

Superfluid Fraction of the Inner Crust of Neutron Stars

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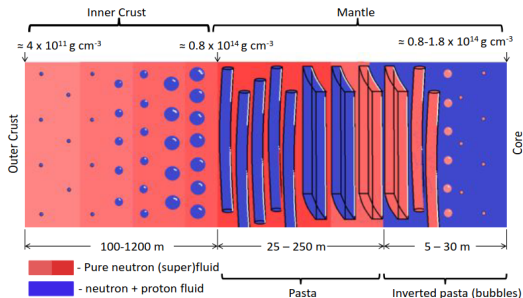


Outline

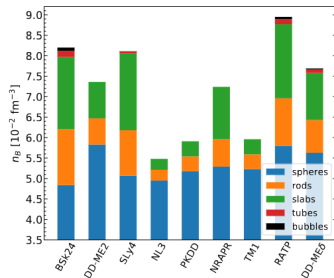
- ▶ Motivation
- ▶ Hartree-Fock-Bogoliubov in the slab and rod phases
- ▶ Conclusions

Inner crust of neutron stars: crystal and pasta phases

- ▶ Coulomb lattice of **Clusters** surrounded by a **superfluid neutron gas**
- ▶ in the deep layers of the inner crust, the clusters can take rod or slab shape (“pasta” phases)



[W.G. Newton]



[H. Dinh Thi et al.,
A&A 654, A114 (2021)]

Entrainment: band theory vs. hydrodynamics

Relative flow of neutrons vs. clusters:
Some neutrons are entrained by the clusters.
How many neutrons are superfluid?

- ▶ Normal band theory

[N. Chamel & P. Haensel,
Liv. Rev. Relativity 11 (2008)]

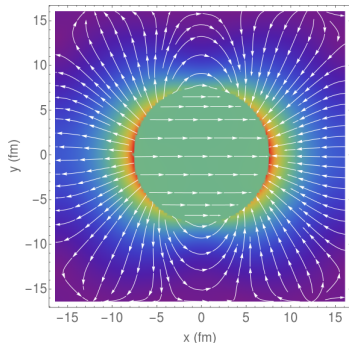
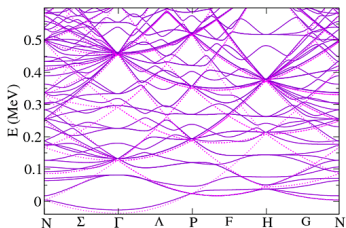
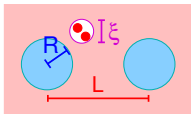
valid for weak coupling ($\Delta \rightarrow 0$)

- ▶ Superfluid hydrodynamics

[N. Martin & MU 2016, PRC 94 (2016)]

valid for strong coupling

$$\xi \propto \frac{k_F}{\pi m \Delta} \ll L$$



Superfluid fraction and Vela glitches

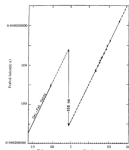
- ▶ **Superfluid hydrodynamics** predicts

much weaker entrainment
 (= larger superfluid fraction)
 than normal band theory

$\xi \sim R \rightarrow$ parameter $\delta < 1$ to account for
 reduction of superfluidity inside clusters
 (nuclear moments of inertia suggest $\delta \sim 0.5$)

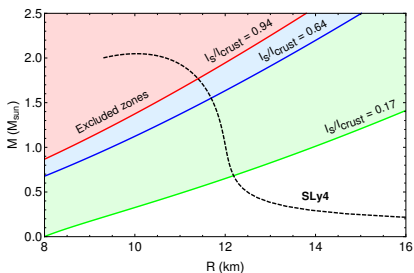
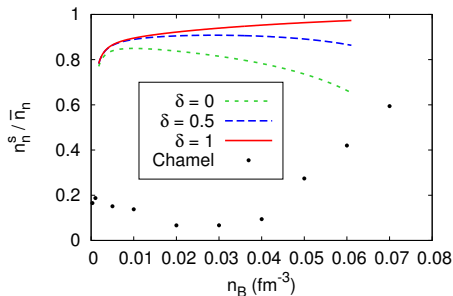
- ▶ Pulsar glitches

[Radhakrishnan
 & Manchester
 Nature 222 (1969)]



- ▶ Superfluid part of the crust moment of inertia in band theory ~ 0.17
 \rightarrow too small to explain Vela glitches

- ▶ In hydrodynamics, even with $\delta = 0$,
 there is enough superfluid density in
 the crust



Superfluid fraction and phonon velocities

larger superfluid fraction
(= weaker entrainment) implies:

