

Critical fluctuations in heavy-ion collisions

Marlene Nahrgang, Subatech Nantes

August 1, 2024



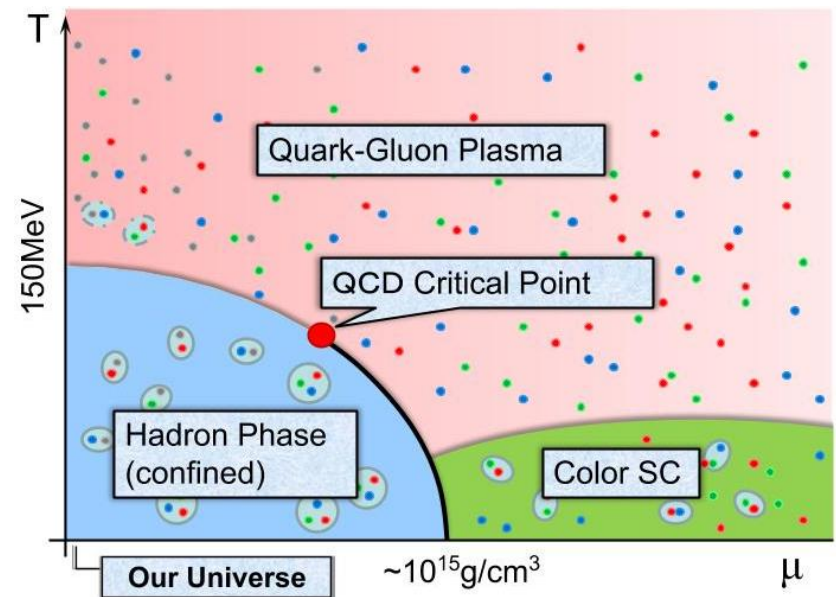
Institute for Nuclear Theory, Seattle
International Visitor Program &
“Heavy Ion Physics in the EIC Era”

Extreme QCD in heavy-ion collisions

Understanding the dynamics of the strong interaction under extreme conditions of temperature and density!

Important questions:

- Onset of deconfinement and chiral symmetry restoration?
- Properties of the strongly coupled QGP?
- Existence of a phase transition with **critical end point**?
- What are the dof in the core of compact stars?



Connect first-principle QCD calculations with experimental observables via a realistic dynamical modeling of heavy-ion collisions and astrophysical events!

How to observe the critical point in HIC

- At a **critical point**, the correlation length ξ diverges and so do the fluctuations.
- Observable in higher-order cumulants of net-baryon number.

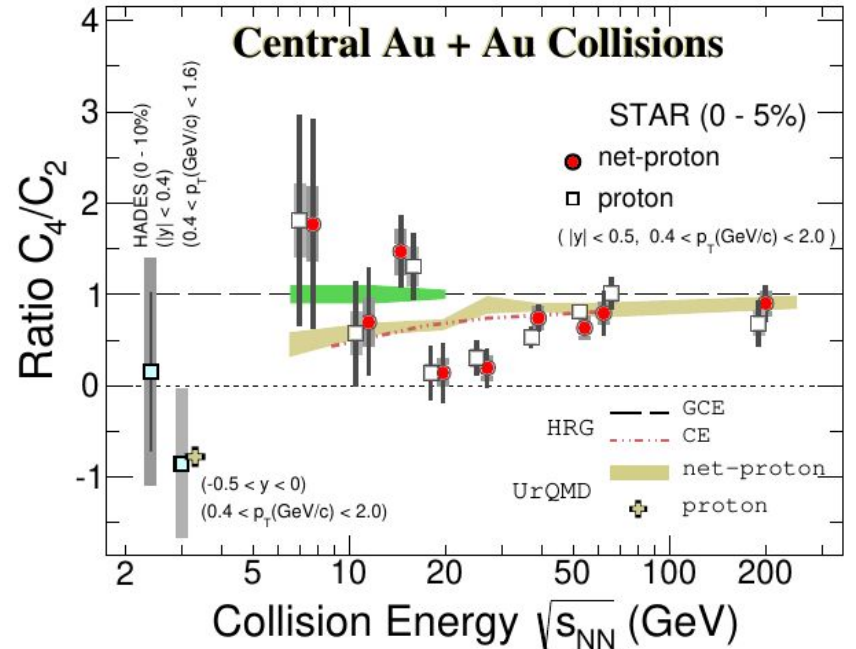
$$\chi_1 = \frac{1}{VT^3} \langle N \rangle, \quad \chi_2 = \frac{1}{VT^3} \langle (\Delta N)^2 \rangle, \quad \chi_3 = \frac{1}{VT^3} \langle (\Delta N)^3 \rangle,$$

$$\chi_4 = \frac{1}{VT^3} \langle (\Delta N)^4 \rangle_c \equiv \frac{1}{VT^3} (\langle (\Delta N)^4 \rangle - 3 \langle (\Delta N)^2 \rangle^2).$$

- To 0th order in V fluctuations:

$\frac{\chi_2}{\chi_1} = \frac{\sigma^2}{M}$	$\frac{\chi_3}{\chi_2} = S\sigma$	$\frac{\chi_4}{\chi_2} = \kappa\sigma^2$
variance	Skewness	Kurtosis

- At a **CP** the τ_{relax} diverges with ξ which leads to critical slowing down



STAR, PRL 128 (2022)

Interesting deviations from the baseline in the experimental data... are they due to the critical point of QCD?

Need a dynamical model!

Importance of dynamical modeling

In a grand-canonical ensemble the system is...

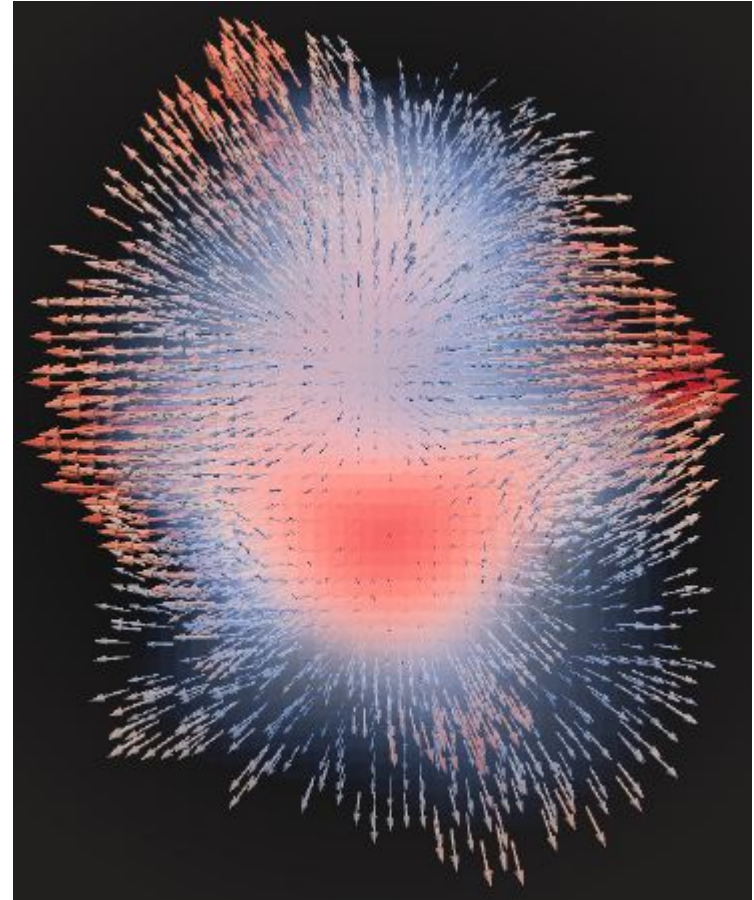
- in thermal equilibrium (= long-lived)
- in equilibrium with a particle heat bath
- spatially infinite
- and static

Systems created in a heavy-ion collision are

- short-lived
- spatially small
- inhomogeneous
- and highly dynamical!

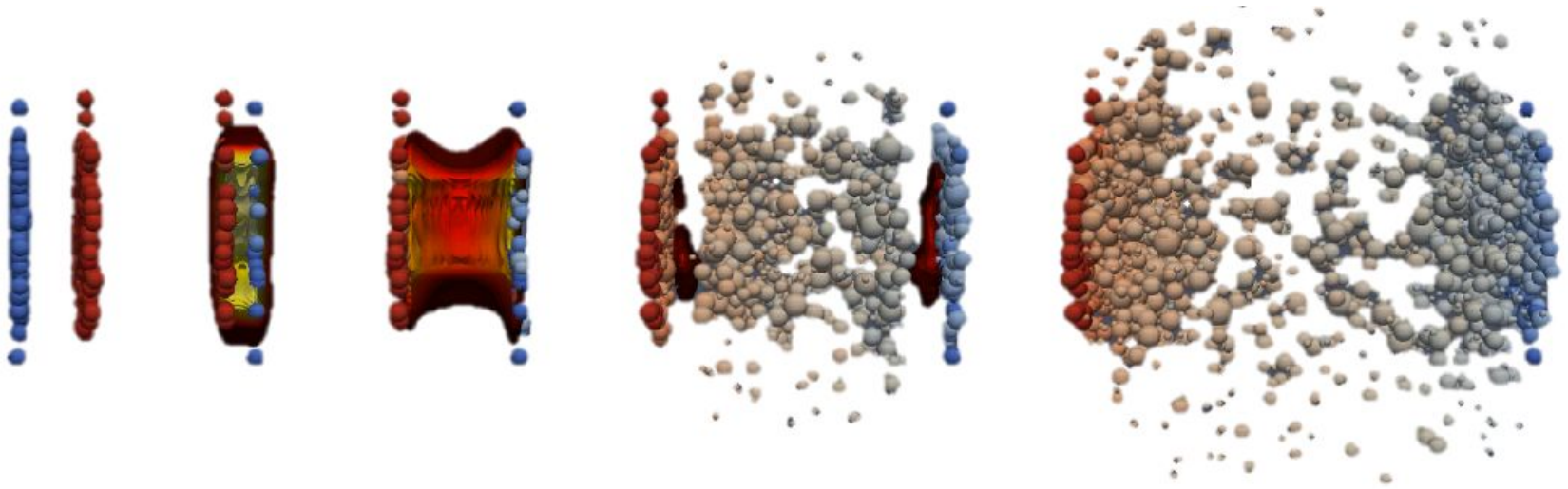
Solution: Develop dynamical models to describe the phase transition in heavy-ion collisions

Event-by-event dynamical modeling allows us in addition to study different particle species, experimental cuts, hadronic final interactions, etc.



madai.us

Fluctuations all along the way



- Initial state fluctuations due to quantum mechanical fluctuations and multiplicity fluctuations
- Thermal fluctuations, including the formation of **critical fluctuations**
- Fluctuations due to the hadronization process
- **Fate of fluctuations in the hadronic phase**
- Imperfect detection efficiency and finite acceptance

madai.us

Approaches to fluid dynamical fluctuations

There are two main approaches of describing fluid dynamics with noise:

Hydro-kinetics

- Set of deterministic kinetic equations for n-point functions of fluid dynamical fields
- Renormalization (perturbatively) performed during the derivation
- Statistical average performed in the derivation of deterministic equations

Stochastic/fluctuating fluid dynamics

- Numerical implementation of the fluid dynamical equations with stochastic conservation law:

$$\partial_\mu T^{\mu\nu} = 0, \quad T^{\mu\nu} = T_{\text{ideal}}^{\mu\nu} + T_{\text{viscous}}^{\mu\nu} + S_{\text{noise}}^{\mu\nu},$$
$$\partial_\mu J^\mu = 0, \quad J^\mu = J_{\text{ideal}}^\mu + J_{\text{viscous}}^\mu + I_{\text{noise}}^\mu.$$

- Sample discretized noise event-by-event
- Observables are calculated from statistical averaging over events.
- Can easily be integrated in standard event generators of HIC!
- Many challenges....

