

# Probing dark photon superradiance through follow-up searches

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Princeton Gravity Initiative

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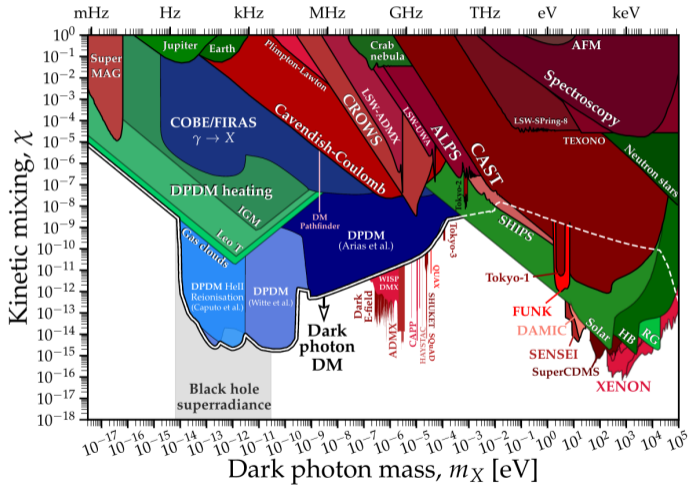
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[Caputo et al., 2021]



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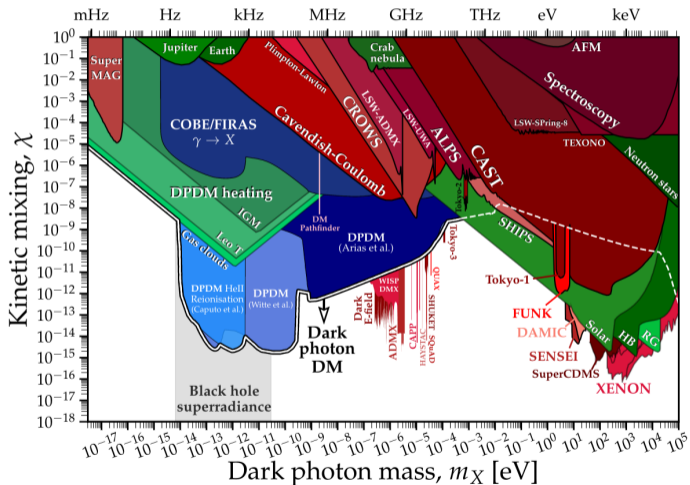
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Unique features:

- Only gravitational interaction
- No dark matter assumption

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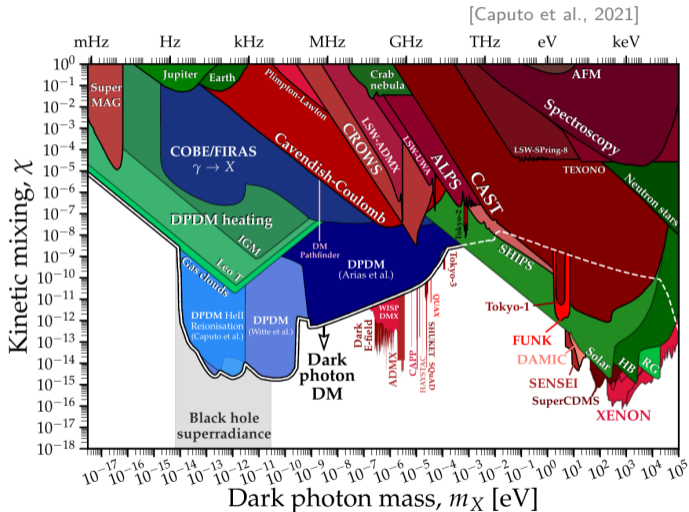
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Requirements:

1. Isolate search strategy
2. Understand theory



# Waveform modelling

Large body of work

[Arvanitaki, Baryakhtar, Brito, East, Yoshino, Kodama, ...]

- Growth rates, power, frequency evolution
- Across parameter space, various modes
- Scalar & vector fields
- (Semi-)analytic, numerical techniques

