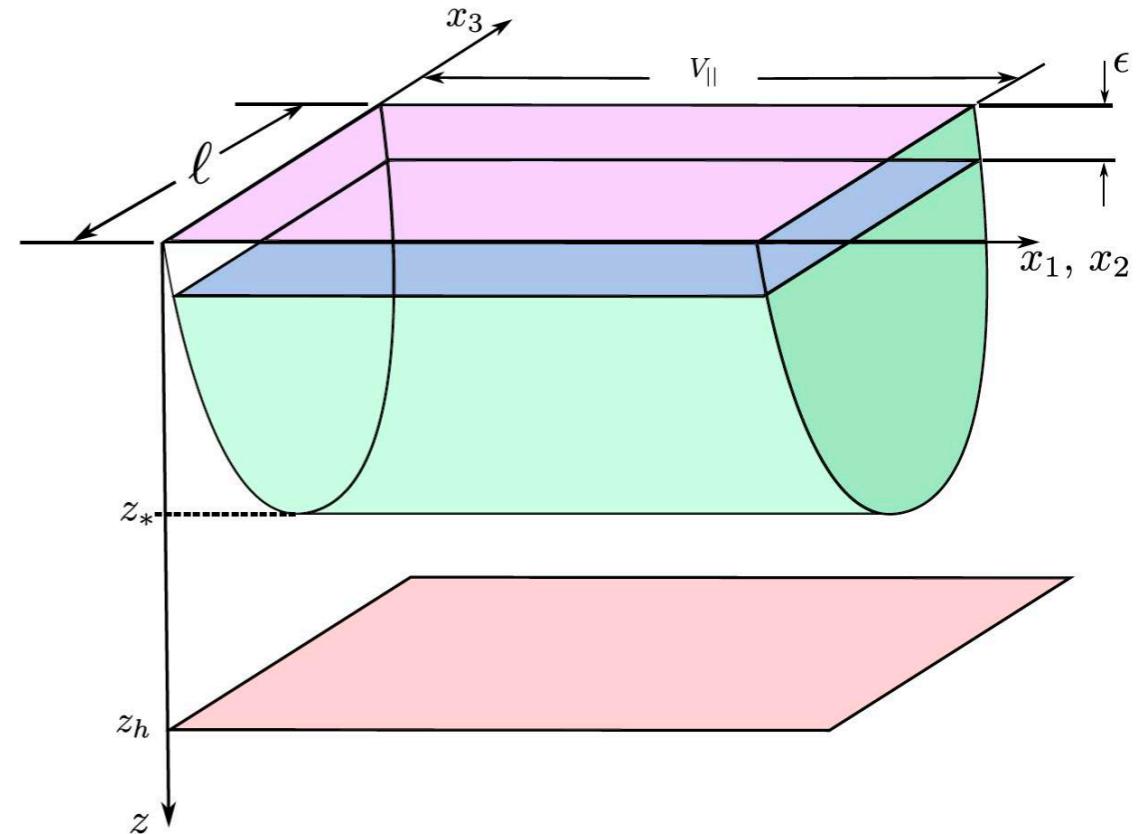


→ calculate a *geodesic* in conformally deformed AdS metric

	Transverse	Longitudinal
Embedding Coordinates	$\chi^\mu = (z(\sigma), t(\sigma), x_1(\sigma), x_2, x_3)$	$\chi^\mu = (z(\sigma), t(\sigma), x_1, x_2, x_3(\sigma))$
Surface Coordinates	$\sigma^i = (\sigma, x_2, x_3)$	$\sigma^i = (\sigma, x_1, x_2)$

Recall:

$$S_a = \frac{1}{4G_5} \mathcal{A}$$



Entanglement entropy

$$S_a = \frac{1}{4G_5} V_{\parallel} \int d\sigma \sqrt{\frac{v(z(\sigma))^4 \left(w(z(\sigma))^2 x'_3(\sigma)^2 + \frac{z'(\sigma)^2}{U(z(\sigma))} \right)}{z(\sigma)^6}}$$