A.Andreev, Sov. Phys. JETP 32 no. 5 (1971) and 48 no. 3 (1978);
Y. Akamatsu et al., PRC 95 no. 1 (2017) and 97 no. 2 (2018); M.
Stephanov et al., PRD 98 (2018); M. Martinez et al., PRC 99 no. 5
(2019); X. An et al., PRC 100 no. 2 (2019), PRL 127 (2021); L. Du et
al., PRC 102 (2020); K. Rajagopal, NPA 1005 (2021)

Fluctuating Dissipative Fluid Dynamics

The correlators of the thermal noise terms in the energy momentum tensor and the conserved currents :

$$\langle S^{\mu\nu}(x_1)S^{\alpha\beta}(x_2)\rangle = 2T \left[\begin{array}{l} \eta (\Delta^{\mu\alpha} \Delta^{\nu\beta} + \Delta^{\mu\beta} \Delta^{\nu\alpha}) \\ + \left(\zeta - \frac{2}{3}\eta \right) \Delta^{\mu\nu} \Delta^{\alpha\beta} \end{array} \right] \delta^{(4)}(x_1 - x_2).$$

$$\langle I^\mu(x_1)I^\nu(x_2)\rangle = 2T\sigma\Delta^{\mu\nu}\delta^{(4)}(x_1 - x_2).$$

Several issues arise from the discretization of the Dirac delta function in the noise

- Stochastic noise introduces a lattice spacing dependence.
- Correction terms due to renormalization become large for small lattice spacings.
- Large noise contributions can locally lead to negative energy densities.
- Large gradients introduced by the uncorrelated noise is a problem for PDE solvers.

Fluctuating Dissipative Fluid Dynamics

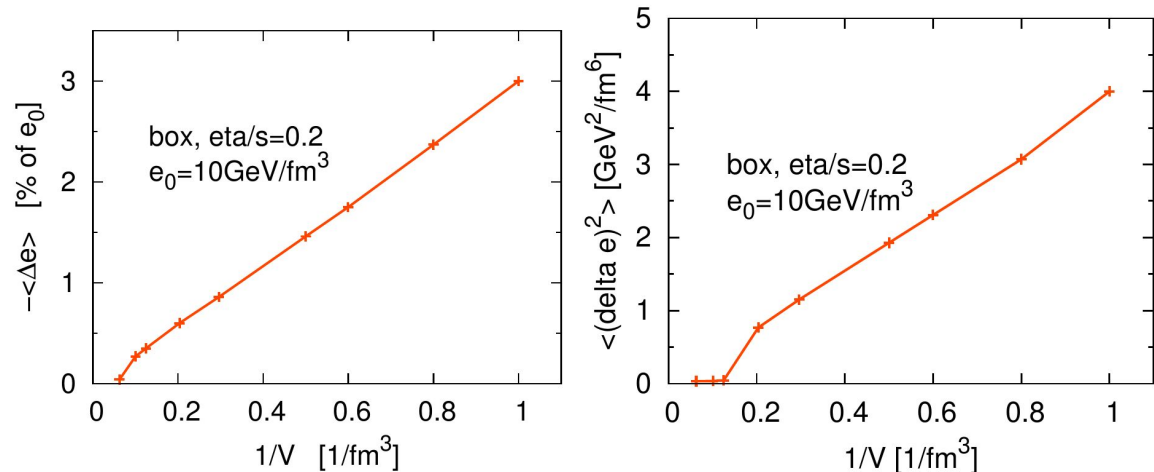
First implementations of FDFD have shown: need to limit the resolution scale; simulate noise down to a particular filter length scale, for which:

$$l_{\text{grid}} < l_{\text{filter}} \lesssim l_{\text{noise}} \ll l_{\text{hydro}}$$

Murase et al.: noise is smeared by Gaussians with widths of 1-1.5 fm (choice not discussed), large enhancement of flow observed. [K. Murase et al, NPA 956 \(2016\);](#)

Nahrgang et al.: noise is coarse-grained over distances of approx. 1fm, lattice spacing dependence of the energy density and its fluctuations observed.

[M. Nahrgang et al, Acta Phys. Polon. 10 \(2017\);](#)



Singh et al.: high-mode Fourier filter with a coarse graining scale of $> 1\text{fm}$, multiplicities and flow are little affected by the inclusion of fluctuations

[M. Singh et al, NPA 982 \(2019\);](#)

Renormalizing critical dynamics in model A of Hohenberg-Halperin in 3 dimensions

Cassol-Seewald et al. 0711.1866
Farakos et al. 9412091, 9404201
Gleiser, Ramos 9311278

Stochastic relaxation equation of chiral order parameter

$$\frac{\partial^2 \varphi}{\partial t^2} - \nabla^2 \varphi + \eta \frac{\partial \varphi}{\partial t} + \frac{\partial V_{\text{eff}}}{\partial \varphi} = \xi$$

Ginzburg-Landau effective potential, ϵ encodes phase transition

$$V_{\text{eff}}(\varphi) = \frac{1}{2}\epsilon\varphi^2 + \frac{1}{4}\lambda\varphi^4$$

White thermal noise

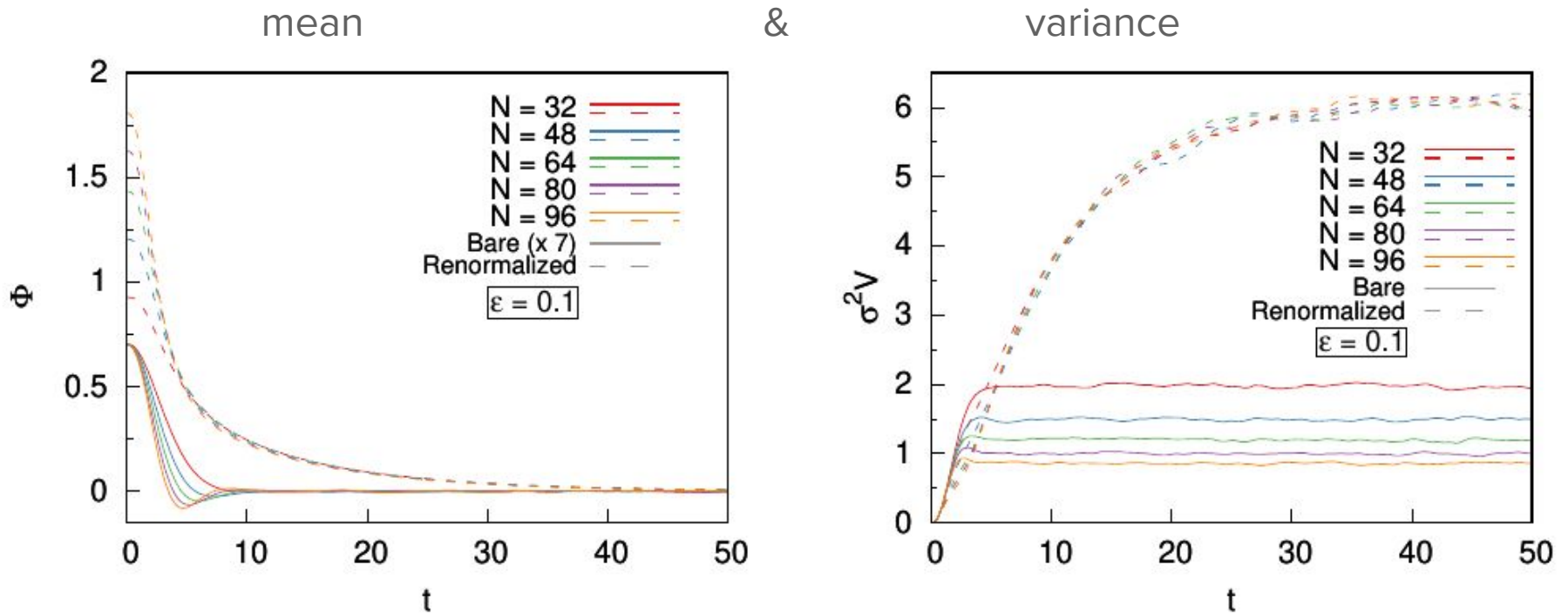
$$\langle \xi(\vec{x}, t) \rangle = 0 \quad \text{and} \quad \langle \xi(\vec{x}, t) \xi(\vec{x}', t') \rangle = 2\eta T \delta(\vec{x} - \vec{x}') \delta(t - t')$$

Counterterm for mass renormalization

$$V_{\text{CT}} = \left\{ -\frac{3\lambda\Sigma T}{4\pi dx} + \frac{3}{8} \left(\frac{\lambda T}{\pi} \right)^2 \left[\ln \left(\frac{6}{M dx} \right) + \zeta \right] \right\} \frac{\varphi^2}{2}$$

Renormalizing critical dynamics in model A

Time evolution near a critical point of

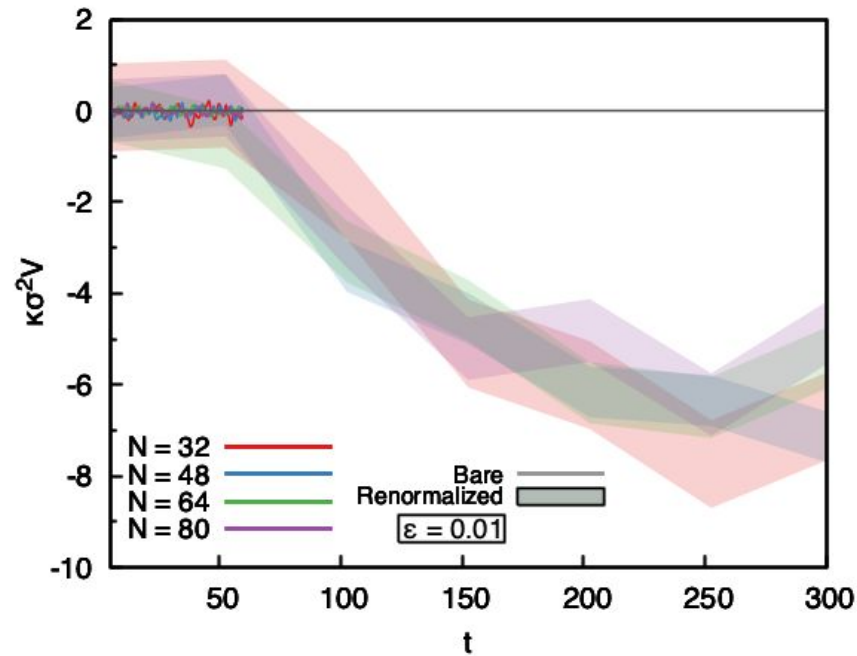


Restoration of lattice spacing independence after including the counterterm

Renormalizing critical dynamics in model A

Time evolution near a critical point of

kurtosis



A nonzero kurtosis is observed after inclusion of the counterterm. Lattice spacing dependence or independence cannot be resolved within the available statistics.