- ▶ decreased effective cluster mass
- ▶ higher speeds v_i of phonons

coupling between lattice and superfluid phonons

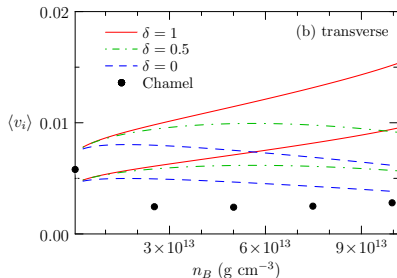
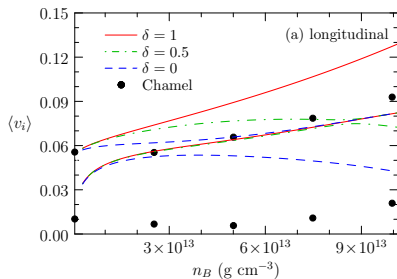
[C. Pethick, S. Reddy & N. Chamel, Prog. Theor. Phys. Suppl. 186 (2010);
D. Durel & MU, PRC 97, 065805 (2018)]

- ▶ suppressed lattice specific heat:

$$c_v = \sum_{i=1}^4 \frac{2\pi^2 T^3}{15 \langle v_i \rangle^3}$$

angle average:

$$\langle v_i \rangle = \left(\int \frac{d\Omega}{4\pi} \frac{1}{v_i^3} \right)^{-1/3}$$



[D. Durel & MU, PRC 97, 065805 (2018)]

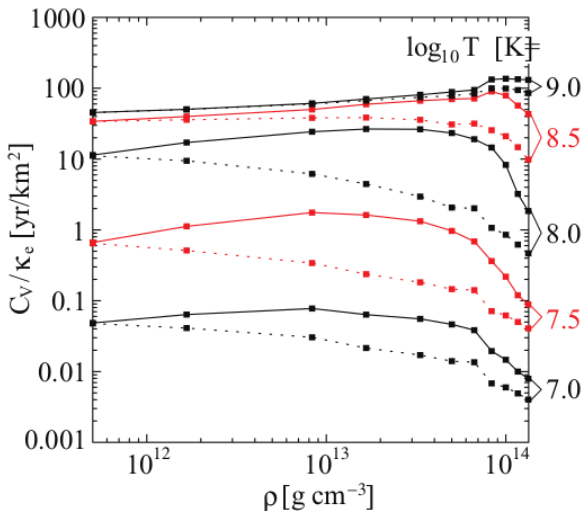
Relevance for cooling

Effect of entrainment on specific heat C_V and heat conductivity κ_e :

Figure taken from Chamel, Page & Reddy, PRC 87, 035803 (2013)

- ▶ solid lines: with strong entrainment (from normal band theory)
- ▶ dotted lines: without entrainment

(curves include phonon and electron specific heat, neutron specific heat suppressed by pairing)



Hartree-Fock-Bogoliubov (HFB) with periodicity

HFB can interpolate between normal band theory in weak coupling and superfluid hydrodynamics in strong coupling

$$\begin{pmatrix} \hbar - \mu & -\Delta \\ -\Delta^\dagger & -\hbar + \mu \end{pmatrix} \begin{pmatrix} U_\alpha^* \\ -V_\alpha \end{pmatrix} = E_\alpha \begin{pmatrix} U_\alpha^* \\ -V_\alpha \end{pmatrix}$$

work in momentum space \rightarrow matrices in discrete (band) indices, diagonal in the (continuous) Bloch and parallel (for rods, slabs) momenta

$$h_{kk'} = \left(\frac{1}{2m} \right)_{kk'} k \cdot k' + U_{kk'} - \hbar k \cdot \mathbf{v} \delta_{kk'}$$

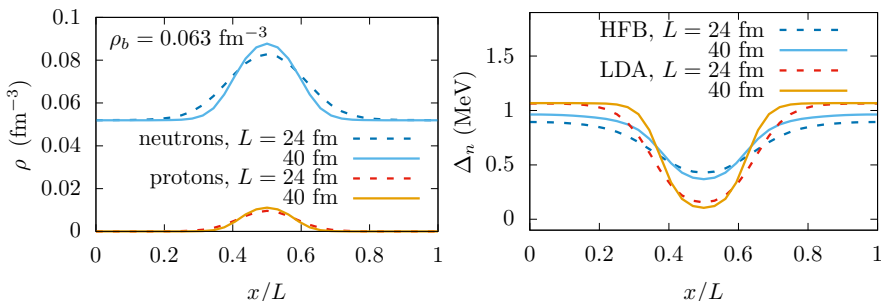
\mathbf{v} = velocity of the slabs or rods in the rest frame of the superfluid

mean field: $U_{kk'} = - \sum_{pp'} V_{kpk'p'} \rho_{p'p}$ (Skyrme functional)

gap: $\Delta_{kk'} = - \sum_{pp'} V_{kk'p'p} \kappa_{p'p}$ (separable interaction $\sim V_{\text{low}-k}$)

“BCS approximation” (only diagonal elements of Δ are retained) not sufficient
[Minami & Watanabe, Phys. Rev. Res. 4 (2022)]