$$V_{\parallel} = \int_{-a}^a \int_{-b}^b dx_1 dx_3$$

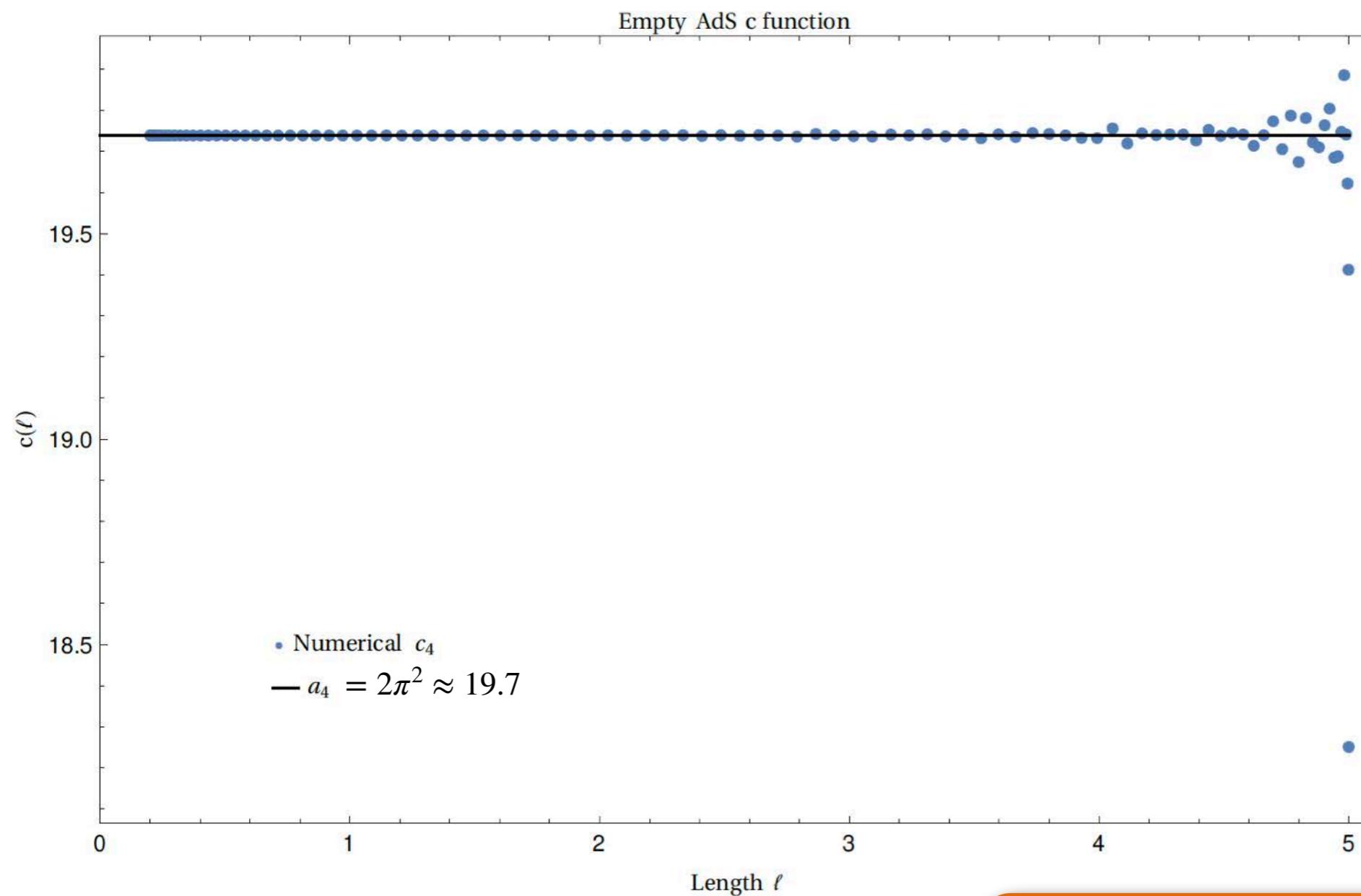
minimal surface area \mathcal{A}

Reminder: metric is

$$ds^2 = \frac{1}{z^2} \left(\frac{dz^2}{U(z)} - U(z) dt^2 + v(z)^2 (dx_1^2 + dx_2^2) + w(z)^2 (dx_3^2 + c(z)dt)^2 \right)$$

Entropic c-function *in N=4 SYM vacuum state*

[Cartwright, Kaminski; arXiv: 2107.12409]