Net-baryon diffusion: model B in 1+1 dimensions

- In the long-time, equilibrium limit the net-baryon density is the slowest mode near the CP.
- For baryonic matter that decouples from the energy flow of the system (model B of Hohenberg, Halperin), the diffusive dynamics follows the minimization of the free energy F

$$\partial_t n_B(t, \mathbf{x}) = \kappa \nabla^2 \left(\frac{\delta \mathcal{F}[n_B]}{\delta n_B} \right) + \nabla \mathbf{J}(t, \mathbf{x})$$

with the stochastic current
(Gaussian, white noise)

$$\mathbf{J}(t, \mathbf{x}) = \sqrt{2T\kappa} \zeta(t, \mathbf{x}), \quad \kappa = \frac{Dn_c}{T}$$

MN and M. Bluhm, PRD 99 (2019) and PRD 102 (2020)

- Apply the evolution in a 1+1 dimensional, boost-invariant Bjorken expansion

Fluctuations in an expanding background, e.g. J. Kapusta et al, PRC 85 (2012); Y. Akamatsu et al. PRC 95 (2017), M. Martinez et al, PRC 99 (2019)

- The nonlinear stochastic diffusion equation transforms as:

$$\partial_\tau n_B = \frac{Dn_c}{\tau \chi_2(\tau)} \partial_y^2 n_B - \frac{Dn_c K(\tau)}{\tau} \partial_y^4 n_B + \frac{Dn_c}{6 \tau \chi_4(\tau)} \partial_y^2 n_B^3 - \partial_y \xi.$$

In Gauss limit: M. Sakaida et al PRC 95 (2017); nonlinear (only critical): M. Kitazawa, G.Pihan, N. Touroux, M. Bluhm, MN NPA 1005 (2021)

Singular and regular susceptibilities

- Parametrize the susceptibilities $\chi_2(\tau)$ and $\chi_4(\tau)$ with a regular part using the argument in

M. Asakawa, U. Heinz, B. Müller, PRL 85 (2000)

$$\chi_n(\tau) = \frac{\langle \Delta N_B^n \rangle}{S} \Big|_{\text{QGP/HRG}} = \frac{\chi_B^n}{s/T^3} \Big|_{\text{QGP/HRG}}$$

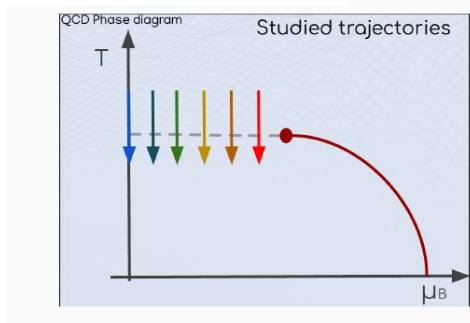
With χ_B^n and the entropy fixed to lattice results at $T=280$ MeV for the QGP and $T=130$ MeV for the HRG, matched via tanh function.

- Couple with the singular contribution (3D Ising) via

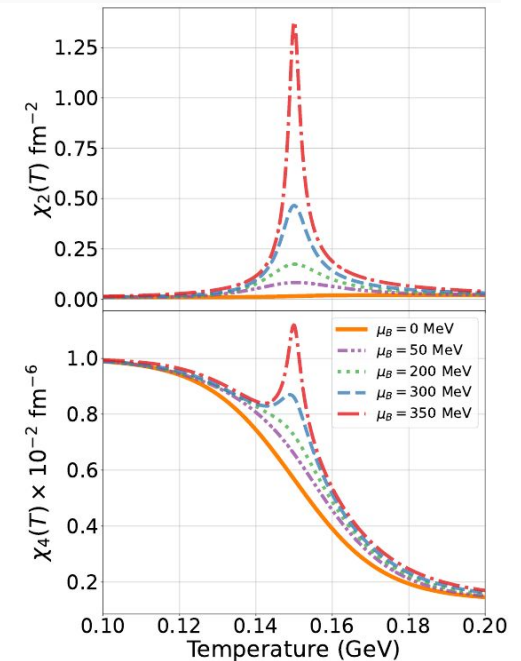
$$\chi_n(T) = \chi_n^{\text{sing}}(T) + \chi_n^{\text{reg}}(T)$$

- Match to the coefficients in the expansion of the free energy density functional.

$$\chi_n(\tau) = \tau \left(\frac{\delta^n \mathcal{F}}{\delta n_B^n} \Big|_{\Delta n_B=0} \right)^{-1}$$



- Investigate several trajectories in the QCD phase diagram.



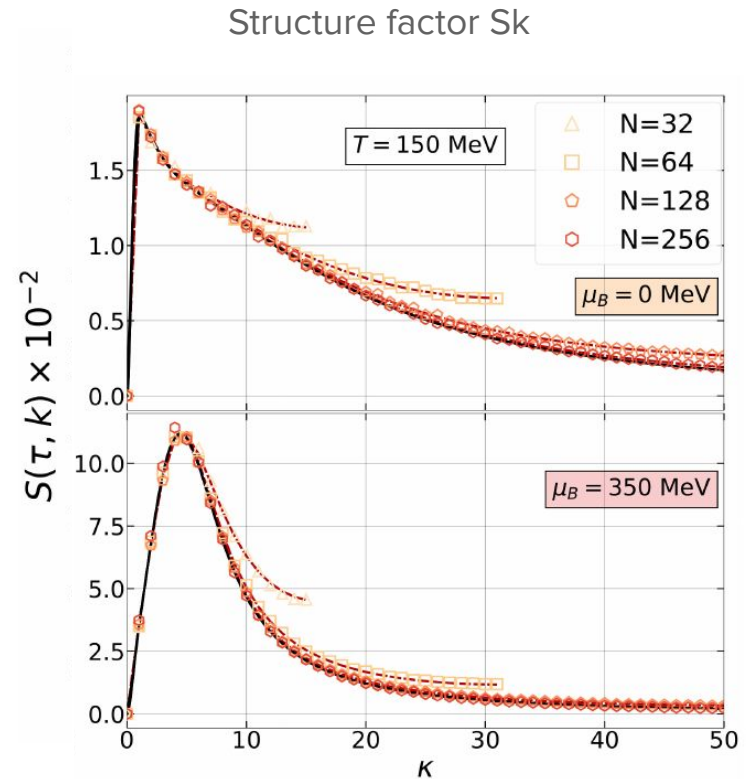
Validating the linear model

only here
 $\lambda_4 = 0!$

Important step for all fluctuating codes:

validation of the appropriate linear model:

- Structure factor and equal-time correlation function are well reproduced
- Approach to continuum as resolution is increased.
- Lower wavenumbers well described with the maximal resolution chosen for this work.
- Enhancement of fluctuations with low wavenumbers at $T_c = 150$ MeV.
- Discretization and baryon conservation effects under control.



G. Pihan, M. Bluhm, M. Kitazawa, T. Sami, MN, PRC 107 (2023)

Anticorrelations as a signal for the critical point

$$\frac{\chi_2}{\chi_1} = \frac{\sigma^2}{M}$$

variance

$$\frac{\chi_3}{\chi_2} = S\sigma$$

Skewness

$$\frac{\chi_4}{\chi_2} = \kappa\sigma^2$$

Kurtosis

Higher-order moments are more sensitive

- to the divergence of the correlation length
- and to any other noncritical aspect of HIC...

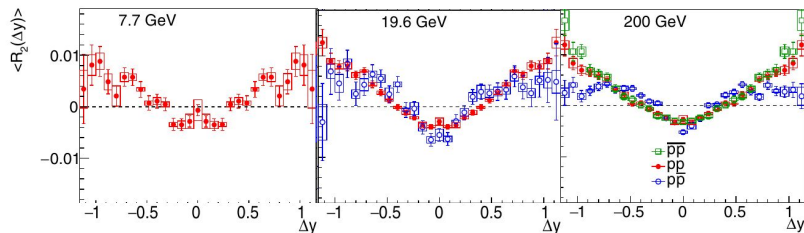
Here: dynamical fluctuations of net-baryon density

- Large fluctuations are balanced by large anti-correlations (net-baryon conservation)
- Due to the dynamics these anti-correlations cannot diffuse fast enough
- Approaching T_c they are visible at $y \sim 1-2$
- At lower T the minimum becomes smaller and moves to larger y



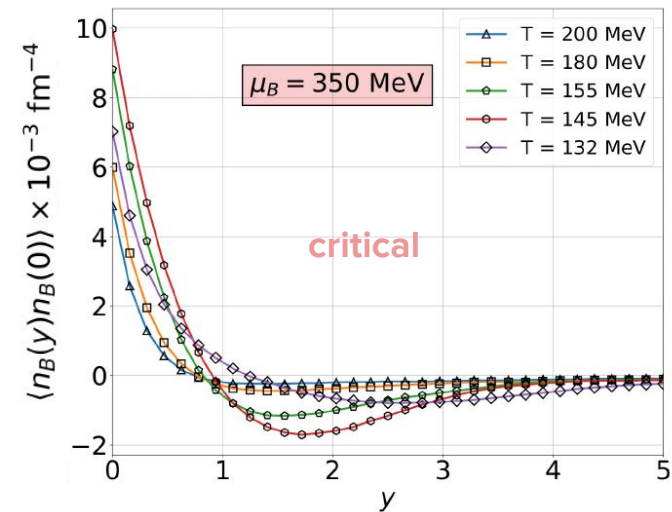
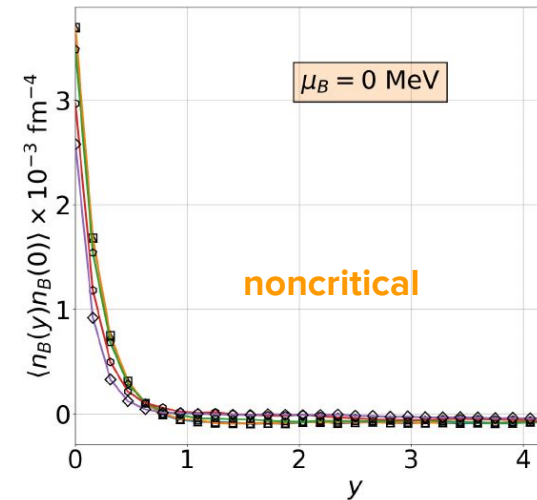
Possible detection depends crucially on T_{FO}