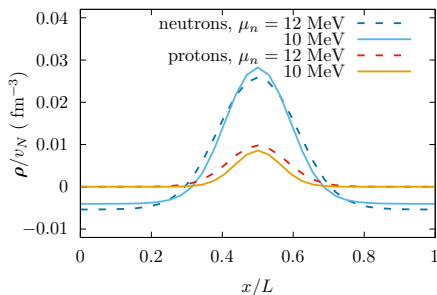
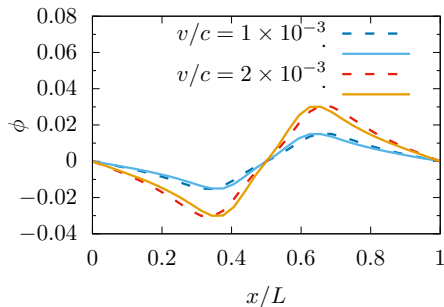
Slab phase (lasagna): density profile and gap



Almirante & Urban, Phys. Rev. C 109, 045805 (2024)

- ▶ gap inside the slab is smaller than in the neutron gas
- ▶ local-density approximation (LDA) overestimates this suppression

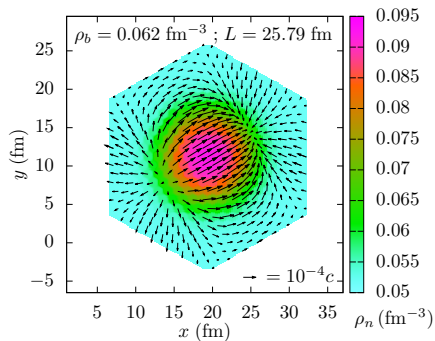
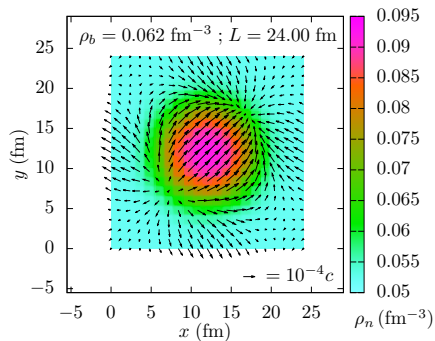
Slab phase (lasagna): phase of the gap and current



Almirante & Urban, Phys. Rev. C 109, 045805 (2024)

- ▶ phase $\phi \propto v$ → linear regime
- ▶ proton current = $v \times$ proton density
- ▶ neutron current shifted down by a constant → superfluid part doesn't move

Rod phase (spaghetti): density and current in square and hexagonal lattices



Almirante & Urban, Phys. Rev. C 110, 065802 (2024)

Results for superfluid fraction

Spaghetti

μ_n (MeV)	L (fm)	ρ_b (fm ⁻³)	$\rho_S/\bar{\rho}_n$ (HFB %)	$\rho_S/\bar{\rho}_n$ (Carter ^a %)
12	24	0.0619	94.5	75
	28	0.0617	95.7	
12.5	24	0.0670	95.4	82
	28	0.0668	96.7	

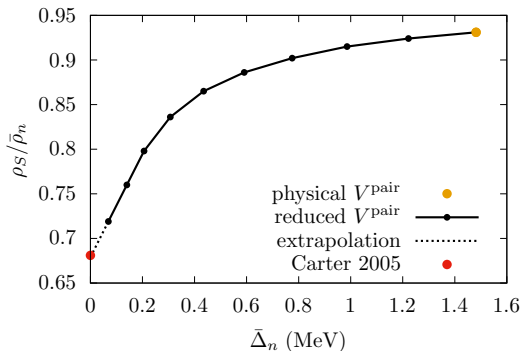
Lasagna

μ_n (MeV)	L (fm)	ρ_b (fm ⁻³)	$\rho_S/\bar{\rho}_n$ (HFB %)	$\rho_S/\bar{\rho}_n$ (Carter ^a %)
13	20	0.0723	96.3	93
	24	0.0720	96.2	
13.5	20	0.0768	97.2	94
	24	0.0766	97.1	

^a normal band theory: B. Carter, N. Chamel & P. Haensel, NPA 748 (2005).

Band structure effects vs. pairing gap

- ▶ Normal band theory should be valid in the weak-coupling limit ($\Delta \rightarrow 0$)
- ▶ Superfluid hydrodynamics only valid for $\xi \ll L \rightarrow \Delta \gg \frac{k_F}{\pi mL}$
- ▶ HFB should be valid all the way between these two limits!
- ▶ Varying artificially the strength of the pairing interaction:



Rod phase,
 $\bar{\rho}_n = 0.059 \text{ fm}^{-3}$
 $L = 27.17 \text{ fm}$,
for same conditions as in
Carter, Chamel & Haensel,
NPA 748 (2005)

Conclusions

- ▶ Superfluid fraction important for glitches and cooling
- ▶ Strong discrepancy between normal band theory (valid for weak coupling) and superfluid hydrodynamics (valid for strong coupling)
- ▶ HFB theory interpolates between these two limits
- ▶ HFB for slab and rod phases gives larger superfluid fraction (closer to hydrodynamic result) than normal band theory
- ▶ HFB for crystalline phase: work in progress