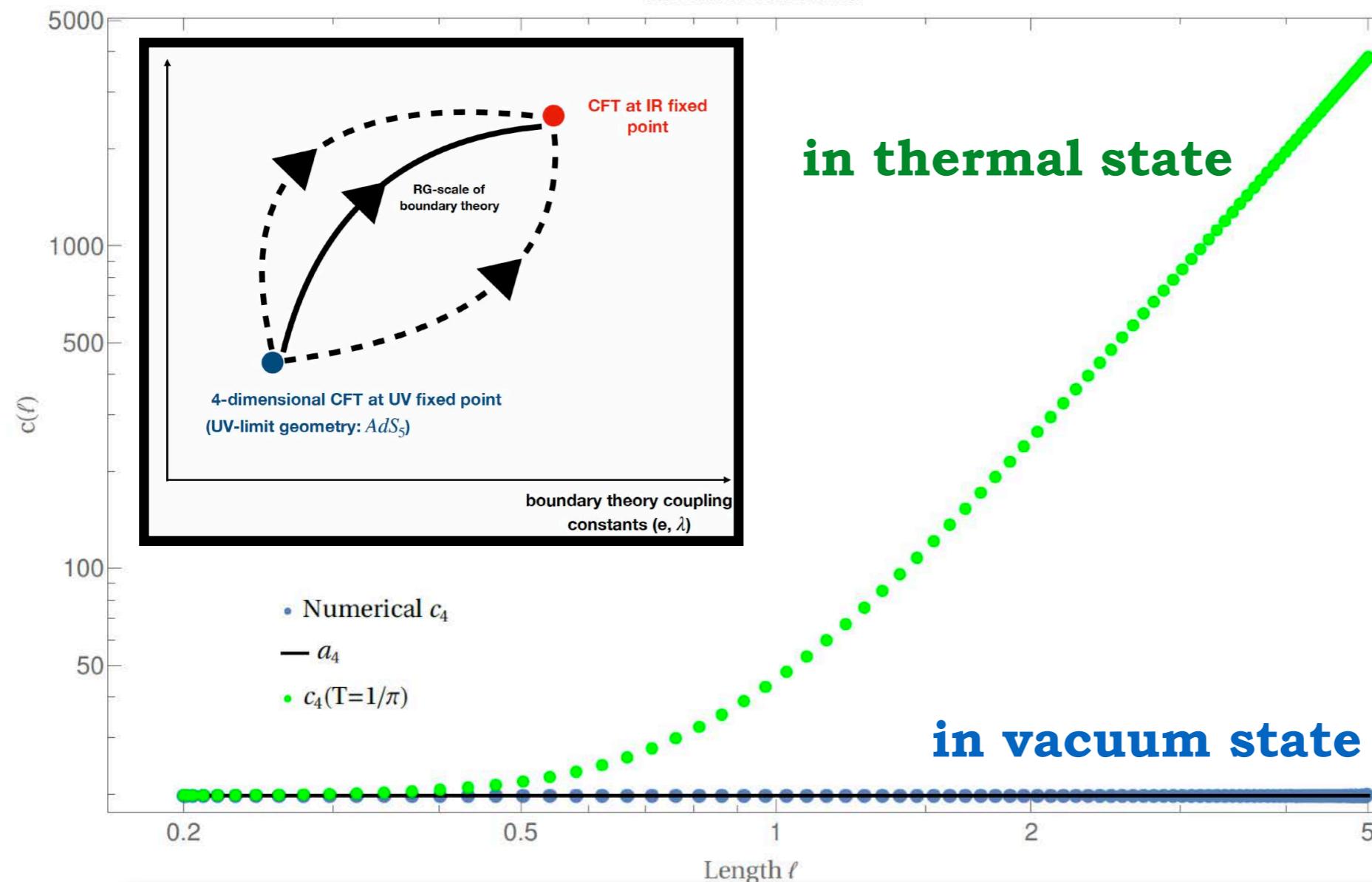
**zero temperature,
no magnetic field,
vanishing charge**

→ **c-function at all scales
equal to central charge of
 $N=4$ SYM, which is a CFT**

Entropic c-function *increases* in *thermal state*

Thermal c function

[Cartwright,Kaminski; JHEP (2021)]



$$a_4 = \beta_4 \frac{\ell^3}{H^2} \frac{\partial S_a}{\partial \ell}$$

[Casini,Huerta; Phys.Lett.B (2004)]

[Nishioka,Takayanagi; JHEP (2007)]

[Myers,Sinha; JHEP (2011)]