- Interesting experimental data:
(STAR AuAu, 30-40% most central, $0.4 < p_T < 2 \text{ GeV}$)



STAR Collab, PRC 101 (2021)

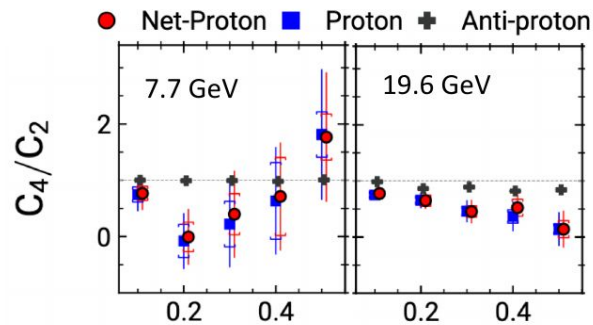
correlation function in rapidity



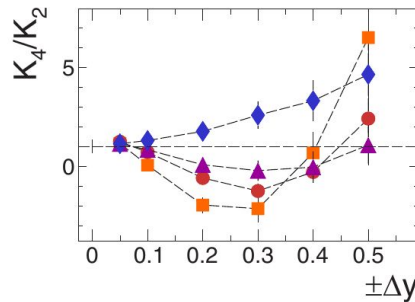
Word of caution: not yet an apple-to-apple comparison possible

Non-monotonic kurtosis as a signal for the critical point

- Monotonic increase in the variance
- Non-monotonic Kurtosis only for the trajectories with **critical point**.
- This **non-monotonic behavior** of the kurtosis survives the rapid expansion for a diffusion length $D = 1$ fm
- strong indication for the presence of the **critical point**.
- For increasing D the minimum moves to larger distances in rapidity.
- Essential for the experiment to cover a wide range in rapidity to see the non-monotonicity.
- Interesting experimental data:
(STAR AuAu, 0-5% most central, $0.4 < p_T < 2$ GeV, $|y| < y_{max}$)



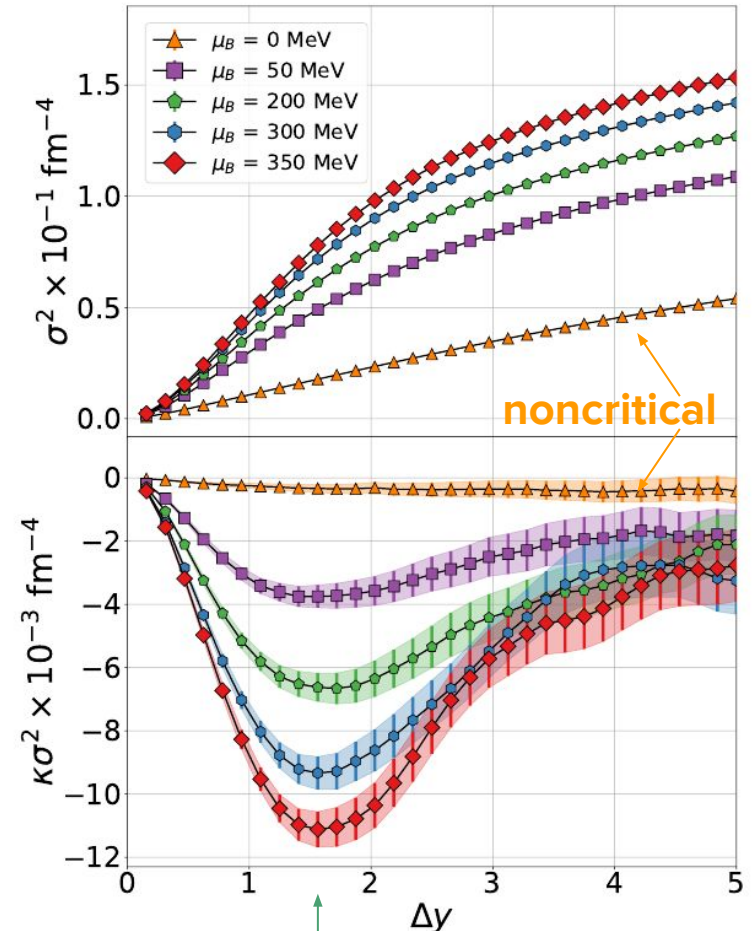
STAR Collab, PRC 104 (2021)



HADES Collab, PRC 102 (2020)

Word of caution: not yet an apple-to-apple comparison possible

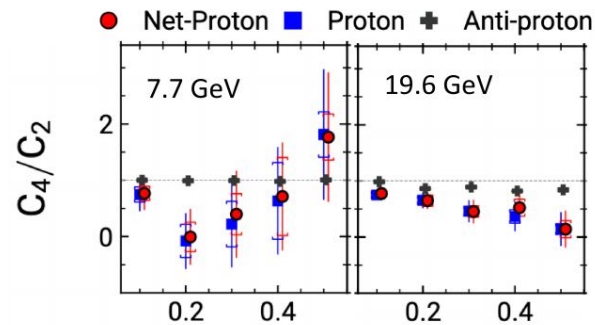
rapidity dependence of fluctuation observables



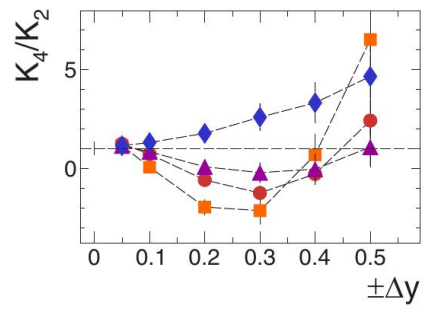
STAR iTPC

Non-monotonic kurtosis as a signal for the critical point

- Monotonic increase in the variance
- Non-monotonic Kurtosis only for the trajectories with **critical point**.
- This **non-monotonic behavior** of the kurtosis survives the rapid expansion for a diffusion length $D = 1$ fm
- strong indication for the presence of the **critical point**.
- For increasing D the minimum moves to larger distances in rapidity.
- Essential for the experiment to cover a wide range in rapidity to see the non-monotonicity.
- Interesting experimental data:
(STAR AuAu, 0-5% most central, $0.4 < p_T < 2 \text{ GeV}$, $|y| < y_{\text{max}}$)



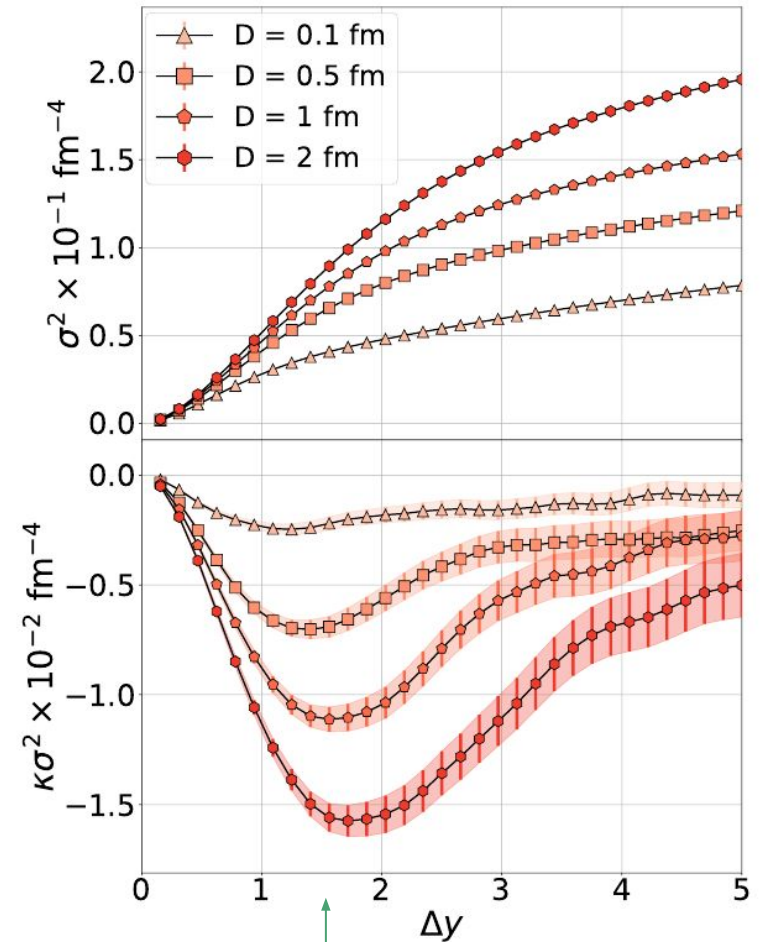
STAR Collab, PRC 104 (2021)



HADES Collab, PRC 102 (2020)

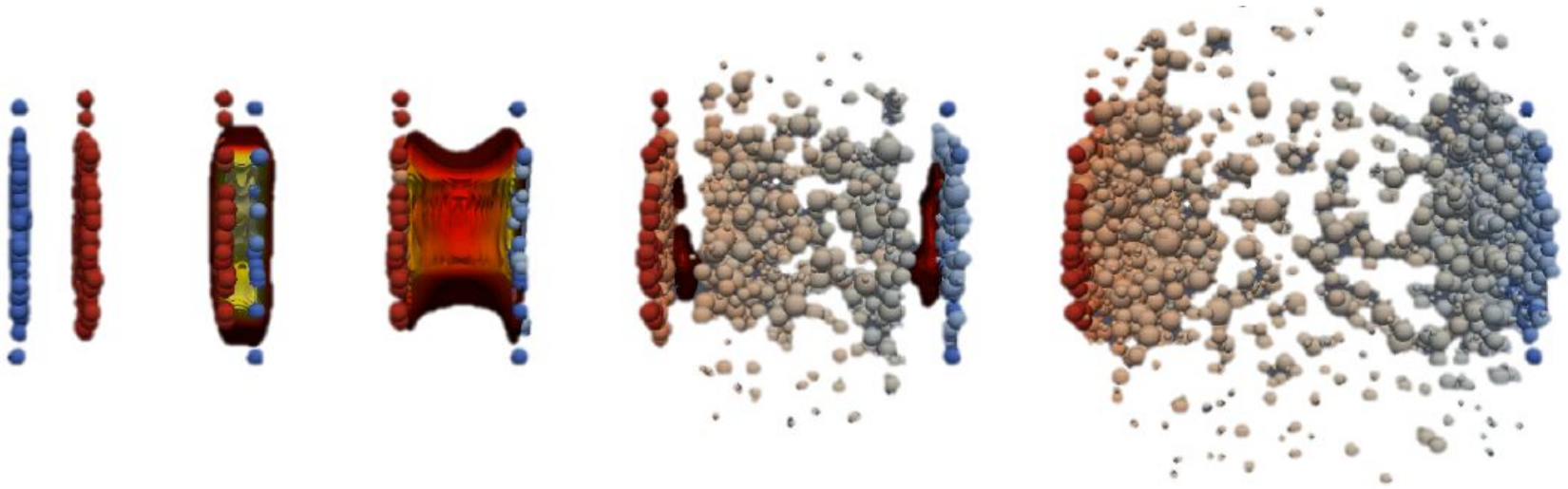
Word of caution: not yet an apple-to-apple comparison possible