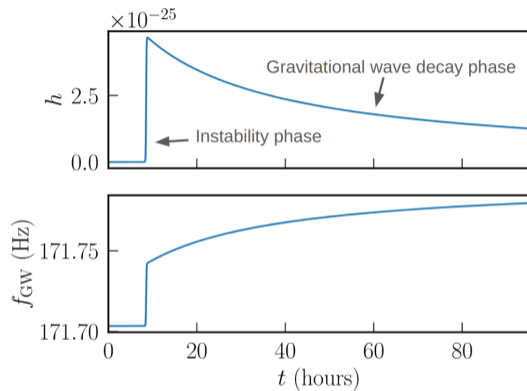


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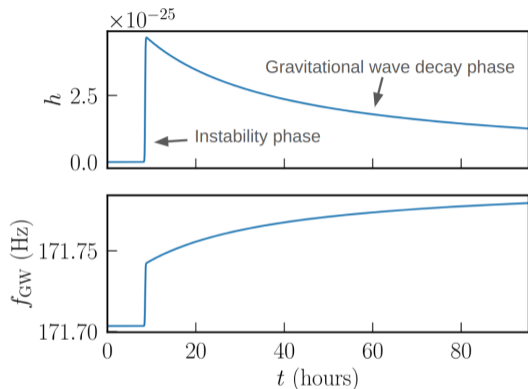
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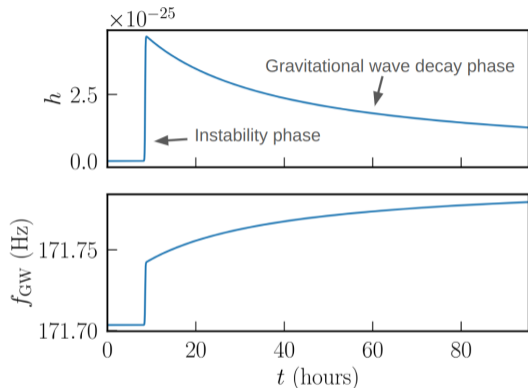
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## SuperRad v2 [May et al. (incl. NS), 2024]

- Accurately model the frequency evolution & understand theoretical uncertainties
- Current waveform phase error  $\Delta\phi \lesssim 1$  across parameter space



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Well yes, but no.

⇒ There is a need to understand non-gravitational interactions

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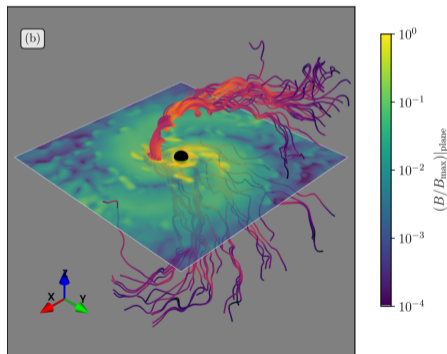
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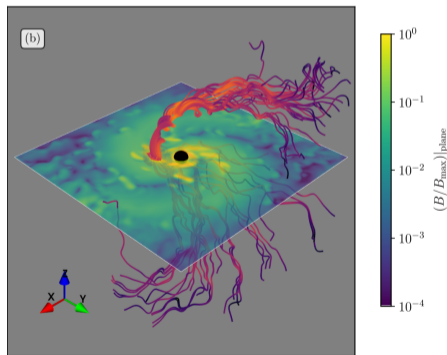
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  - Luminosity:  $L \lesssim 10^{43}$  erg/s
- ⇒ Superradiance mechanism largely unaffected
- Evidence for periodicity ⇒ “fake pulsars”
  - $B \lesssim 10^8$  Gauss ⇒ X-ray &  $\gamma$ -ray
- ⇒ High-energy electromagnetic signatures



# Leverage follow-up searches

How is the gravitational wave signal modified?

[NS, Modino, Egaña-Ugrinovic, Huang, Baryakhtar, East, 2023]

- GW amplitude may decay more rapidly
  - GW frequency may evolve more quickly
- ⇒ Understand impact on GWs as function of  $\varepsilon$



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[Jones et al. (incl. NS), in prep]

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  - Population of merger remnants
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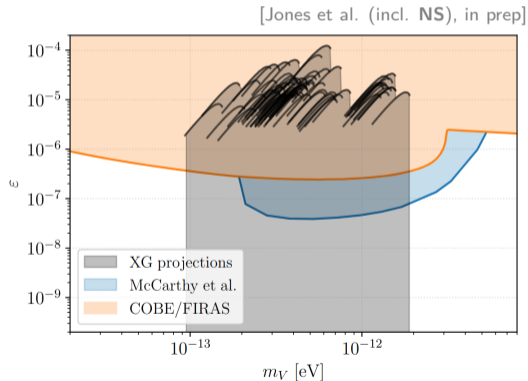
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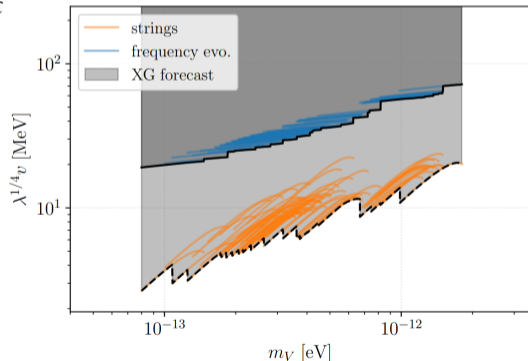
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Higgs-Abelian model:

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# Summary

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- Waveform modelling basically complete
- Understanding non-gravitational interactions is important

## Outlook dark photon superradiance:

- Impact on accretion disks
- Microscopic understanding of plasma dynamics

## Outlook:

- Waveform modelling: higher-order modes? get frequency evolution to matched-filter level?
- Weakly nonlinear effects in Higgs-Abelian sector
- What about LISA?