H : IR-regulator

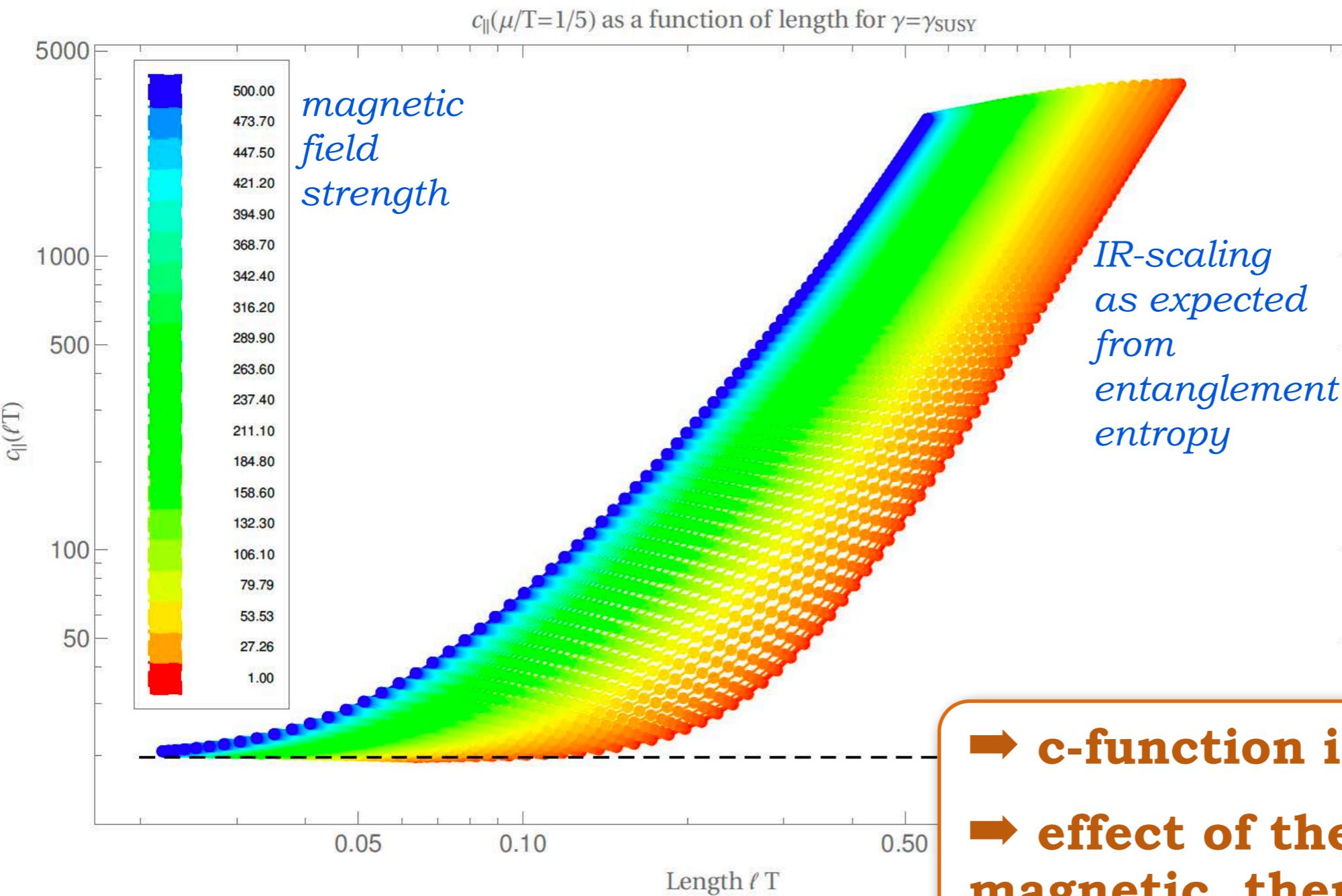
β_4 : known constant

- ➡ c-function increases in thermal state (violates c-theorem?)
- ➡ constant in vacuum of CFT (c-theorem valid)
- ➡ IR limit: thermal entropy

[Zamolodchikov; JETP Lett.(1986)]
 [Komargodski,Schwimmer; (2011)]
 [Osborn; Phys.Lett.B(1988)]
 [Cardy; Phys.Lett.B(1988)]

Entropic c-Function *increases* in *thermal state*

[Cartwright, Kaminski; arXiv: 2107.12409]