rapidity dependence of fluctuation observables



STAR iTPC

Fluctuations all along the way

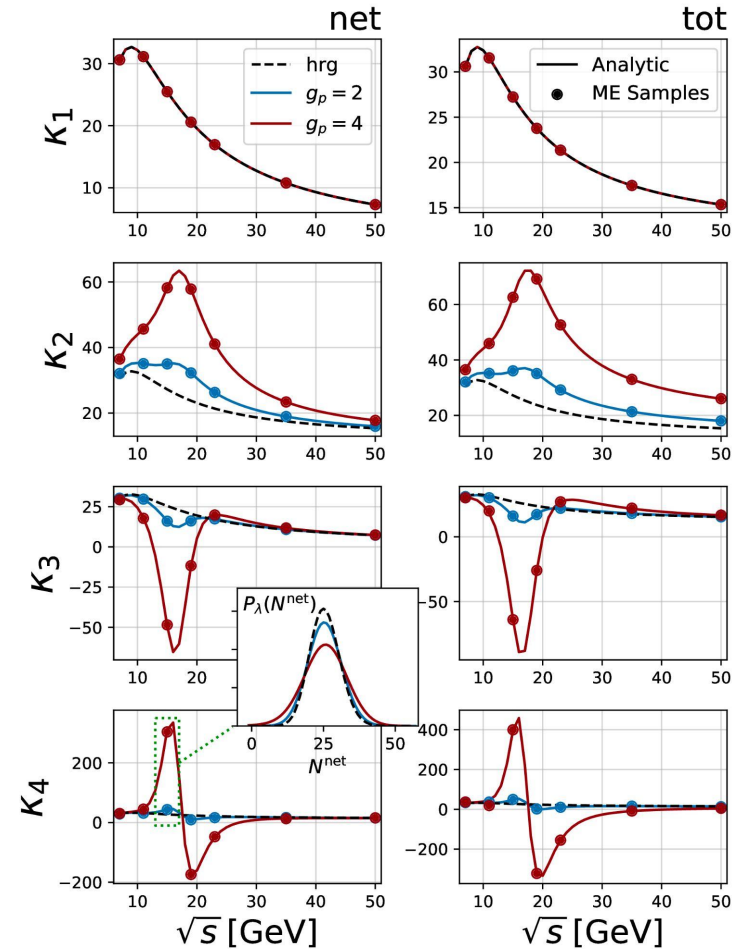
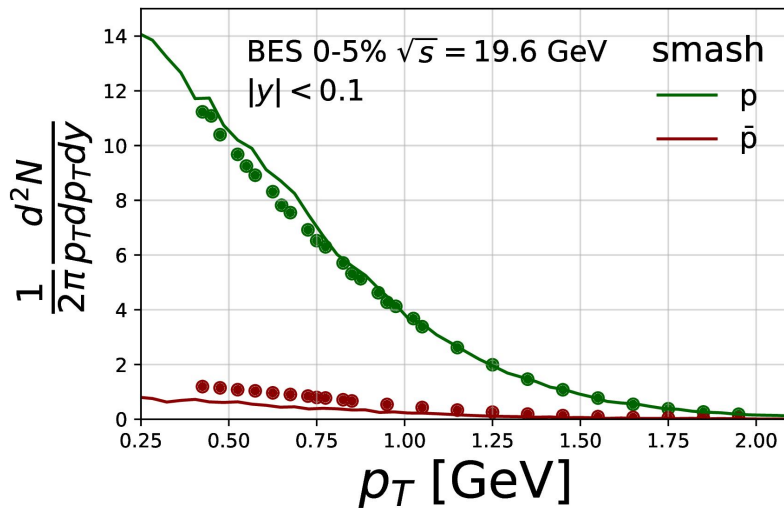


- Initial state fluctuations due to quantum mechanical fluctuations and multiplicity fluctuations
- Thermal fluctuations, including the formation of **critical fluctuations**
- Fluctuations due to the hadronization process
- **Fate of fluctuations in the hadronic phase**
- Imperfect detection efficiency and finite acceptance

madai.us

Fate of critical fluctuations in the hadronic phase

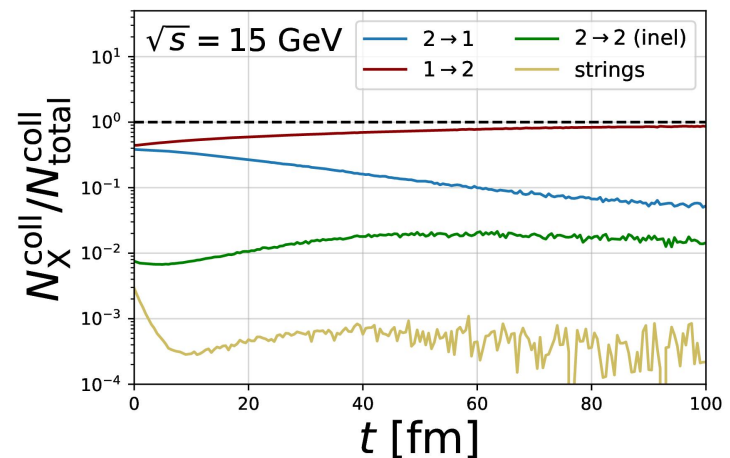
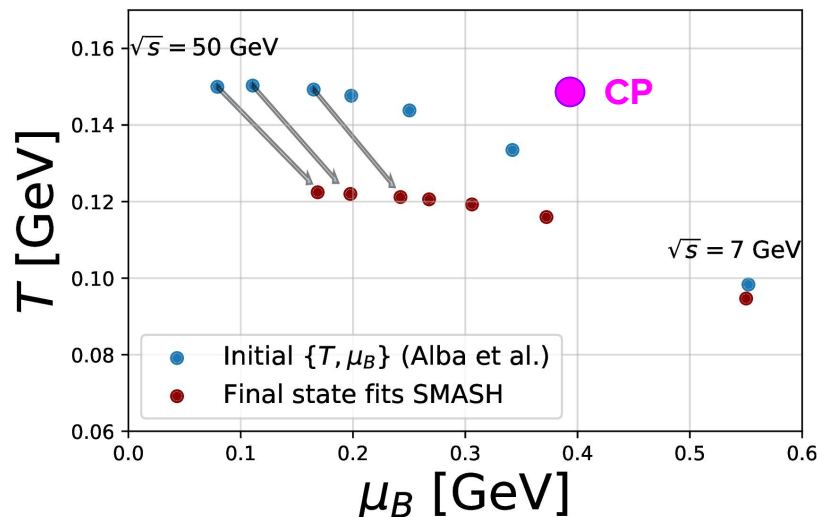
- Calculate up to 4th order cumulants of critical fluctuations from an 3d Ising model mapping to QCD and couple it to HRG cumulants ($g_c=2,4$).
- Reconstruct the particle distributions from the cumulants + maximum entropy constraint
- Assume simple geometry at particlization: uniform spatial distribution in a sphere ($R=9\text{fm}$)
Momentum distribution $f_{i,k} = e^{-u \cdot k_i/T}$
velocity $\vec{u}(r) = \vec{e}_r u_0 r/R$ with $u_0=0.5$



Fate of critical fluctuations in the hadronic phase

- Apply smash (<https://smash-transport.github.io>) to the final hadronic interactions of the initialized particles.
- Resonance decay and regeneration are the dominant processes during the hadronic expansion.

Particle	Mass [GeV/c ²]	Degeneracy
π	0.138	3
ρ	0.776	6
K	0.494	4
$K^*(892)$	0.892	8
N	0.938	8
Δ	1.232	32
Λ	1.116	2
Σ	1.189	12



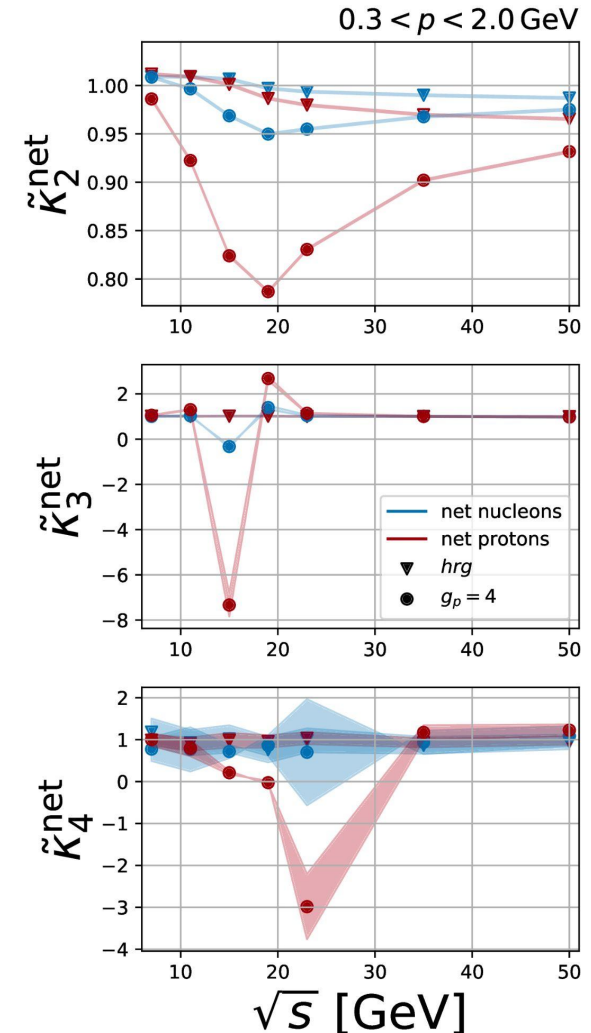
Impact of resonance dynamics vs decay

- Compare the dynamical effect of the resonance decay and regeneration compared to only resonance decay:

$$\tilde{\kappa}_n = \frac{\kappa_n^{\text{dynamical}}}{\kappa_n^{\text{decays}}}$$

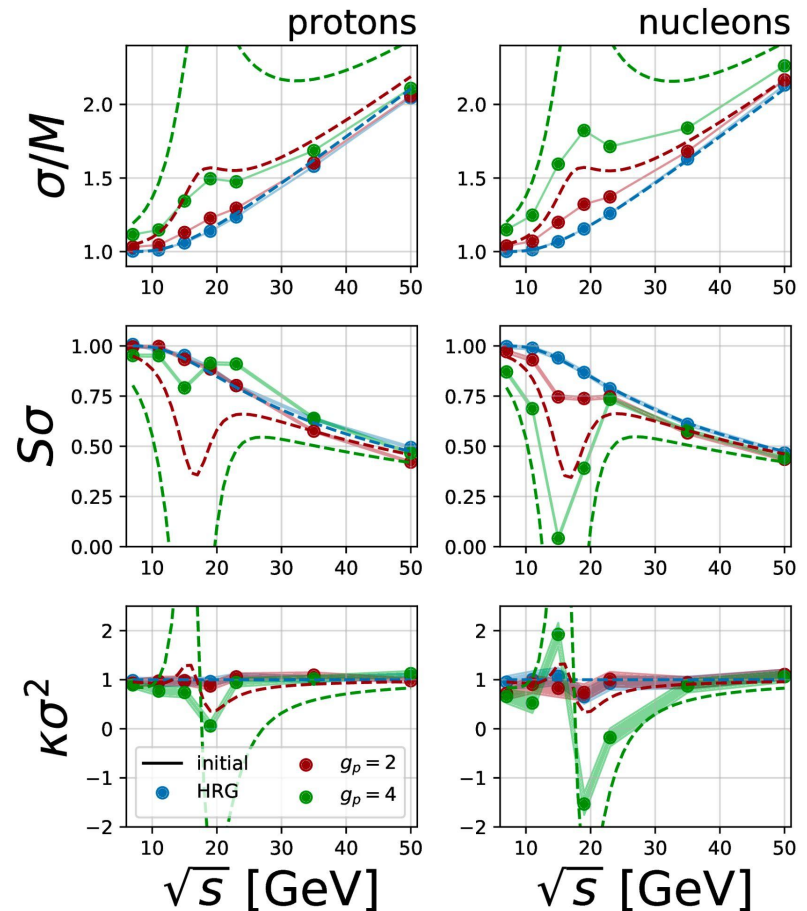
- Net-proton cumulants are strongly impacted by the hadron dynamics compared to the net-nucleons -> importance of isospin randomization processes.