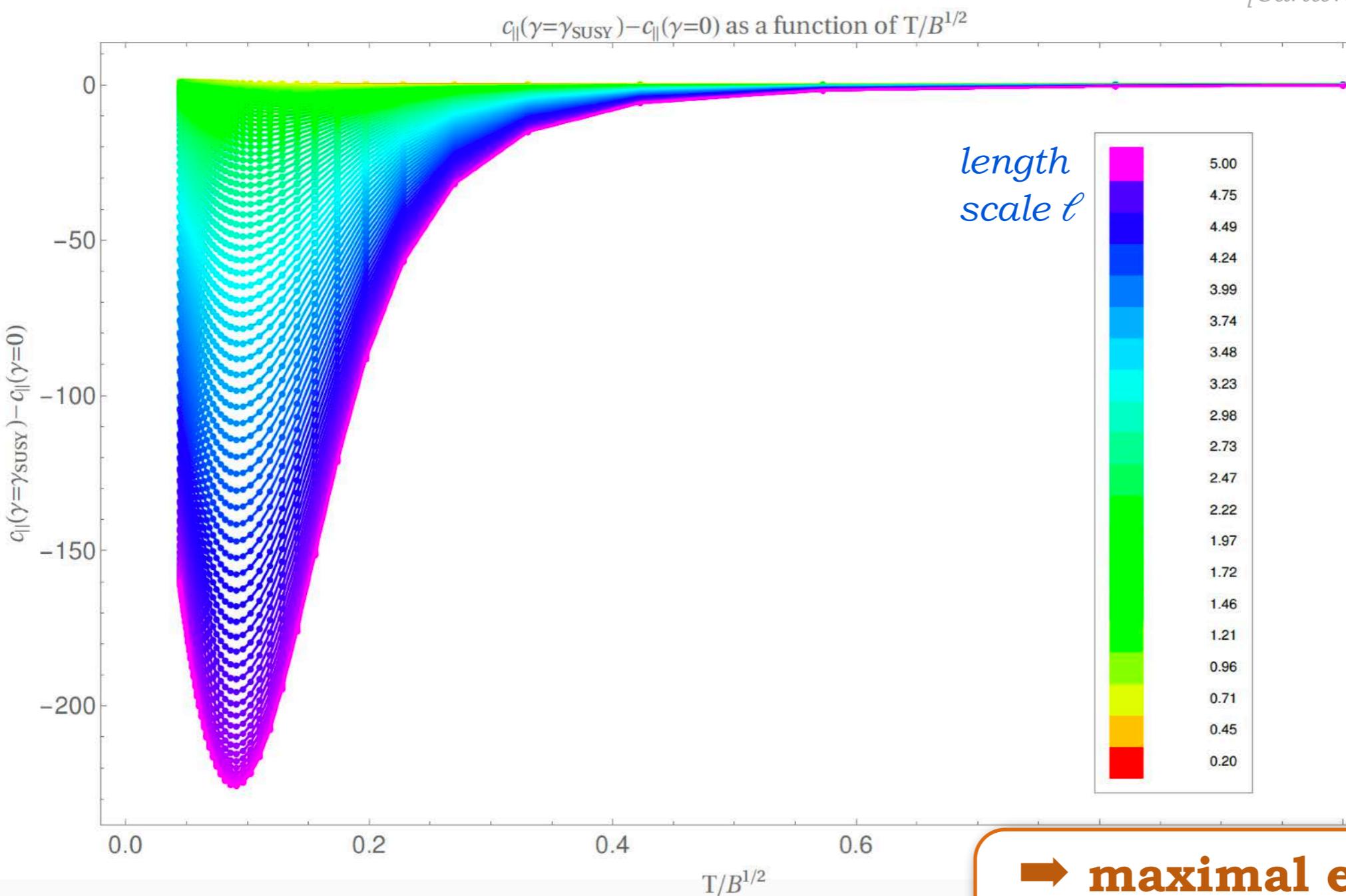


now with temperature, magnetic field, charge, chiral anomaly

- c-function increases
- effect of the charged, magnetic, thermal state
- IR limit: thermal entropy
- proposal: measure of occupation number

Effect of the chiral anomaly

[Cartwright, Kaminski; arXiv: 2107.12409]