M. Bluhm, MN, S. Bass, T. Schaefer, Eur.Phys.J.C 77 (2017);
 MN, M. Bluhm, P. Alba, R. Bellwied, C. Ratti Eur.Phys.J.C 75
 (2015); M. Kitazawa, M. Asakawa, Phys.Rev.C 85 (2012)



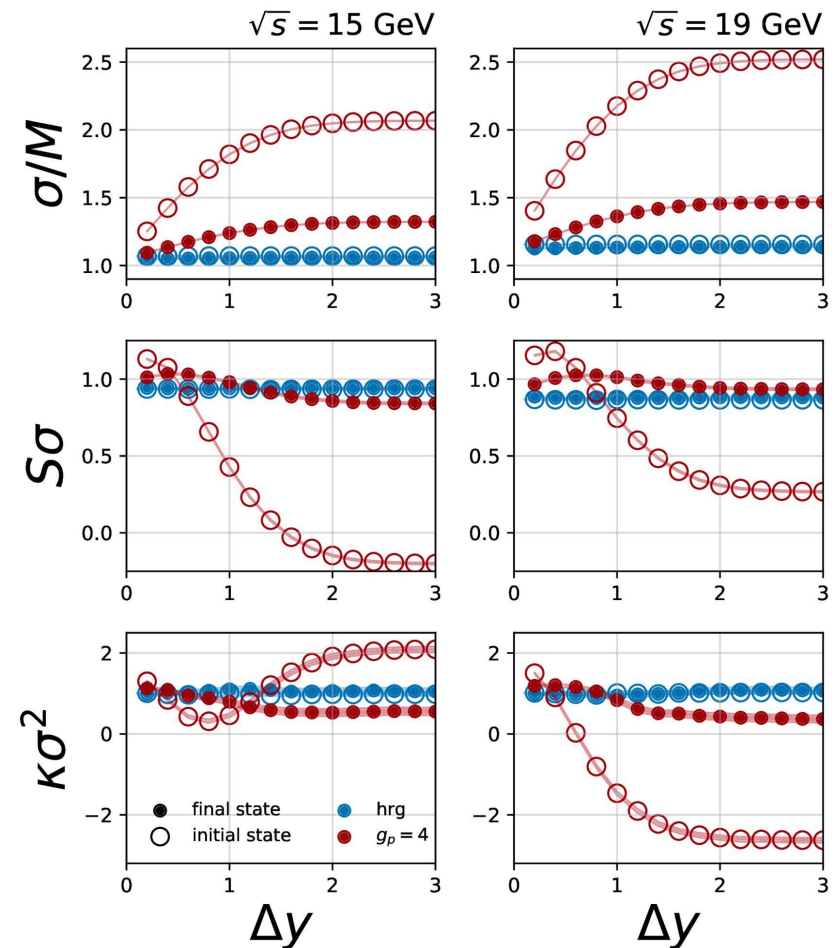
Final proton and nucleon cumulant ratios

- For $g_c = 2$, no critical signal is seen in the net-proton variance and skewness, a very small signal in the kurtosis survives.
- For $g_c = 4$, the net-proton variance shows critical features \rightarrow not compatible with experiment.
- The nucleon critical signal is significantly more pronounced than for protons only.
- Signal depends strongly on the rapidity acceptance and can even change sign in the kurtosis.



Final proton and nucleon cumulant ratios

- For $g_c = 2$, no critical signal is seen in the net-proton variance and skewness, a very small signal in the kurtosis survives.
- For $g_c = 4$, the net-proton variance shows critical features \rightarrow not compatible with experiment.
- The nucleon critical signal is significantly more pronounced than for protons only.
- Signal depends strongly on the rapidity acceptance and can even change sign in the kurtosis.



Workshop at Subatech, Nantes, October 28-31, 2024

Hydrodynamics and related observables in heavy-ion collisions

<https://indico.in2p3.fr/event/32630/>

Broad topics:

- hydrodynamic fluctuations
- finite baryon density
- spin-hydrodynamics
- anisotropic hydrodynamics
- stable first-order hydrodynamics
- jets and heavy-flavor dynamics
- rapid hydrodynamization
- numerical developments



Conclusions

Treating the dynamics of fluctuations near the **Critical Point** is important for quantitative statements about its existence based on heavy-ion collision data!

- Net-baryon fluctuations are strongly impacted by the expansion dynamics.
- **Anticorrelations of baryons can signal the CP.**
- **Non-monotonic dependence of the kurtosis on the rapidity window near the CP.**
- Resonance decay and regeneration strongly affects the critical fluctuations.

Currently explored:

- Renormalizing (chiral) fluid dynamics [with N. Attieh](#)
- Implementing (renormalized) fluctuating fluid Dynamics [with J. Sterba, I. Karpenko, B. Tomasik](#)

APPENDIX

Importance of fluctuations for transport coefficients

$$\eta \sim \int d^3x dt \langle T^{ij}(\mathbf{x}, t) T^{ij}(0, 0) \rangle$$

Included in fluid dynamics

NOT included in fluid dynamics

- symmetrized correlator:

$$G_S^{xyxy}(\omega, \mathbf{0}) = \int d^3x dt e^{i(\omega t - \mathbf{k} \cdot \mathbf{x})} \left\langle \frac{1}{2} \{ T^{xy}(t, \mathbf{x}), T^{xy}(0, \mathbf{0}) \} \right\rangle$$

- for the shear-shear contribution \Rightarrow

$$G_{R, \text{shear-shear}}^{xyxy}(\omega, \mathbf{0}) = -\frac{7T}{90\pi^2} \Lambda^3 - i\omega \frac{7T}{60\pi^2} \frac{\Lambda}{\gamma_\eta} + (i+1)\omega^{3/2} \frac{7T}{90\pi^2} \frac{1}{\gamma_\eta^{3/2}}$$

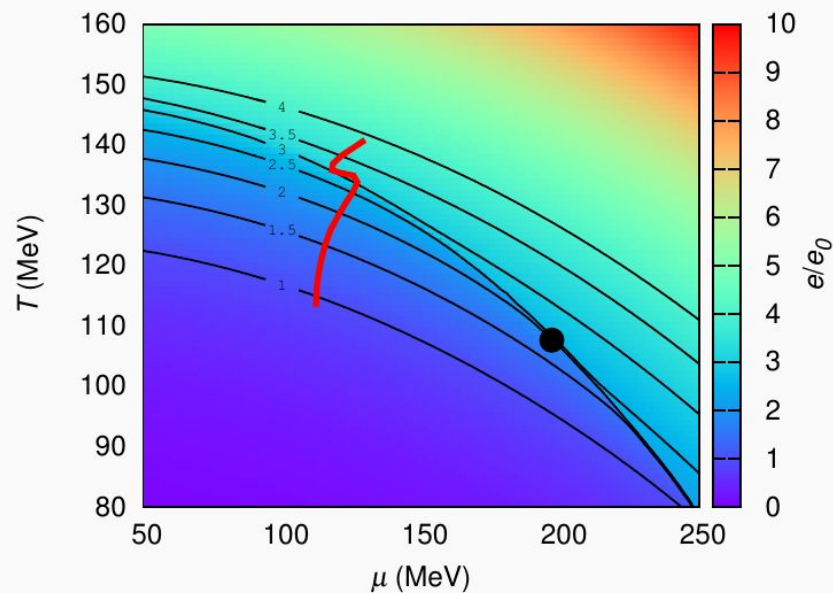
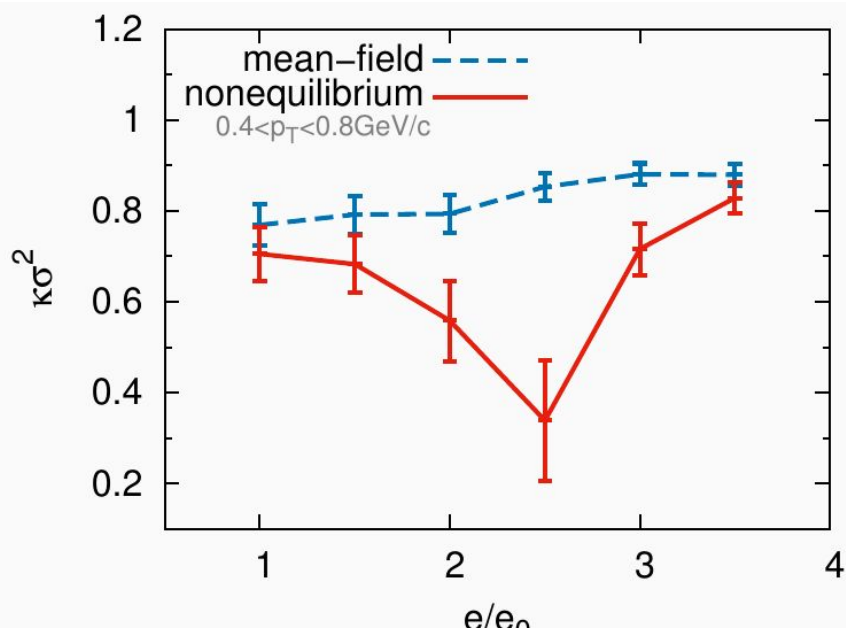
cutoff-dependent
fluctuation contribution
to the pressure

cutoff-dependent
correction to η

frequency-dependent
contribution to
 η and τ_π

Net-proton fluctuations near the critical point

- UrQMD initial conditions rescaled to the EoS of the effective model.
- From densities to particles via Cooper-Frye particlization.
- At particlization: densities of the sigma field coupled to the FD densities.