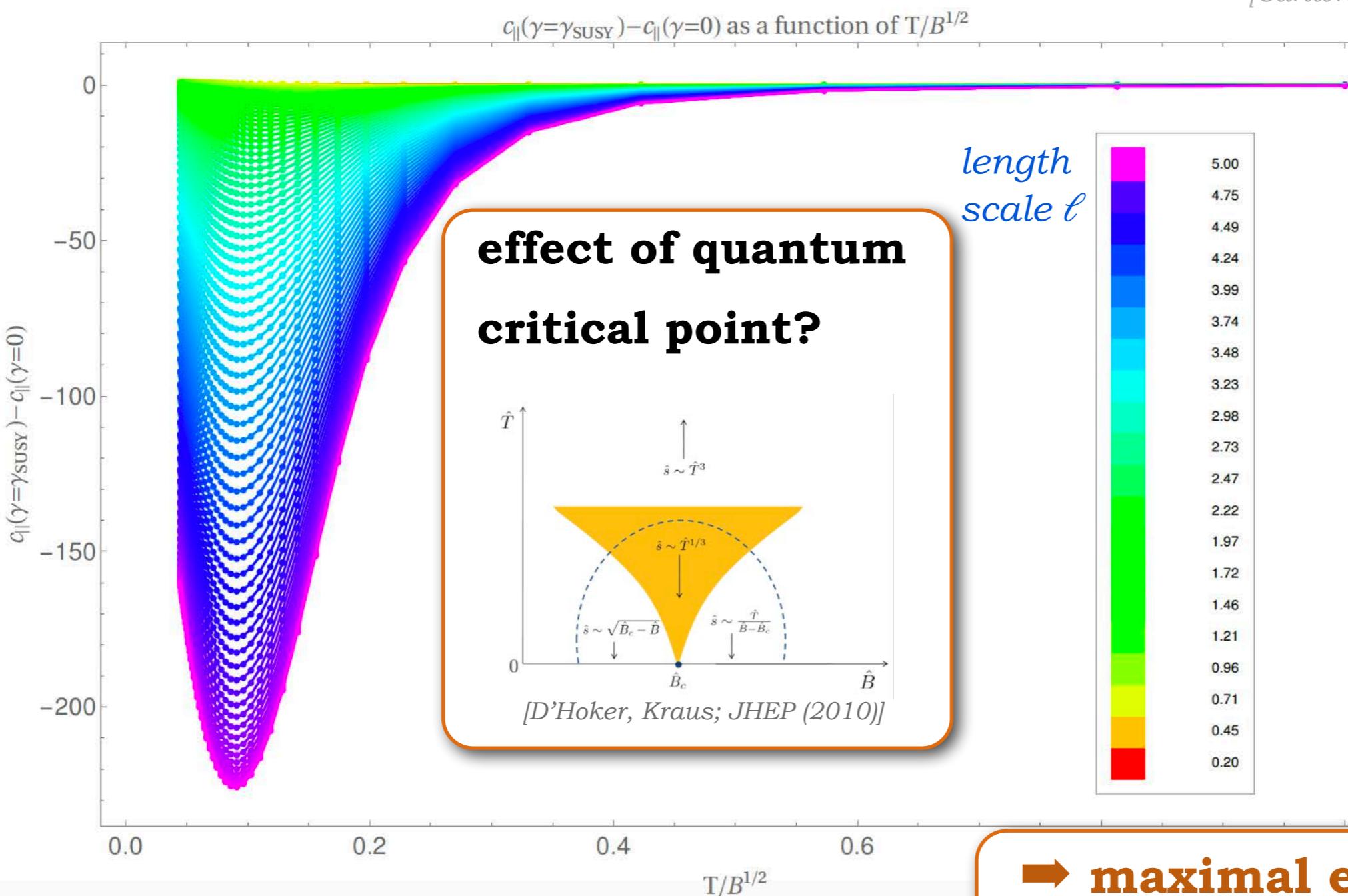


→ maximal effect at 0.1
(thermal entropy has no maximum)

now with temperature, magnetic field, charge, chiral anomaly

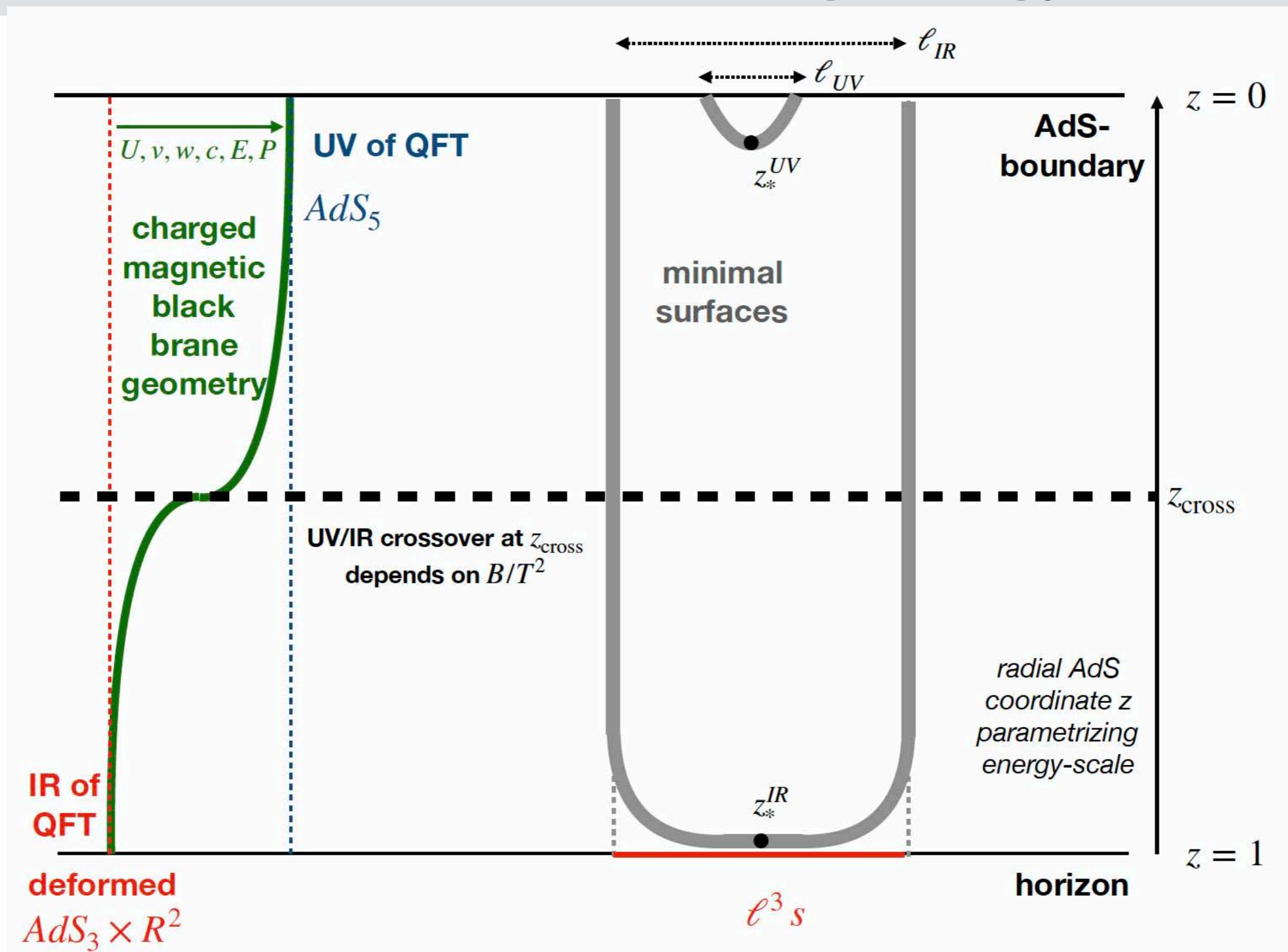
Effect of the chiral anomaly

[Cartwright, Kaminski; arXiv: 2107.12409]