C. Herold, MN, Y. Yan and C. Kobdaj, PRC 93 (2016) no.2

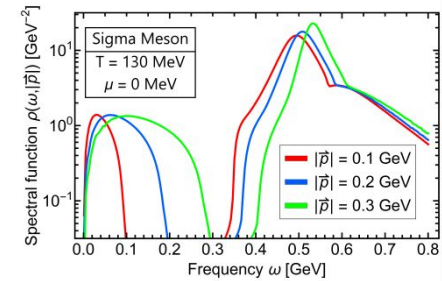
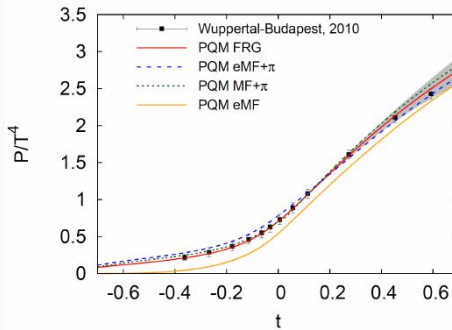
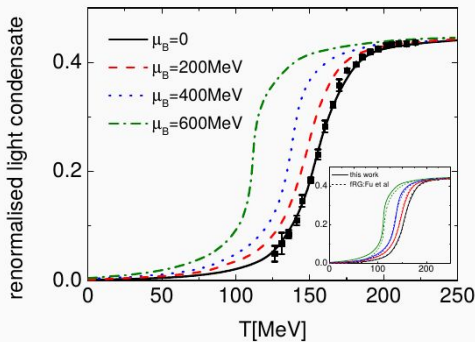
- No non-monotonic behavior in pure mean-field equilibrium calculations.
- Clear signal for criticality in net-proton fluctuations at transition energy density!
- Overall decreasing trend probably due to net-baryon number conservation

NchiFD + FRG \longrightarrow QCD assisted transport

- Include effective potential beyond mean field, momentum dependent equilibrium sigma spectral function \Rightarrow linear response regime of QCD.

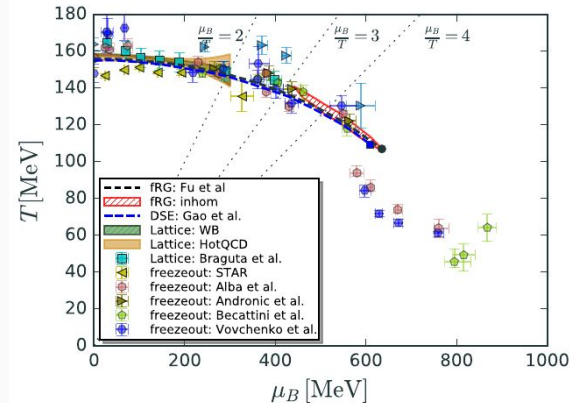
First-principle approach to QCD from the Functional Renormalization Group (FRG)

Cyrol, Mitter, Pawłowski, Strodthoff PRD97 (2018)



F. Gao, J. Pawłowski, 2010.13705; T. Herbst et al, PLB731 (2014); T. Herbst PRD88 (2013); F. Rennecke, J. Pawłowski, N. Wink

- Excellent description of phase structure at vanishing chemical potential.
- Phase structure qualitatively similar to the conjectured QCD phase diagram.
- Obtain spectral functions from analytical continuation.



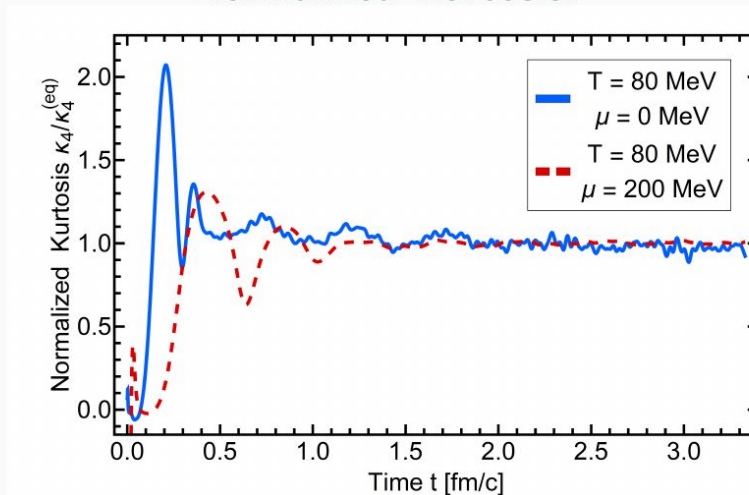
9 / 2

NchiFD + FRG \longrightarrow QCD assisted transport

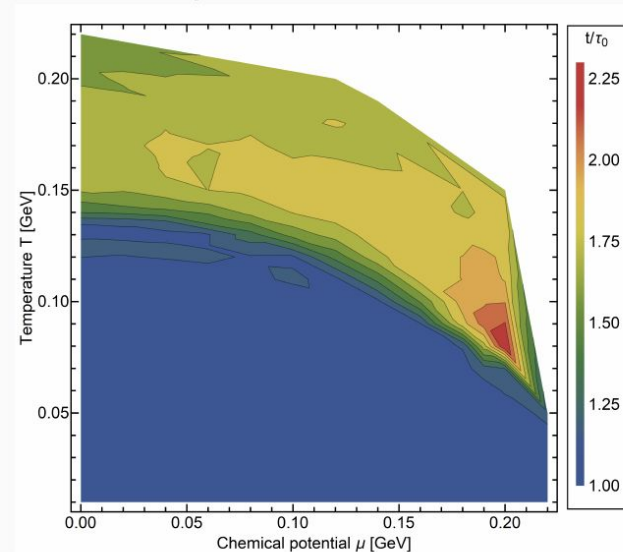
M. Bluhm et al., NPA982 (2019)

Transport equation: $\frac{\delta\Gamma}{\delta\sigma} = \xi$, where $\{\Re\Gamma_\sigma^{(2)}(\omega, \vec{p}), \Im\Gamma_\sigma^{(2)}(\omega, \vec{p}), U\} \in \Gamma$

Normalized Kurtosis:



Equilibration time:



- Critical end point and the phase structure are clearly identifiable.
- Critical slowing down in the vicinity of the critical point, **but no dramatic enhancement of τ_{relax} in a dynamic setup!**

Diffusive dynamics of net-baryon density

- In the long-time, equilibrium limit the net-baryon density is the slowest mode near the CP.
- For baryonic matter that decouples from the energy flow of the system (model B of Hohenberg, Halperin), the diffusive dynamics follows the minimization of the free energy F

$$\partial_t n_B(t, \mathbf{x}) = \kappa \nabla^2 \left(\frac{\delta \mathcal{F}[n_B]}{\delta n_B} \right) + \nabla \mathbf{J}(t, \mathbf{x})$$

with the stochastic current
(Gaussian, white noise)

$$\mathbf{J}(t, \mathbf{x}) = \sqrt{2T\kappa} \zeta(t, \mathbf{x}), \quad \kappa = \frac{Dn_c}{T}$$

MN and M. Bluhm, PRD 99 (2019) and PRD 102 (2020)

First: static box with fixed temperature

Critical energy density from 3d Ising Model

Ginzburg-Landau

$$\mathcal{F}[n_B] = T \int d^3 \left(\frac{m^2}{2n_c^2} \Delta n_B^2 + \frac{K}{2n_c^2} (\nabla \Delta n_B)^2 + \frac{\lambda_3}{3n_c^3} \Delta n_B^3 + \frac{\lambda_4}{4n_c^4} \Delta n_B^4 + \frac{\lambda_6}{6n_c^6} \Delta n_B^6 \right)$$

The couplings depend on temperature via the correlation length $\xi(T)$:

$$m^2 = 1/(\xi_0 \xi^2) \quad \text{Gauss + surface}$$

$$K = \tilde{K}/\xi_0$$

$$\lambda_3 = n_c \tilde{\lambda}_3 (\xi/\xi_0)^{-3/2}$$

$$\lambda_4 = n_c \tilde{\lambda}_4 (\xi/\xi_0)^{-1}$$

$$\lambda_6 = n_c \tilde{\lambda}_6$$

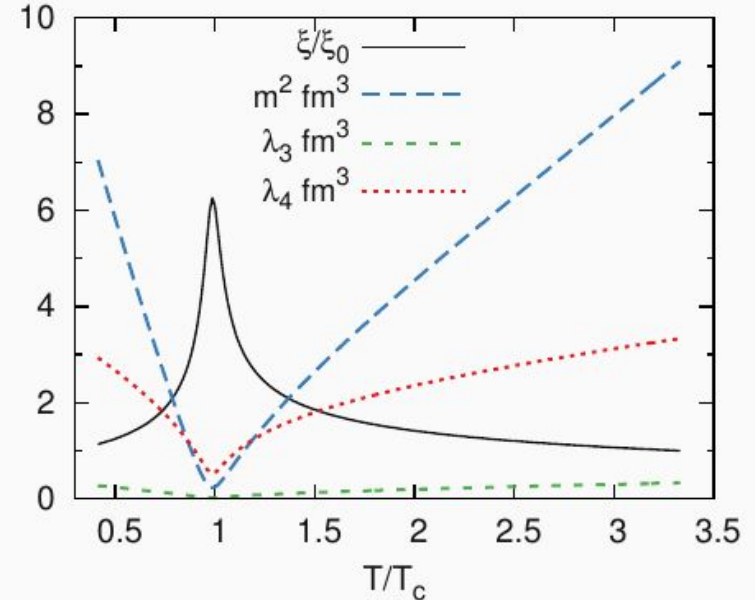
M. Tsypin PRL 73 (1994); PRB 55 (1997)

parameter choice: $\Delta n_B = n_B - n_c$

$\xi_0 \sim 0.5 \text{ fm}$, $T_c = 0.15 \text{ GeV}$, $n_c = 1/3 \text{ fm}^{-3}$

$K = 1/\xi_0$ (surface tension)

$\tilde{\lambda}_3, \tilde{\lambda}_4, \tilde{\lambda}_6$ (universal, but mapping to QCD)