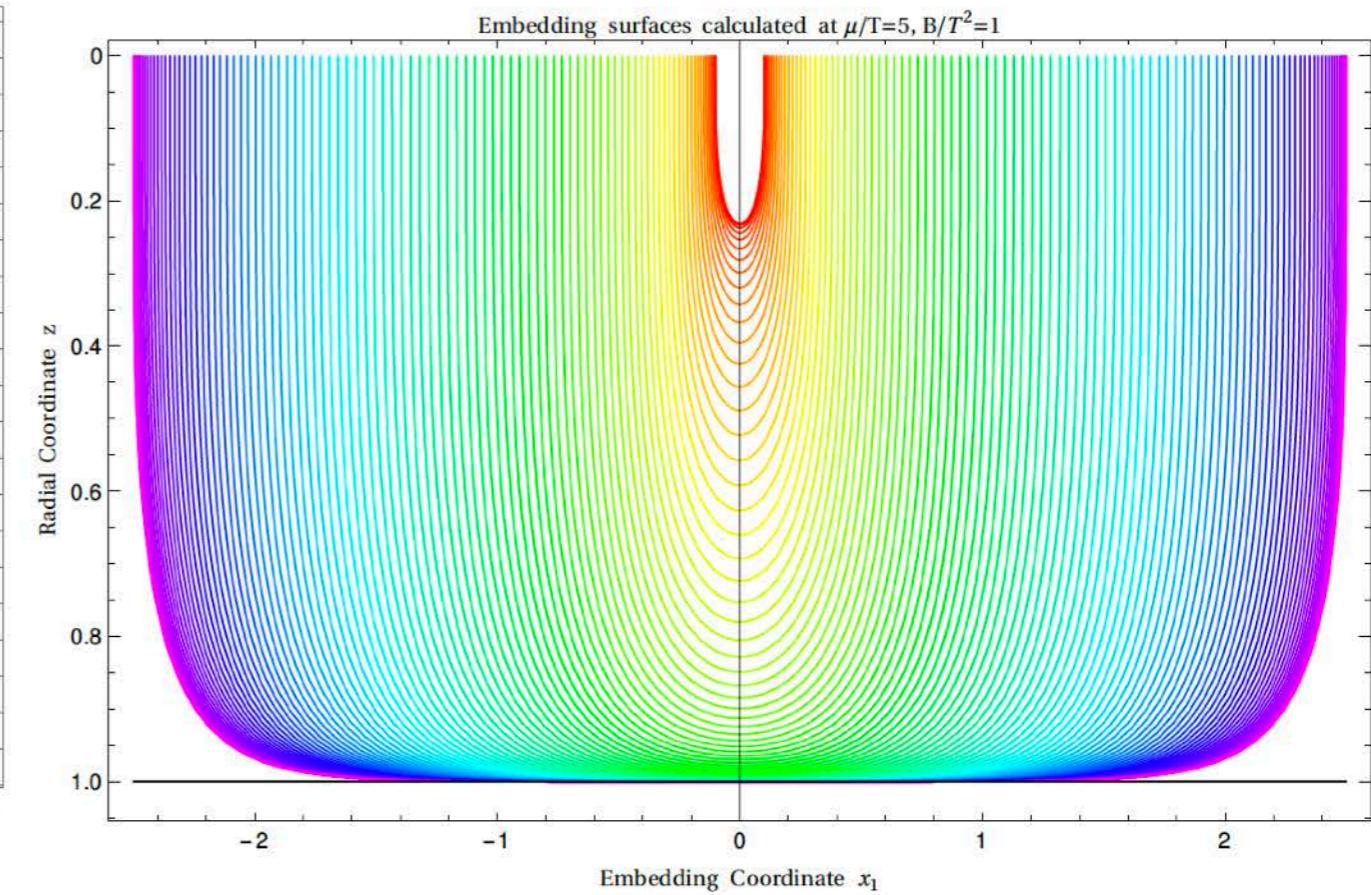
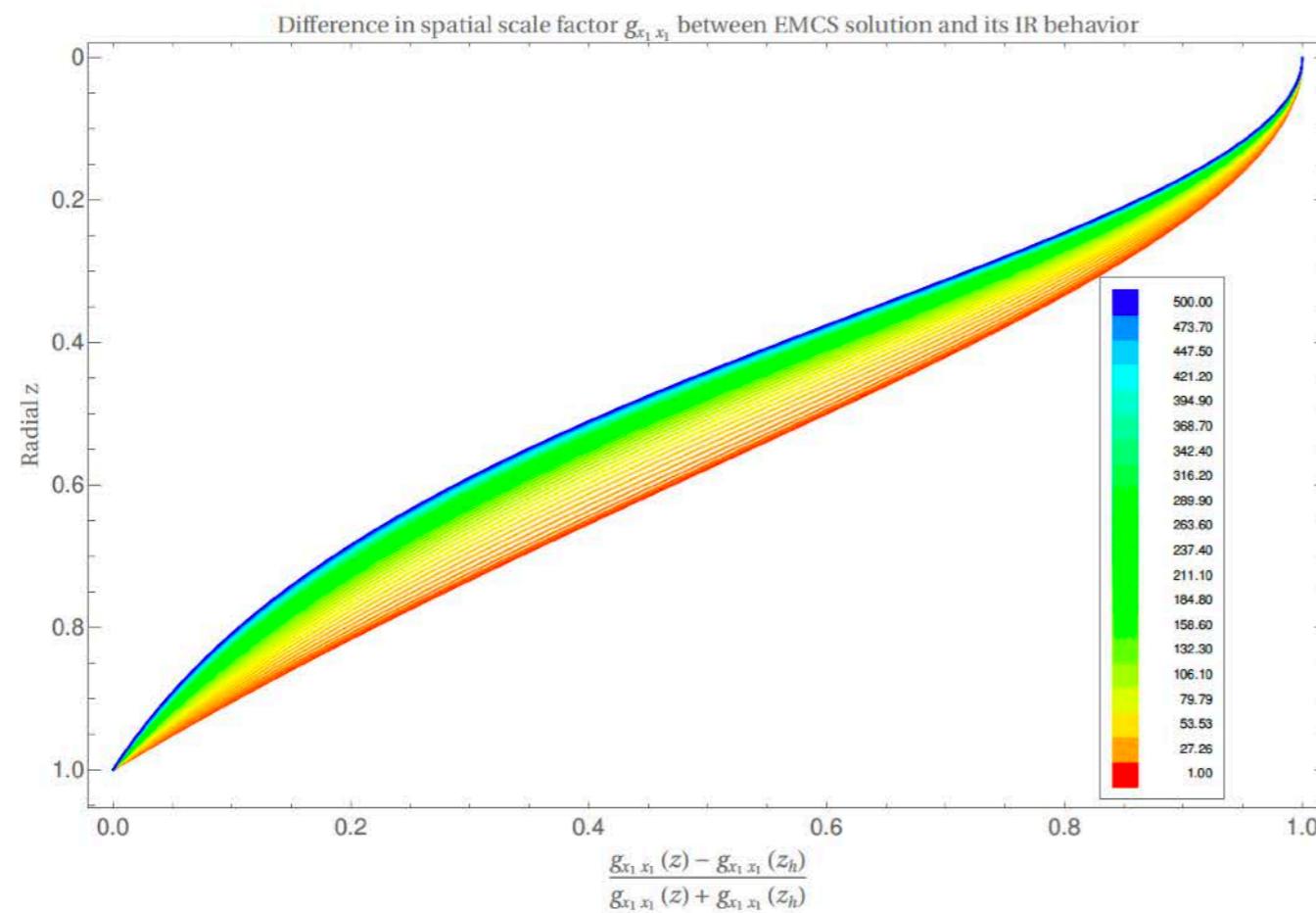
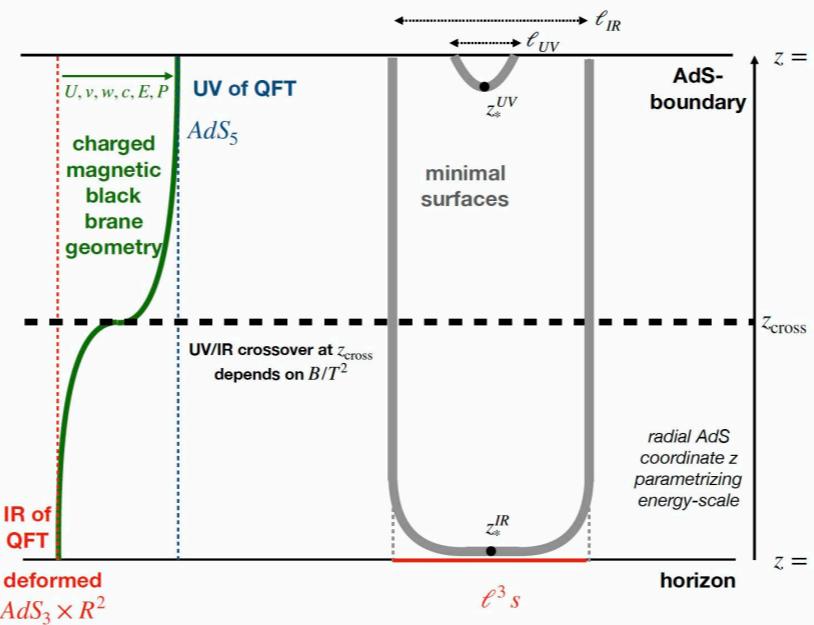


→ maximal effect at 0.1
(thermal entropy has no maximum)

Schematic picture: probing energy scales



Numerical data confirming schematic picture



Thermal entropy

Thermal entropy density in the Einstein–Maxwell and Einstein–Maxwell–Chern–Simons theory