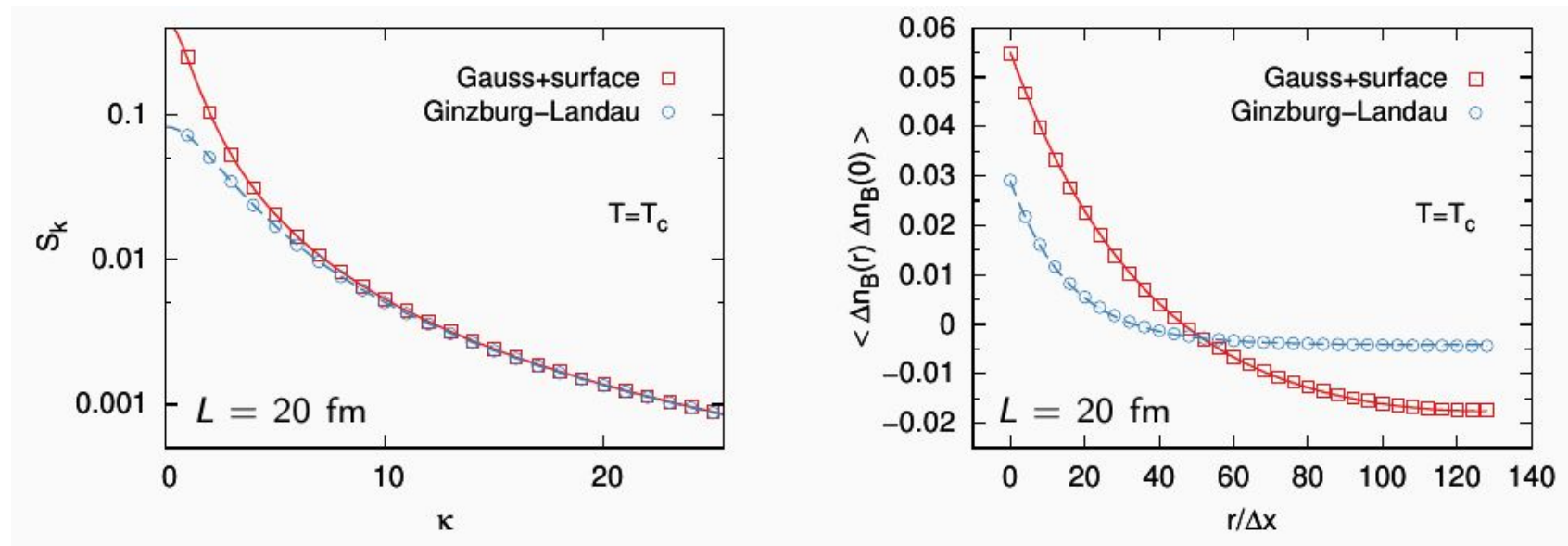
in this Fig.: $\tilde{\lambda}_3 = 1$, $\tilde{\lambda}_4 = 10$

Studied in a static and cooling box of QGP

Validated in the equilibrium limit for the Gauss + surface model:

- Structure factor and equal-time correlation function are well reproduced
- Approaches continuum as resolution is increased
- Baryon conservation effects under control

Important step for all fluctuating codes!



Scaling of equilibrium cumulants

- Expected scaling in an infinite system

($\xi \ll V$): M. Stephanov PRL 102 (2009)

$$\sigma_V^2 \propto \xi^2, \quad (S\sigma)_V \propto \xi^{2.5}, \quad (\kappa\sigma^2)_V \propto \xi^5$$

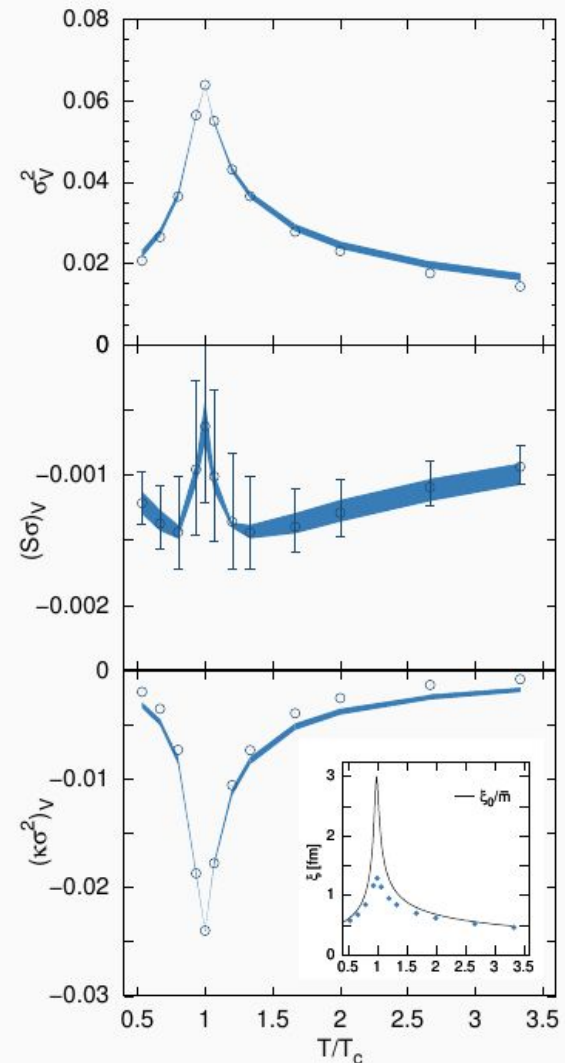
- Here, a finite system with **exact baryon conservation** ($\xi \lesssim V$)! Can be systematically studied in $\xi/V \Rightarrow$ affects equilibrium scaling!

- E.g. for the skewness terms $\propto \lambda_3\lambda_4$ and $\propto \lambda_3\lambda_6$ contribute with opposite sign.

$$\sigma_V^2 \propto \xi^{1.3 \pm 0.05}$$

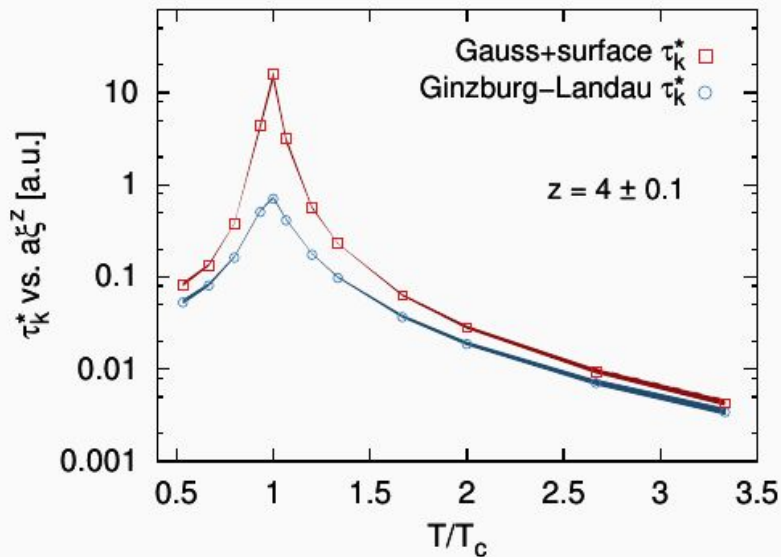
$$(S\sigma)_V \propto -\#\xi^{1.47 \pm 0.05} + \#\xi^{2.4 \pm 0.05}$$

$$(\kappa\sigma^2)_V \propto \xi^{2.5 \pm 0.1}$$



Dynamical critical scaling

- Dynamical structure factor for Gaussian model in continuum:
 $S(k, t) = S(k) \exp(-t/\tau_k)$ with $\tau_k^{-1} = \frac{Dm^2}{n_c} \left(1 + \frac{K}{m^2} k^2\right) k^2$
- Analyze ξ -dependence of relaxation time for modes with $k^* = 1/\xi$:



for both models: $\tau_k^* = a \xi^z$ with

$$z = 4 \pm 0.1$$

$$a = \frac{n_c \xi_0}{D(1 + \tilde{K})}$$

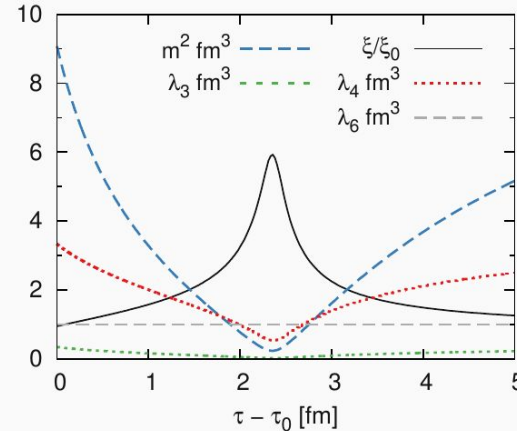
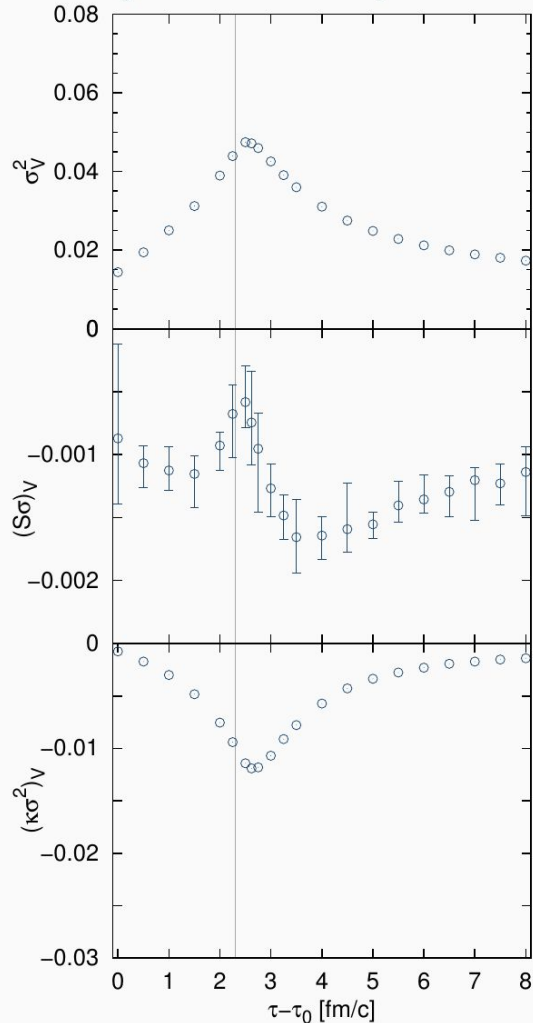
\Rightarrow Simulations reproduce scaling of model B!

For the full dynamics of a HIC, couple to fluctuations in $T^{\mu\nu}$ \longrightarrow model H

According to Hohenberg, Halperin,
 Rev. Mod. Phys. 49 (1977)

Time evolution of critical fluctuations

For a Bjorken-like temperature evolution:



- shift of extrema for variance/kurtosis (retardation effects) to later times corresponding to $T(\tau) < T_c$
- |extremal values| in dyn simulations < equilibrium values (nonequilibrium effects):

$$(\sigma_V^2)_{\text{dyn}}^{\text{max}} \approx 0.75 (\sigma_V^2)_{\text{eq}}^{\text{max}}$$

$$((\kappa\sigma^2)_V)_{\text{dyn}}^{\text{min}} \approx 0.5 (\kappa\sigma_V^2)_{\text{eq}}^{\text{min}}$$

- expected behavior with varying D and c_s^2 (expansion rate)

The role of numerics and computational resources

Time for two heavy nuclei to collide and produce particles:

$\sim 10^{-23}$ seconds

Time for a simulation of two colliding heavy nuclei and particle production:

~ 1 hour

Example: even with a Gaussian Process Emulator the Bayesian analysis of the temperature dependence of bulk and shear viscosity costs 100 Mio CPU hours.

This assumes $O(10^4)$ events per point, fluctuation studies easily require $O(10^8)$ events per point...

